


Determination of the fission-cycling onset in the r-process*

Mei-Yue-Nan Ma (马梅月楠)¹  Zhi-Hong Li (李志宏)^{1,2,3†} Ge-Xing Li (李歌星)¹ Chao Dong (董超)¹
Na Song (宋娜)¹ Chen Chen (陈晨)¹ Jun-Wen Tian (田竣文)¹ Jia-Ying-Hao Li (李家英豪)¹
Zhi-Cheng Zhang (张智程)¹ Hui-Ling Tian (田慧玲)¹

¹China Institute of Atomic Energy, Beijing 102413, China

²University of Chinese Academy of Science, Beijing 101408, China

³Jinping Deep Underground Frontier Science and Dark Matter Key Laboratory of Sichuan Province, Liangshan 615000, China

Abstract: Identifying the thermodynamic conditions marking the onset of fission cycling is crucial for modeling heavy-element production in the r-process. In this work, we develop a framework to determine this onset over the (T_9, n_n) plane. We define a heavy-region condition band near $N \approx 184$ and construct an equilibrium path band based on the effective neutron separation energy $S_n^0(T_9, n_n)$. We then compare the lifetimes of neutron-induced fission and β -decay for nuclei with $94 \leq Z \leq 106$. Within this framework, we construct a continuous map of actinide nuclei along the equilibrium path band, identifying where neutron-induced fission first overtakes β -decay. We find that, with increasing temperature and neutron density, the onset shifts toward nuclei with lower proton number (Z) and smaller mass number (A), transitioning from the Es-Cf region to the Am region. These results provide a quantitative benchmark for identifying the conditions under which fission cycling occurs in heavy r-process environments.

Keywords: r-process, fission cycling, $N=184$ shell, equilibrium path, neutron-induced fission

DOI: 10.1088/1674-1137/ae5ef5 **CSTR:** 32044.14.ChinesePhysicsC.50074104

I. INTRODUCTION

The rapid neutron-capture process (r-process) is responsible for roughly half of the elements heavier than iron [1–4]. In neutron-rich environments, the reaction flow can proceed to very heavy nuclei, where neutron-induced fission competes with neutron capture and β -decay [5–7]. Once fission dominates, fission recycling regulates the production of actinides and shapes the late-time abundance patterns of heavy elements [8–11]. Consequently, identifying the onset of fission cycling is a central problem in r-process nucleosynthesis.

For a given astrophysical environment (T_9, n_n) , the onset of fission cycling depends on whether the r-process path reaches nuclei for which neutron-induced fission is faster than β -decay and on the first nucleus along the path that satisfies this condition [12]. These nuclei are expected to be located primarily in the actinide and light transuranic region near the predicted shell closure at $N \approx 184$ [10, 13]. In this region, nuclear masses, fission barriers, and decay properties remain highly uncertain, with significant discrepancies among global nuclear models [7, 13, 14]. Consequently, both the predicted location and the robustness of fission cycling vary considerably

across existing studies, highlighting their strong dependence on nuclear physics inputs [8, 10, 12, 15].

Most previous studies of fission cycling have relied on reaction-network nucleosynthesis, either through post-processing of hydrodynamic ejecta trajectories (*e.g.*, [16–18]) or via parameterized surveys across a fixed set of outflow conditions (*e.g.*, [19]). Complementary studies have focused on deriving neutron-induced fission rates and quantifying their model dependence (*e.g.*, [12]). Although these network-based approaches have yielded valuable insights, they offer limited guidance for establishing a systematic correlation between the onset of fission cycling and the underlying nuclear structure inputs across the full (T_9, n_n) plane. In particular, the correlation between the equilibrium r-process path and the onset of fission cycling, as a function of temperature and neutron density, has not been explored in a continuous and reproducible manner [20].

In this work, we develop a systematic framework to identify the onset of fission cycling across the (T_9, n_n) plane. Building on the structural trends of neutron-rich isotopic chains near the predicted $N \approx 184$ shell closure, we first determine the thermodynamic conditions under which the equilibrium r-process path enters the very

Received 4 February 2026; Accepted 14 April 2026; Accepted manuscript online 15 April 2026

* This work was supported by the National Natural Science Foundation of China (12475151, 12405164), the Continuous-Support Basic Scientific Research Project (BJ010261223284)

† E-mail: zhli@ciae.ac.cn

©2026 Chinese Physical Society and the Institute of High Energy Physics of the Chinese Academy of Sciences and the Institute of Modern Physics of the Chinese Academy of Sciences and IOP Publishing Ltd. All rights, including for text and data mining, AI training, and similar technologies, are reserved.

heavy mass region, thereby forming a continuous band in the (T_9, n_n) plane [21–23].

We then construct an equilibrium path band based on the effective neutron separation energy S_n^0 as a structural criterion. This band includes nuclei whose separation-energy systematics are consistent with the equilibrium r-process path in this region. In particular, we focus on nuclei with $94 \leq Z \leq 106$, corresponding to actinide and early transuranic nuclei that are most relevant for fission cycling.

Finally, we perform point-by-point evaluations of neutron-induced fission and β -decay lifetimes using a consistent set of nuclear inputs. For each (T_9, n_n) condition, the first nucleus satisfying $\tau_{n,f} < \tau_\beta$ is identified as marking the onset of fission cycling.

By combining the heavy-region condition band with these lifetime evaluations, the framework yields a global and transparent map of fission-cycling onset nuclei across the (T_9, n_n) plane. This provides a complementary perspective to traditional trajectory-based r-process studies.

II. NUCLEAR INPUTS AND OVERALL STRATEGY

The framework developed in Sec. III relies on a consistent set of nuclear inputs that characterize the structural properties of neutron-rich nuclei and govern the competition between β -decay and neutron-induced fission.

A. Nuclear masses and separation energies

All structural quantities used in constructing the equilibrium path are derived from the FRDM2012 global mass model [24]. The one- and two-neutron separation energies are defined as

$$\begin{aligned} S_n(Z, A) &= B(Z, A) - B(Z, A - 1), \\ S_{2n}(Z, A) &= B(Z, A) - B(Z, A - 2). \end{aligned} \quad (1)$$

These considerations form the basis for evaluating equilibrium abundance ratios and for identifying structural trends among adjacent isotopic chains, including evolution toward the predicted neutron shell closure at $N \approx 184$. The behavior of S_{2n} across the actinide region determines the selection of a consistent Z window for constructing the heavy-region condition band in the (T_9, n_n) plane.

B. Neutron-induced fission reaction rates and lifetimes

Neutron-induced fission lifetimes are derived from stellar reaction rates computed with the TALYS 2.0 nuclear reaction code in astrophysical-rate mode [25]. Throughout this work, a uniform set of nuclear-physics inputs is employed to ensure consistency across the nuclear region under study.

