# Oblique parameters in gauged baryon and lepton numbers with a 125 GeV Higgs＊ 

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#### Abstract

In an extension of the standard model，where baryon number and lepton number are local gauge sym－ metries，we analyze the effect of corrections from exotic fermions and scalars on the oblique parameters $S, T, U$ ． Because a light neutral Higgs $h_{0}$ with mass around $124-126 \mathrm{GeV}$ strongly constrains the corresponding parameter space of this model，we also investigate the gluon fusion process $\mathrm{gg} \rightarrow \mathrm{h}_{0}$ and two photon decay of the lightest neutral Higgs $\mathrm{h}_{0} \rightarrow \gamma \gamma$ at the Large Hadron Collider．


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## 1 Introduction

The main physics goals of the Large Hadron Collider （LHC）are to understand the origin of electroweak sym－ metry breaking，and to search for the neutral Higgs bo－ son predicted by the standard model（SM）and its var－ ious extensions．Recently，ATLAS and CMS have re－ ported significant excess events that are interpreted to be most probably related to a neutral Higgs with mass $m_{\mathrm{h}_{0}} \sim 124-126 \mathrm{GeV}$ ．This implies that the Higgs mech－ anism of breaking electroweak symmetry possibly has a solid experimental cornerstone．

The oblique parameters $S, T, U$［1］are extracted from electroweak precision data（EWPD）observations that probe the radiative corrections with sufficient ac－ curacy．A light neutral Higgs with mass $124-126 \mathrm{GeV}$ also affects the theoretical evaluations of the oblique pa－ rameters $S, T, U$ through loop corrections to the gauge boson propagators，which contain the neutral Higgs as a virtual field．In extensions of the SM，the corrections from the exotic fields to the gauge boson propagators can be expressed in terms of shifts of the parameters $S, T$ ， $U$［2］．

A broken baryon number $(B)$ conservation can ex－ plain the origin of the matter－antimatter asymmetry in the Universe in a natural way．The heavy majorana neu－
trinos contained in the seesaw mechanism can induce the tiny observed neutrino masses［3］to explain the results of neutrino oscillation experiments．Hence，the lepton num－ ber $(L)$ is also expected to be broken．In Ref．［4］，two extensions to the SM are examined，where $B$ and $L$ are spontaneously broken gauge symmetries around the TeV scale，while Ref．［5］also investigates the predictions for the Higgs mass and the Higgs decays in a supersymmetric model named BLMSSM，which is a minimal supersym－ metric extension of the SM（MSSM）with local gauged $B$ and $L$ ．Within the framework of the first extension of the SM with spontaneously broken $B$ and $L$［4］，we ana－ lyze the gluon fusion production and then decay into two photons of the Higgs with mass $m_{\mathrm{h}_{0}} \sim 124-126 \mathrm{GeV}$ ．Ad－ ditionally，we also investigate the corrections from exotic fields of the oblique parameters $S, T, U$ ．

This paper is organized as follows．In Section 2，we briefly summarize the main ingredients of an extension of the SM where the baryon and lepton numbers are local symmetries，we then present the mass squared matrices for the neutral Higgs sector．Inspired by the new results from the ATLAS and CMS collaborations，in Section 3 we study in great detail the Higgs production through gluon fusion，followed by the decay of the Higgs boson into two photons．We discuss the constraints on the pa－ rameter space from the oblique parameters $S, T, U$ in Section 4．Our conclusions are given in Section 5.

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## 2 An extension of the SM where baryon and lepton numbers are local gauge symmetries

When baryon and lepton numbers are local gauge symmetries, one can write the gauge group as $S U(3)_{\mathrm{C}} \otimes$ $S U(2)_{\mathrm{L}} \otimes U(1)_{\mathrm{Y}} \otimes U(1)_{\mathrm{B}} \otimes U(1)_{\mathrm{L}}$. In the first extension of the SM proposed in Ref. [4], the exotic particles include new quarks $\mathrm{Q}_{\mathrm{L}}^{\prime}, \mathrm{u}_{\mathrm{R}}^{\prime}, \mathrm{d}_{\mathrm{R}}^{\prime}$, new leptons $\mathrm{l}_{\mathrm{L}}^{\prime}, \mathrm{v}_{\mathrm{R}}^{\prime}, \mathrm{e}_{\mathrm{R}}^{\prime}$ and three scalar singlets $S_{\mathrm{B}}, S_{\mathrm{L}}, S$ along with a scalar doublet $\phi$. The Yukawa couplings are written as