Nuclear masses for the reaction-rate calculations are taken from the Möller finite-range droplet model (FRDM2012), which provides the structural baseline for this study. Neutron-induced fission is treated as a competing decay channel of the compound nucleus. For the baseline onset mapping, fission barriers follow the standard TALYS prescription: experimentally constrained barriers are adopted where available; otherwise, a macroscopic rotating liquid-drop model is used.

These stellar neutron-induced fission rates are then converted to the corresponding fission lifetimes. The sensitivity of the inferred onset nuclei to the adopted nuclear mass model and to alternative fission-barrier prescriptions is examined separately in Sec. V.A and V.B.

C. β -decay lifetimes

β -decay lifetimes are taken from the large-scale QRPA calculations of Marketin *et al.* [26], which provide consistent coverage of neutron-rich heavy nuclei. These lifetimes are used to compare β -decay and neutron-induced fission lifetimes along the actinide segments of the equilibrium path band, and to identify, for each (T_9, n_n) point, the nucleus satisfying $\tau_{n,f} < \tau_\beta$.

D. Overall strategy

Taken together, the FRDM2012 nuclear masses, TALYS stellar reaction rates, and QRPA β -decay lifetimes form the set of nuclear inputs used to construct the heavy-region condition band, define the equilibrium path band, and determine the onset of fission cycling across the (T_9, n_n) plane.

III. METHODS

This section outlines the computational framework used to determine the conditions under which the equilibrium r-process path can reach the very heavy neutron-rich region. The predicted shell closure at $N \approx 184$ is adopted as a structural reference.

The framework also identifies, for each (T_9, n_n) condition, the first nucleus for which neutron-induced fission outpaces β -decay.

Our methodology builds on and extends the waiting-point (WP) formulation of Xu *et al.* [27], incorporating additional structural diagnostics, equilibrium-path filtering, and large-scale reaction-lifetime calculations using TALYS 2.0.

A. Equilibrium abundance ratio and the effective separation energy $S_n^0(T_9, n_n)$

In $(n, \gamma) \rightleftharpoons (\gamma, n)$ equilibrium, the abundance ratio between adjacent isotopes within an isotopic chain follows the standard Saha relation [28]. For completeness, we briefly summarize the key elements of the equilibri-

um formulation introduced in Ref. [27] that are needed for the present work. It is convenient to group the slowly varying prefactors—including the neutron density, thermal phase-space factor, partition function, and mass factors—into a single effective quantity, thereby defining the equilibrium one-neutron separation energy $S_n^0(T_9, n_n)$.

With this definition, the abundance ratio can be written in a compact form

$$\frac{Y(Z, A+1)}{Y(Z, A)} \simeq \exp\left[\frac{S_n(Z, A+1) - S_n^0(T_9, n_n)}{kT}\right], \quad (2)$$

where Y denotes the nuclear abundance and S_n the one-neutron separation energy.

The equilibrium r-process path is defined by the condition $Y(Z, A+1)/Y(Z, A) = 1$, which is equivalent to the separation-energy condition

$$S_n(Z, A) = S_n^0(T_9, n_n). \quad (3)$$

In practical applications, the partition-function ratio $G(Z, A+1)/G(Z, A)$ and the mass factor $(A+1)/A$ are typically set to unity, because their deviations from unity are small and therefore have a negligible impact on S_n^0 over the relevant thermodynamic range. Under this approximation, S_n^0 depends solely on (T_9, n_n) and assumes the analytic form derived by Xu *et al.* [27] for $(n, \gamma) \rightleftharpoons (\gamma, n)$ equilibrium in r-process environments:

$$S_n^0(T_9, n_n) \approx T_9 \left[2.79 + \frac{1.5 \log_{10} T_9 - \log_{10} \left(\frac{n_n}{10^{20} \text{cm}^{-3}} \right)}{5.04} \right]. \quad (4)$$

This expression captures the joint dependence of isotopic abundance ratios on temperature and neutron density and is valid over the ranges $1 \leq T_9 \leq 3$ and $10^{20} \leq n_n \leq 10^{30} \text{cm}^{-3}$, which are consistent with the astrophysical conditions considered in this work.

B. Critical waiting-point nuclei and T_9 - n_n conditions

In this work, we adopt and explicitly implement the *critical waiting-point* (CWP) framework introduced by Xu *et al.* [27], which is based on the equilibrium formulation summarized in Sec. III.A. Within this framework, the equilibrium r-process path, under $(n, \gamma) \rightleftharpoons (\gamma, n)$ conditions, proceeds predominantly along even- N isotopic chains and is characterized by a nearly constant effective neutron separation energy $S_n^0(T_9, n_n)$.

CWP nuclei constitute a subset of waiting-point nuclei that primarily constrain the thermodynamic conditions required for the formation of a given r-process abundance peak. Using the definition of $S_n^0(T_9, n_n)$ from

Sec. III.A, the CWP condition can be expressed in terms of separation-energy systematics along an even- N isotopic chain. For two even- N isotopes (Z, N) and $(Z, N+2)$ within the same isotopic chain, the abundance ratio follows from Eq. (2) as:

$$\frac{Y(Z, A)}{Y(Z, A-2)} \simeq \exp\left[\frac{S_{2n}(Z, A) - 2S_n^0(T_9, n_n)}{kT}\right]. \quad (5)$$

This provides the basis for deriving the CWP condition. Specifically, a nucleus (Z, N) satisfies the CWP condition if $S_n^0(T_9, n_n)$ lies between the two adjacent values of $S_{2n}/2$ within the same even- N isotopic chain, *i.e.*,

$$\frac{1}{2}S_{2n}(Z, N+2) \lesssim S_n^0(T_9, n_n) \lesssim \frac{1}{2}S_{2n}(Z, N). \quad (6)$$

This defines the range of S_n^0 for which (Z, N) serves as a CWP nucleus along the equilibrium path.

For completeness, the canonical CWP nuclei associated with the classical shell closures at $N = 50, 82$, and 126 , as reported in Ref. [27], are listed in Table 1.

For each even- N isotopic chain near a shell closure, we combine Eq. (6) with the functional form of $S_n^0(T_9, n_n)$ to map the CWP condition onto a finite region in the (T_9, n_n) plane. The intersection of these regions for selected CWP nuclei then defines the thermodynamic conditions necessary to produce the corresponding r-process abundance peak under equilibrium assumptions.

Applying this construction to the classical shell closures at $N = 50, 82, 126$ with FRDM2012 separation energies reproduces the systematic trends reported by Xu *et al.* [27] (Fig. 1), thereby validating our implementation of the CWP framework. This result provides a methodological basis for extending the analysis to the region near $N \approx 184$, a predicted shell closure relevant to actinide nucleosynthesis. By adopting the separation-energy systematics of these classical CWP nuclei as a structural baseline, we ensure that the present analysis of the onset of fission cycling remains consistent with the original framework while adapting it to the thermodynamic conditions of heavy-element nucleosynthesis.

C. Structural reference near $N \approx 184$

A wide range of global nuclear mass models, including macroscopic-microscopic approaches and self-consistent mean-field calculations, predict an enhanced neut-

Table 1. Critical waiting-point nuclei.

N	CWP nuclei
50	$^{80}\text{Zn}, ^{79}\text{Cu}, ^{78}\text{Ni}$
82	$^{130}\text{Cd}, ^{129}\text{Ag}, ^{128}\text{Pd}, ^{127}\text{Rh}, ^{126}\text{Ru}$
126	$^{195}\text{Tm}, ^{194}\text{Er}, ^{193}\text{Ho}, ^{192}\text{Dy}, ^{191}\text{Tb}$

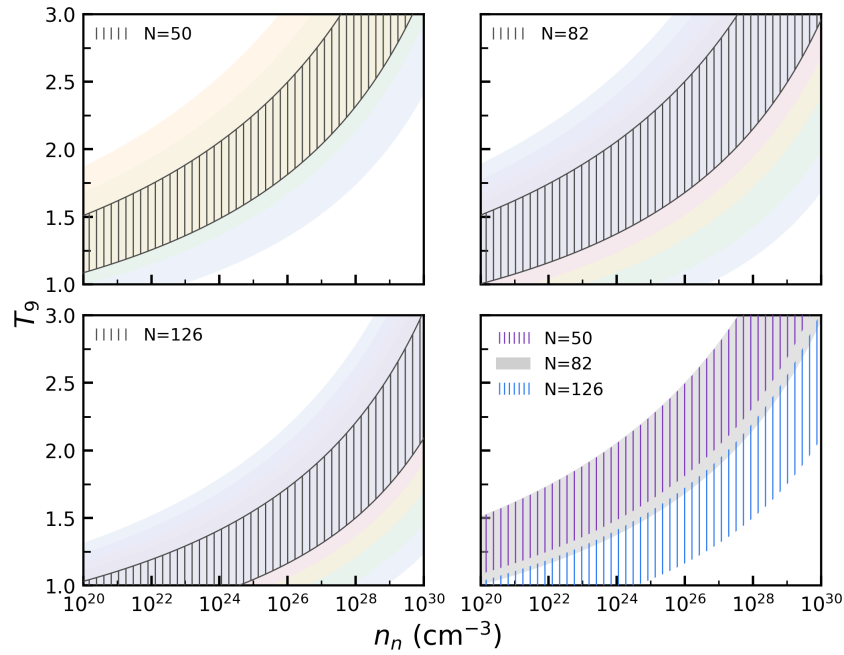


Fig. 1. (color online) The $T_9 - n_n$ conditions required by the $N = 50, 82,$ and 126 CWP nuclei. The lower-right panel compares the relative locations and overlaps of these conditions for the three shell closures. Results are based on FRDM2012 separation energies and reproduce the trends reported by Xu *et al.* [27].

ron shell closure near $N \approx 184$ [13, 29, 30].

Despite model-to-model differences in the predicted shell strength and detailed separation-energy systematics, the region around $N \approx 184$ is widely regarded as a structural benchmark for very neutron-rich heavy nuclei beyond the classical $N = 126$ shell [31, 32].

In this framework, the region near $N \approx 184$ is used as a structural reference to identify when the equilibrium r-process flow enters the very-heavy mass regime, rather than as a strict requirement that the path reach $N = 184$. Accordingly, the extension to the heavy region is formulated in terms of a structural-consistency criterion based on neutron separation energies.

Specifically, the heavy-region condition band is anchored to a common separation-energy window inferred from the classical waiting-point regions at $N = 82$ and $N = 126$, expressed in terms of the $S_{2n}/2$ metric. This construction yields a conceptually model-independent diagnostic, whereas its quantitative realization depends on the separation-energy systematics predicted by the adopted nuclear mass model.

The implementation of this criterion and the resulting structural constraints in the $N \approx 184$ region are presented in the next section.

D. Extension of CWP $T_9 - n_n$ conditions to the heavy region

Building on the validated CWP framework for the classical r-process abundance peaks at $N = 50, 82,$ and 126 (Sec. III.B), we extend the analysis to the very neut-

ron-rich heavy region near $N \approx 184$. As discussed in Sec. III.C, this region is adopted as a structural reference for actinide-producing main r-process conditions.

This extension is restricted to the main r-process regime responsible for the production of heavy nuclei, in particular those in the actinide region, whereas the $N = 50$ peak, commonly attributed to the weak r-process, is excluded from the present consideration [4].

As discussed in Sec. III.B, the classical abundance peaks at $N = 82$ and $N = 126$ are established under similar thermodynamic conditions within the $(n, \gamma) \rightleftharpoons (\gamma, n)$ equilibrium regime. This shared physical basis justifies adopting the corresponding classical S_n^0 window as a structural reference for identifying heavy-region CWP nuclei near $N \approx 184$.

For the heavy-region CWP nuclei identified on this basis, the corresponding S_n^0 intervals are mapped onto the (T_9, n_n) plane through Eq. (4). Their overlap defines the thermodynamic domain within which the r-process can access the heavy region, thereby establishing the corresponding condition band.

To account for residual variations in separation-energy systematics and moderate deviations from ideal equilibrium, we introduce an effective tolerance ΔS , which sets the finite width of this condition band. This band then provides the thermodynamic basis for constructing the equilibrium path band in the following section.

E. Equilibrium path band under CWP conditions

Building on the heavy-region CWP condition band

established in Sec. III.D, the boundary values of the associated separation-energy window provide a natural structural criterion for identifying nuclei accessible to equilibrium r-process flow. We define the equilibrium path band as the set of nuclei in the heavy actinide region whose separation-energy systematics are consistent with this window.

Because the condition band already incorporates the effective tolerance ΔS , which reflects both intrinsic variations in the separation-energy systematics and uncertainty propagation from nuclear masses, no additional uncertainty is introduced into the path-band construction. The band is not restricted to the reference CWP nuclei themselves but also includes neighboring isotopic chains that satisfy the same structural criterion. This extension reflects the role of the $N \approx 184$ region as a structural reference for the onset of access to the heavy region, rather than imposing a restriction to a specific set of nuclei.

A nucleus (Z, N) is deemed to belong to the equilibrium path band if its two-neutron separation energy per neutron, $S_{2n}/2$, lies within this window. This quantity provides a convenient measure of isotopic-chain systematics in very neutron-rich heavy nuclei and remains fully consistent with the CWP-based framework adopted in this work.

The resulting equilibrium path band should therefore be interpreted as a structure-defined envelope of nuclei accessible to equilibrium r-process flow in the heavy actinide region, rather than as a single path corresponding to a fixed thermodynamic trajectory. This set of nuclei provides the nuclear-chart basis for the lifetime comparisons and fission-onset analysis presented in Sec. III.F.

F. Identification of fission-cycling onset along the heavy equilibrium path band

Once the equilibrium r-process flow reaches the actinide region, the competition between neutron-induced fission and β decay governs the removal of material from the heavy equilibrium-path band. Under the CWP T_9 - n_n conditions for the heavy region established in Sec. III. D, this competition is quantified by directly comparing the characteristic lifetimes $\tau_{n,f}(T_9, n_n)$ and τ_β .

Neutron-induced fission lifetimes. These lifetimes are derived from Maxwellian-averaged stellar reaction rates computed using TALYS 2.0. For each nucleus and temperature T_9 , TALYS provides a stellar reaction rate $R_\star(T_9)$, with units of $\text{cm}^3 \text{mol}^{-1} \text{s}^{-1}$, which accounts for the thermal population of excited states. The corresponding rate per target nucleus is

$$\langle \sigma v \rangle_\star(T_9) = \frac{R_\star(T_9)}{N_A}, \quad (7)$$

where N_A denotes Avogadro's number. For a neutron number density of n_n , the fission rate per target nucleus is given by

$$\lambda_{n,f}(T_9, n_n) = n_n \langle \sigma v \rangle_\star(T_9), \quad (8)$$

and the corresponding lifetime is

$$\tau_{n,f}(T_9, n_n) = \lambda_{n,f}^{-1}(T_9, n_n). \quad (9)$$

β -decay lifetimes. These lifetimes are obtained from large-scale covariant RHB+pn-RQRPA calculations by Marketin, Huther, and Martínez-Pinedo, which provide consistent coverage of neutron-rich actinide and transuranic nuclei [26]. For each nucleus, the total decay rate λ_{tot} is used, and the β -decay lifetime is defined as:

$$\tau_\beta = \lambda_{\text{tot}}^{-1}. \quad (10)$$

Application across the path band. The lifetimes described above are evaluated for all nuclei within the structure-defined equilibrium path band (Sec. III. E). At each thermodynamic point (T_9, n_n) within the heavy-region CWP domain, we define the onset of fission cycling in two steps.

First, we identify the minimum proton number Z at which neutron-induced fission becomes competitive with β -decay, *i.e.*, $\tau_{n,f}(T_9, n_n) \leq \tau_\beta$.

Second, for this Z , the onset nucleus is the isotope with the lowest neutron number N that satisfies the same condition.

This procedure provides a consistent and transparent criterion for locating the onset of fission cycling across the (T_9, n_n) plane.

IV. RESULTS

This section presents results from applying the structural and equilibrium-path framework developed in Sec. III to the heavy actinide region.

We first identify the subset of actinide isotopic chains near $N \approx 184$ whose separation-energy systematics are consistent with the common CWP window defined by the classical $N = 82$ and $N = 126$ peaks.

We then map this structurally selected window onto the (T_9, n_n) plane to determine the corresponding thermodynamic domain and to construct the associated equilibrium r-process path band in the actinide region.

Finally, within this restricted set of thermodynamic conditions and nuclear paths, we evaluate the competition between neutron-induced fission and β -decay and identify the onset of fission cycling along the equilibrium

r-process path band.

A. Structural and thermodynamic constraints on the $N \approx 184$ actinide region under CWP T_9-n_n conditions

Applying the construction described in Sec. III.D to the FRDM2012 mass model, we identify a subset of actinide isotopic chains near $N \approx 184$ whose separation-energy systematics are consistent with the common CWP window defined by the classical $N = 82$ and $N = 126$ peaks.

Figure 2 illustrates the systematics of the two-neutron separation energy, $S_{2n}/2$, for isotopic chains near the classical CWP regions at $N = 82$ and $N = 126$, along with candidate actinide chains approaching $N \approx 184$.

Imposing that the $N = 82$ and $N = 126$ peaks be sustained under identical thermodynamic conditions defines a common interval in $S_{2n}/2$ that characterizes the equilibrium r-process path.

Within the FRDM2012 mass model, only the actinide isotopic chains with $Z = 98, 100,$ and 102 (Cf, Fm, and No) satisfy this structural criterion near $N \approx 184$. These chains are therefore identified as the heavy-region CWP nuclei.

The corresponding S_n^0 intervals trace loci in the (T_9, n_n) plane, as shown in Fig. 3. Their overlap defines the core band of thermodynamic conditions for the $N \approx 184$ region.

To account for residual variations in separation-energy systematics and deviations from ideal equilibrium, we adopt an effective tolerance of $\Delta S = 0.2$ MeV. This choice is motivated by the propagation of uncertainties from nuclear masses to S_{2n} ; because S_{2n} is defined as a difference, the typical uncertainty on $S_{2n}/2$ is at the level of a few tenths of MeV. Given the relatively smooth systematics near $N \approx 184$ in FRDM2012, this value is a representative tolerance that captures the expected scale of

model uncertainties and structural variations around the CWP nuclei.

For clarity and to focus on the heavy-region conditions relevant for actinide nucleosynthesis, the thermodynamic domain is restricted to $1 \leq T_9 \leq 2$ and $10^{20} \leq n_n \leq 10^{28} \text{ cm}^{-3}$, consistent with the typical parameter ranges adopted in previous studies of the main r-process. This range constitutes a subset of the broader r-process parameter space ($T_9 \sim 1-3$) and focuses on the region most relevant to the present analysis [33].

This tolerance defines the effective separation-energy window,

$$2.6 \text{ MeV} \leq S_{2n}/2 \leq 3.045 \text{ MeV}.$$

This characterizes both the structural consistency and the thermodynamic domain of equilibrium r-process flow in the $N \approx 184$ region.

The restriction to $Z = 98, 100,$ and 102 reflects the separation-energy systematics predicted by FRDM2012; the sensitivity of this selection to alternative mass models is discussed in Sec. V.A.

B. Equilibrium path band under CWP T_9-n_n conditions

Using the calibrated separation-energy window derived in Sec. IV. A, we construct the equilibrium path band by applying the structural criterion defined in Sec. III. E.

A nucleus is included in the band if its $S_{2n}/2$ value lies within the heavy-region window $2.6 \text{ MeV} \leq S_{2n}/2 \leq 3.045 \text{ MeV}$. The proton-number range is extended from the reference CWP chains ($Z = 98, 100, 102$) to the broader actinide region, covering $94 \leq Z \leq 106$, without imposing an odd-even restriction.

Because the calibrated window already accounts for the relevant variations in separation-energy systematics

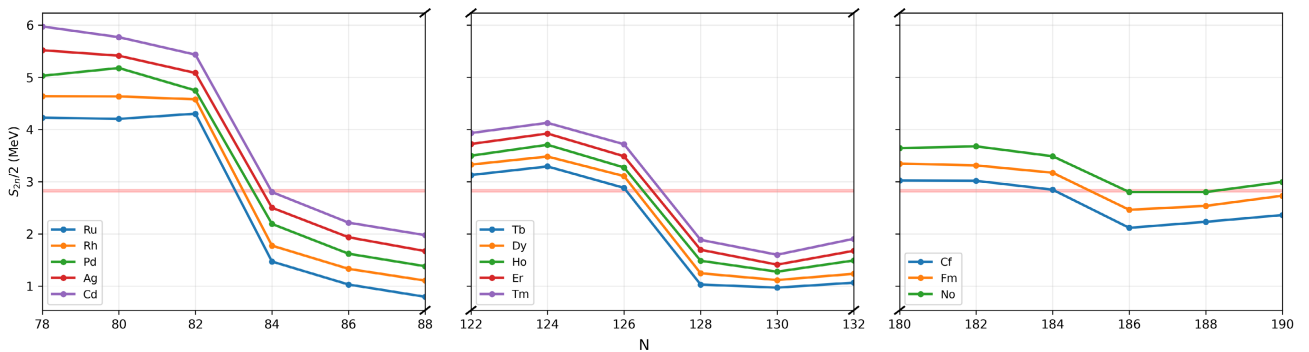


Fig. 2. (color online) Shown are the systematics of the two-neutron separation-energy metric $S_{2n}/2$ along selected isotopic chains near the classical waiting points (CWPs) at $N = 82$ and $N = 126$, together with candidate actinide chains approaching $N \approx 184$. The shaded horizontal band indicates the common CWP S_n^0 window required to sustain the abundance peaks at $N = 82$ and $N = 126$ under identical thermodynamic conditions. Applying this structural consistency criterion to the $N \approx 184$ region selects only the $Z = 98, 100,$ and 102 chains (Cf, Fm, and No), whereas neighboring chains deviate systematically as N approaches 184.

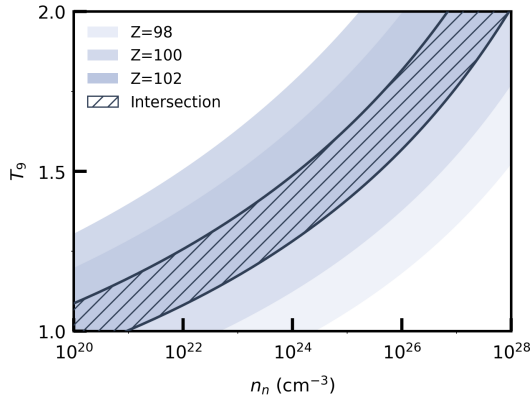


Fig. 3. (color online) Thermodynamic conditions (T_9, n_n) required by the heavy-region CWP nuclei near $N \approx 184$. Shaded bands show the individual $T_9 - n_n$ loci for the $Z = 98, 100,$ and 102 isotopic chains (Cf, Fm, and No). The hatched region indicates their overlap, which defines the condition band for equilibrium r-process flow in the heavy actinide region.

and for moderate deviations from ideal equilibrium, no further broadening of the path band is applied.

Figure 4 shows the resulting equilibrium path band superimposed on the nuclear chart for $94 \leq Z \leq 106$. Colored tiles denote nuclei selected by the heavy-region structural criterion, with the color scale indicating the relative competition between neutron-induced fission and β -decay (see Sec. IV.C).

The band spans the full actinide region and is centered near the reference CWP chains at $Z = 98, 100,$ and 102 . It defines the nuclear domain for subsequent lifetime comparisons and for identifying the onset of fission cycling.

C. Onset of fission cycling along the equilibrium path band: representative environments

We investigate the onset of fission cycling along the equilibrium path band (Sec. IV.B) under three representative thermodynamic conditions drawn from the heavy-region CWP (T_9, n_n) domain shown in Fig. 3.

These points span the band of conditions: $T_9 = 1.0, n_n = 7.0 \times 10^{19} \text{ cm}^{-3}$; $T_9 = 1.5, n_n = 7.0 \times 10^{24} \text{ cm}^{-3}$; and $T_9 = 2.0, n_n = 2.5 \times 10^{27} \text{ cm}^{-3}$.

For each (T_9, n_n) pair, we evaluate the neutron-induced fission and β -decay lifetimes as described in Sec. III. F and identify the onset of fission cycling along the path band.

The results are shown in Fig. 4, where blue regions indicate fission-dominated nuclei and red regions indicate β -decay-dominated nuclei. According to the operational criterion defined in Sec. III.F, the onset nuclides for the three representative conditions are ^{290}Es , ^{279}Cf , and ^{275}Cm , respectively.

At the lowest neutron density, β -decay dominates over most of the path band, and neutron-induced fission

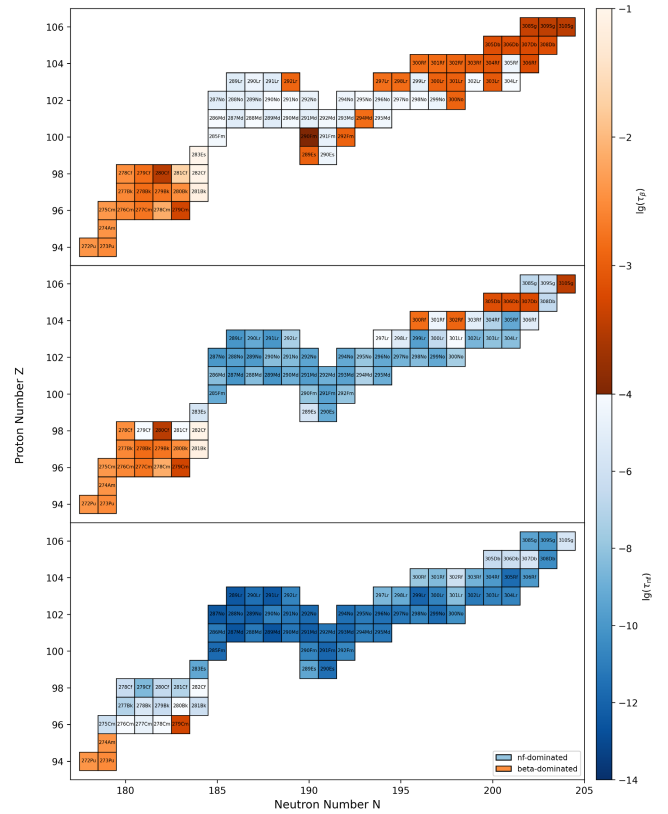


Fig. 4. (color online) Neutron-induced fission and β -decay lifetimes along the equilibrium path band in the actinide region. The three panels correspond to representative thermodynamic conditions (top to bottom): $T_9 = 1.0, n_n = 7.0 \times 10^{19} \text{ cm}^{-3}$; $T_9 = 1.5, n_n = 7.0 \times 10^{24} \text{ cm}^{-3}$; and $T_9 = 2.0, n_n = 2.5 \times 10^{27} \text{ cm}^{-3}$. Each tile represents a nucleus within the path band defined in Sec. IV. B. Colors indicate the relative dominance of neutron-induced fission and β -decay ($\tau_{n,f}$ vs. τ_β), with blue denoting fission-dominated nuclei and red denoting β -decay-dominated nuclei.

becomes competitive only at relatively high Z and mass number A . As temperature and neutron density increase, the fission-dominated region expands systematically toward lower Z and lower mass number A , shifting the onset of fission cycling from ^{290}Es to ^{279}Cf and finally to ^{275}Cm .

This trend indicates that more extreme thermodynamic conditions lead to an earlier onset of fission cycling along the equilibrium path band.

D. Global mapping of the fission-cycling onset across the $T_9 - n_n$ conditions

We extend the lifetime analysis across the full $T_9 - n_n$ parameter space and construct a global map of the onset of fission cycling in the actinide region.

Figure 5 shows the onset nucleus identified at each thermodynamic point (T_9, n_n) within the CWP domain. Each colored point corresponds to a distinct set of condi-

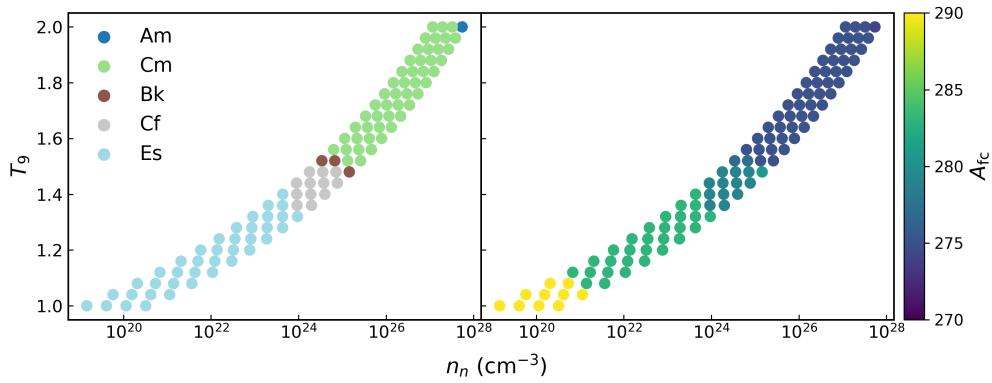


Fig. 5. (color online) Onset of fission cycling across the T_9 - n_n parameter space. **Left:** Proton number Z_{fc} of the nucleus at onset. **Right:** Corresponding mass number A_{fc} .

tions, with the onset characterized by its proton number Z_{fc} and mass number A_{fc} .

The left panel displays Z_{fc} , while the right panel shows the corresponding A_{fc} .

The full set of (T_9, n_n) points, together with the corresponding values of Z_{fc} and A_{fc} , is listed in Appendix A (Table A1).

The map indicates that the onset of fission cycling spans a well-defined region of the actinide chart, extending from americium (Am) to einsteinium (Es). As temperature and neutron density increase, the onset shifts systematically toward lower mass numbers, from $A \approx 290$ to $A \approx 270$.

Rather than occurring at a fixed neutron number, the onset follows a continuous trajectory along the equilibrium boundary, reflecting the strong environmental dependence of the competition between neutron-induced fission and β -decay.

V. DISCUSSION

The framework developed here combines structural diagnostics, the T_9 - n_n condition band, equilibrium-path filtering, and point-by-point lifetime comparisons to identify the onset of fission cycling under equilibrium r-process conditions.

It applies to regimes in which an approximate $(n, \gamma) \rightleftharpoons (\gamma, n)$ equilibrium holds and a waiting-point-like description of the reaction flow remains valid [3, 4, 34]. Such conditions are expected during the high-temperature, high-neutron-density phases of the r-process, for example, in the early, hot stages of compact-object merger ejecta [16].

We now examine the key assumptions and nuclear-physics uncertainties that determine the robustness of the inferred fission-onset boundary.

A. Mass-model dependence and robustness

Within the FRDM2012 mass model adopted here, both the extent of the T_9 - n_n condition band and the set of

actinide isotopic chains included in the equilibrium-path band are determined by the systematics of the predicted neutron-separation energies.

The structural diagnostic introduced in Sec. IV.A, based on the $S_{2n}/2$ metric near $N \approx 184$, selects a subset of even- Z chains for which the separation energies vary smoothly.

The effective bandwidth ΔS is chosen to encompass the spread in S_{2n} among the selected isotopic chains, while reflecting both intrinsic variations in the separation-energy systematics and the propagation of nuclear-mass uncertainties to S_{2n} ; because of its difference form, the latter sets a characteristic scale of a few tenths of an MeV. It thus provides a consistent measure of the structural uncertainty associated with the heavy-region condition band.

Across the condition band, the equilibrium separation energy S_n^0 varies systematically with (T_9, n_n) , while the relative positioning of the selected isotopic chains is essentially unchanged. As a result, the location of the fission-cycling onset is governed primarily by the competition between $\tau_{n,f}$ (which depends strongly on n_n) and τ_β , rather than by moderate variations in the separation-energy systematics. This suggests that the inferred fission-onset boundary is relatively robust against moderate mass-model uncertainties, provided that the global S_{2n} trends near $N \approx 184$ are reasonably reproduced.

Although FRDM2012 is adopted here as a consistent structural baseline, the framework itself is not tied to a specific mass model and can be applied to other global mass models. Systematic comparisons with alternative mass tables will be carried out in future work to further quantify the robustness of the condition-band construction and the resulting fission-onset map.

B. Sensitivity to fission-barrier systematics

The neutron-induced fission rates calculated with TALYS in this work depend on the adopted fission-barrier systematics. The baseline onset map presented in Sec. IV uses the standard TALYS prescription, combining experi-

mentally constrained barriers where available (`fismodel = 1`) with a rotating liquid-drop description as a fallback (`fismodelalt = 4`).

To assess this dependence, we performed a complementary set of calculations using a fully theoretical barrier set, based on the macroscopic–microscopic barriers of Mamdouh *et al.* (`fismodel = 2`) together with the Rotating Finite-Range Model of Sierk (`fismodelalt = 3`), and recomputed the onset map on the same (T_9, n_n) grid.

The resulting onset nuclei exhibit systematic yet bounded shifts under the alternative barrier prescription. Across all sampled thermodynamic points, the typical shift is +1 in both Z and A , with roughly one-third of the points remaining unchanged. The effect is predominantly local: 98.7% of the grid points satisfy $|\Delta Z| \leq 3$ and $|\Delta A| \leq 8$, and only two points show larger deviations, confined to a narrow region in which the $\tau_{n,f} - \tau_\beta$ crossing is particularly steep.

These results indicate that, although fission-barrier systematics can significantly modify individual fission lifetimes, their impact on the fission-onset boundary remains limited. The changes are restricted to local reassignments among neighboring actinide nuclei, without altering the global structure of the onset map across the equilibrium-path band.

VI. CONCLUSION

We have developed a framework to identify the onset

of fission cycling in the r-process across a broad range of astrophysical conditions. By combining structural diagnostics, thermodynamic constraints in the $T_9 - n_n$ plane, equilibrium-path filtering, and lifetime comparisons, the framework provides a systematic approach to determine when fission begins to influence the r-process flow.

Applying this framework to the actinide region near $N \approx 184$, we find a clear, systematic trend across the $T_9 - n_n$ band of conditions. As temperature and neutron density increase, the onset shifts toward lighter actinides along the equilibrium path, with the dominant fissioning nuclei evolving from heavier species (*e.g.*, Es) to Am and neighboring elements. The corresponding mass numbers decrease from $A \approx 290$ toward $A \approx 270$.

These trends are consistent with previous r-process studies that place the dominant fissioning region near the $N \sim 184$ shell closure, with typical masses around $A \approx 280$ [3, 35]. Within this framework, our analysis further shows that the onset of fission cycling is governed primarily by the competition between neutron-induced fission and β -decay, rather than by shifts in the equilibrium path itself.

The framework is flexible and can be extended to alternative nuclear inputs and astrophysical conditions, providing a transparent basis for future studies of fission feedback and actinide production in heavy-element nucleosynthesis.

A. Numerical Data for Fission-Cycling Onset Nuclei

Table A1. Numerical values for the nuclei at the onset of fission cycling, $(A_{fc}, Z_{fc}, \text{elem}_{fc})$, are listed as functions of the temperature T_9 (in GK) and the neutron density n_n (in cm^{-3}), corresponding to the results shown in Fig. 5. For each temperature T_9 , two rows are used, with two sets of $(n_n, \log_{10} n_n, A_{fc}, Z_{fc}, \text{elem}_{fc})$ per row.

T_9 /GK	n_n / cm^{-3}	$\log_{10} n_n$	A_{fc}	Z_{fc}	elem_{fc}	n_n / cm^{-3}	$\log_{10} n_n$	A_{fc}	Z_{fc}	elem_{fc}
1.000	1.457×10^{19}	19.163	290	99	Es	4.092×10^{19}	19.612	290	99	Es
1.000	1.149×10^{20}	20.060	290	99	Es	3.229×10^{20}	20.509	290	99	Es
1.040	5.780×10^{19}	19.762	290	99	Es	1.560×10^{20}	20.193	290	99	Es
1.040	4.212×10^{20}	20.625	290	99	Es	1.137×10^{21}	21.056	290	99	Es
1.080	2.075×10^{20}	20.317	290	99	Es	5.400×10^{20}	20.732	290	99	Es
1.080	1.405×10^{21}	21.148	283	99	Es	3.657×10^{21}	21.563	283	99	Es
1.120	6.814×10^{20}	20.833	283	99	Es	1.714×10^{21}	21.234	283	99	Es
1.120	4.310×10^{21}	21.634	283	99	Es	1.084×10^{22}	22.035	283	99	Es
1.160	2.065×10^{21}	21.315	283	99	Es	5.031×10^{21}	21.702	283	99	Es
1.160	1.226×10^{22}	22.088	283	99	Es	2.986×10^{22}	22.475	283	99	Es
1.200	5.823×10^{21}	21.765	283	99	Es	1.377×10^{22}	22.139	283	99	Es
1.200	3.257×10^{22}	22.513	283	99	Es	7.701×10^{22}	22.887	283	99	Es
1.240	1.538×10^{22}	22.187	283	99	Es	3.538×10^{22}	22.549	283	99	Es
1.240	8.138×10^{22}	22.911	283	99	Es	1.872×10^{23}	23.272	283	99	Es
1.280	3.830×10^{22}	22.583	283	99	Es	8.582×10^{22}	22.934	283	99	Es

Continued on next page

Table A1-continued from previous page

T_9 /GK	n_n / cm^{-3}	$\log_{10} n_n$	A_{fc}	Z_{fc}	elem_{fc}	n_n / cm^{-3}	$\log_{10} n_n$	A_{fc}	Z_{fc}	elem_{fc}
1.280	1.923×10^{23}	23.284	283	99	Es	4.310×10^{23}	23.634	283	99	Es
1.320	9.035×10^{22}	22.956	283	99	Es	1.976×10^{23}	23.296	283	99	Es
1.320	4.321×10^{23}	23.636	283	99	Es	9.449×10^{23}	23.975	283	99	Es
1.360	2.029×10^{23}	23.307	283	99	Es	4.337×10^{23}	23.637	283	99	Es
1.360	9.268×10^{23}	23.967	279	98	Cf	1.981×10^{24}	24.297	279	98	Cf
1.400	4.357×10^{23}	23.639	283	99	Es	9.112×10^{23}	23.960	279	98	Cf
1.400	1.905×10^{24}	24.280	279	98	Cf	3.985×10^{24}	24.600	279	98	Cf
1.440	8.977×10^{23}	23.953	279	98	Cf	1.839×10^{24}	24.265	279	98	Cf
1.440	3.768×10^{24}	24.576	279	98	Cf	7.720×10^{24}	24.888	279	98	Cf
1.480	1.781×10^{24}	24.251	279	98	Cf	3.578×10^{24}	24.554	279	98	Cf
1.480	7.191×10^{24}	24.857	279	98	Cf	1.445×10^{25}	25.160	281	97	Bk
1.520	3.411×10^{24}	24.533	277	97	Bk	6.730×10^{24}	24.828	277	97	Bk
1.520	1.328×10^{25}	25.123	275	96	Cm	2.619×10^{25}	25.418	275	96	Cm
1.560	6.326×10^{24}	24.801	275	96	Cm	1.227×10^{25}	25.089	275	96	Cm
1.560	2.378×10^{25}	25.376	275	96	Cm	4.611×10^{25}	25.664	275	96	Cm
1.600	1.139×10^{25}	25.056	275	96	Cm	2.171×10^{25}	25.337	275	96	Cm
1.600	4.141×10^{25}	25.617	275	96	Cm	7.896×10^{25}	25.897	275	96	Cm
1.640	1.993×10^{25}	25.300	275	96	Cm	3.742×10^{25}	25.573	275	96	Cm
1.640	7.024×10^{25}	25.847	275	96	Cm	1.319×10^{26}	26.120	275	96	Cm
1.680	3.401×10^{25}	25.532	275	96	Cm	6.289×10^{25}	25.799	275	96	Cm
1.680	1.163×10^{26}	26.066	275	96	Cm	2.151×10^{26}	26.333	275	96	Cm
1.720	5.664×10^{25}	25.753	275	96	Cm	1.033×10^{26}	26.014	275	96	Cm
1.720	1.882×10^{26}	26.275	275	96	Cm	3.432×10^{26}	26.535	275	96	Cm
1.760	9.225×10^{25}	25.965	275	96	Cm	1.659×10^{26}	26.220	275	96	Cm
1.760	2.983×10^{26}	26.475	275	96	Cm	5.365×10^{26}	26.730	275	96	Cm
1.800	1.471×10^{26}	26.168	275	96	Cm	2.612×10^{26}	26.417	275	96	Cm
1.800	4.636×10^{26}	26.666	275	96	Cm	8.228×10^{26}	26.915	275	96	Cm
1.840	2.301×10^{26}	26.362	275	96	Cm	4.034×10^{26}	26.606	275	96	Cm
1.840	7.072×10^{26}	26.850	275	96	Cm	1.240×10^{27}	27.093	275	96	Cm
1.880	3.534×10^{26}	26.548	275	96	Cm	6.121×10^{26}	26.787	275	96	Cm
1.880	1.060×10^{27}	27.025	275	96	Cm	1.837×10^{27}	27.264	275	96	Cm
1.920	5.334×10^{26}	26.727	275	96	Cm	9.134×10^{26}	26.961	275	96	Cm
1.920	1.564×10^{27}	27.194	275	96	Cm	2.679×10^{27}	27.428	275	96	Cm
1.960	7.922×10^{26}	26.899	275	96	Cm	1.342×10^{27}	27.128	275	96	Cm
1.960	2.273×10^{27}	27.357	275	96	Cm	3.850×10^{27}	27.585	275	96	Cm
2.000	1.159×10^{27}	27.064	275	96	Cm	1.942×10^{27}	27.288	275	96	Cm
2.000	3.255×10^{27}	27.513	275	96	Cm	5.456×10^{27}	27.737	274	95	Am

References

- [1] E. M. Burbidge, G. R. Burbidge, W. A. Fowler *et al.*, *Rev. Mod. Phys.* **29**(4), 547 (1957)
- [2] J. J. Cowan, F.-K. Thielemann and J. W. Truran, *Physics Reports* **208**(4), 267 (1991)
- [3] J. J. Cowan, C. Sneden, J. E. Lawler *et al.*, *Rev. Mod. Phys.* **93**(1), 015002 (2021)
- [4] M. Arnould, S. Goriely and K. Takahashi, *Phys. Rep.* **450**(4-6), 97 (2007)
- [5] I. V. Panov, E. Kolbe, B. Pfeiffer *et al.*, *Nucl. Phys. A* **747**(2), 633 (2005)
- [6] S. Goriely, J.-L. Sida, J.-F. Lemaître *et al.*, *Phys. Rev. Lett.* **111**(24), 242502 (2013)
- [7] M. R. Mumpower, R. Surman, G. C. McLaughlin *et al.*, *Prog. Part. Nucl. Phys.* **86**, 86 (2016)
- [8] J. Beun, G. C. McLaughlin, R. Surman *et al.*, *Phys. Rev. C* **77**(3), 035804 (2008)
- [9] O. Korobkin, S. Rosswog, A. Arcones *et al.*, *Mon. Not. R. Astron. Soc.* **426**(3), 1940 (2012)
- [10] M. Eichler, A. Arcones, A. Kelic *et al.*, *Astrophys. J.* **808**(1), 30 (2015)
- [11] S. Goriely, *The European Physical Journal A* **51**(2), 22 (2015)
- [12] I. V. Panov, I. Y. Korneev, T. Rauscher *et al.*, *Astron. Astrophys.* **513**, A61 (2010)
- [13] J. Erler, N. Birge, M. Kortelainen *et al.*, *Nature* **486**(7404), 509 (2012)
- [14] N. Schunck and L. M. Robledo, *Rep. Prog. Phys.* **79**(11), 116301 (2016)
- [15] W.-P. Liu, B. Guo, Z. An *et al.*, *Nucl. Sci. Tech.* **35**(12), 217 (2024)
- [16] C. Freiburghaus, S. Rosswog and F.-K. Thielemann, *Astrophys. J.* **525**, L121 (1999)
- [17] S. Wanajo, Y. Sekiguchi, N. Nishimura *et al.*, *Astrophys. J. Lett.* **789**(2), L39 (2014)
- [18] J. D. J. Mendoza-Temis, M.-R. Wu, K. Langanke *et al.*, *Phys. Rev. C* **92**(5), 055805 (2015)
- [19] J. Lippuner and L. F. Roberts, *Astrophys. J.* **815**(2), 82 (2015)
- [20] I. Petermann, K. Langanke, G. Martínez-Pinedo *et al.*, *Phys. Rev. C* **81**, 014308 (2010)
- [21] T. Kodama and K. Takahashi, *Nucl. Phys. A* **239**, 489 (1975)
- [22] S. Goriely and M. Arnould, *Astron. Astrophys.* **312**, 327 (1996)
- [23] P. Möller, W. D. Myers, H. Sagawa *et al.*, *Phys. Rev. Lett.* **108**(5), 052501 (2012)
- [24] P. Möller, A. J. Sierk, T. Ichikawa *et al.*, *At. Data Nucl. Data Tables* **109-110**, 1 (2016)
- [25] A. Koning, S. Hilaire and S. Goriely, *The European Physical Journal A* **59**(6), 131 (2023)
- [26] T. Marketin, L. Huther and G. Martínez-Pinedo, *Phys. Rev. C* **93**(2), 025805 (2016)
- [27] X. D. Xu, B. Sun, Z. M. Niu *et al.*, *Phys. Rev. C* **87**(1), 015805 (2013)
- [28] W. Hillebrandt, F. K. Thielemann, H. V. Klapdor *et al.*, *Astron. Astrophys.* **99**, 195 (1981)
- [29] S. Goriely, N. Chamel and J. M. Pearson, *Phys. Rev. C* **88**(6), 061302 (2013)
- [30] J. J. Li, W. H. Long, J. Margueron *et al.*, *Phys. Lett. B* **732**, 169 (2014)
- [31] Q. Mo, M. Liu, and N. Wang, *Phys. Rev. C* **90**(2), 024320 (2014)
- [32] N. Wang, M. Liu, X. Wu *et al.*, *Phys. Rev. C* **93**(1), 014302 (2016)
- [33] F.-K. Thielemann, A. Arcones, R. Käppeli *et al.*, *Prog. Part. Nucl. Phys.* **66**, 346 (2011)
- [34] T. Kajino, W. Aoki, A. Balantekin *et al.*, *Prog. Part. Nucl. Phys.* **107**, 109 (2019)
- [35] S. A. Giuliani, G. Martínez-Pinedo, M.-R. Wu *et al.*, *Phys. Rev. C* **102**(4), 045804 (2020)