$$
\begin{align*}
-\mathcal{L}_{\mathrm{Y}}= & \left\{\sum_{I, J=1}^{3}\left[\left(Y_{\mathrm{U}}\right)_{I J} \bar{Q}_{\mathrm{L}}^{I} \tilde{H} u_{\mathrm{R}}^{J}+\left(Y_{\mathrm{D}}\right)_{I J} \bar{Q}_{\mathrm{L}}^{I} H d_{\mathrm{R}}^{J}\right]\right. \\
& +Y_{\mathrm{U}}^{\prime} \bar{Q}_{\mathrm{L}}^{\prime} Q_{\mathrm{L}}^{\prime \mathrm{c}} S_{\mathrm{B}}+\sum_{I=1}^{3}\left[\left(Y_{1}\right)_{I} \bar{Q}_{\mathrm{L}}^{\prime} \tilde{\phi} u_{\mathrm{R}}^{J}+\left(Y_{2}\right)_{I} \bar{Q}_{\mathrm{L}}^{I} \phi d_{\mathrm{R}}^{\prime}\right] \\
& + \text { h.c. }\}+\left\{\sum_{I, J=1}^{3}\left[\left(Y_{\nu}\right)_{I J} \bar{L}_{\mathrm{L}}^{I} \tilde{H} \nu_{\mathrm{R}}^{J}+\left(Y_{\mathrm{E}}\right)_{I J} \bar{L}_{\mathrm{L}}^{I} H e_{\mathrm{R}}^{J}\right]\right. \\
& +Y_{\mathrm{E}}^{\prime} \bar{L}_{\mathrm{L}}^{\prime} L_{\mathrm{L}}^{\prime \mathrm{c}} S_{\mathrm{L}}+\frac{1}{2} \sum_{I, J=1}^{3}\left(\lambda_{\mathrm{a}}\right)_{I J} \bar{\nu}_{\mathrm{R}}^{I, \mathrm{c}} S_{\mathrm{L}}^{*} \nu_{\mathrm{R}}^{J} \\
& \left.+\sum_{I=1}^{3}\left(\lambda_{\mathrm{b}}\right)_{I} \overline{\mathrm{~L}}_{\mathrm{R}}^{I, \mathrm{c}} S_{\mathrm{L}} \nu_{\mathrm{R}}^{\prime}+\text { h.c. }\right\} \tag{1}
\end{align*}
$$

The scalar potential is generally given as follows:

$$
\begin{align*}
-\mathcal{L}_{\mathrm{S}}= & m_{\mathrm{H}}^{2} H^{\dagger} H+m_{\phi}^{2} \phi^{\dagger} \phi+m_{S_{\mathrm{B}}}^{2} S_{\mathrm{B}}^{*} S_{\mathrm{B}}+m_{S_{\mathrm{L}}}^{2} S_{\mathrm{L}}^{*} S_{\mathrm{L}} \\
& +m_{\mathrm{S}}^{2} S^{*} S+\lambda_{\mathrm{HH}}\left(H^{\dagger} H\right)\left(H^{\dagger} H\right)+\lambda_{\phi \phi}\left(\phi^{\dagger} \phi\right)\left(\phi^{\dagger} \phi\right) \\
& +\lambda_{\mathrm{BB}}\left(S_{\mathrm{B}}^{*} S_{\mathrm{B}}\right)\left(S_{\mathrm{B}}^{*} S_{\mathrm{B}}\right)+\lambda_{\mathrm{LL}}\left(S_{\mathrm{L}}^{*} S_{\mathrm{L}}\right)\left(S_{\mathrm{L}}^{*} S_{\mathrm{L}}\right) \\
& +\lambda_{\mathrm{SS}}\left(S^{*} S\right)\left(S^{*} S\right)+\lambda_{\mathrm{H} \phi}\left(H^{\dagger} H\right)\left(\phi^{\dagger} \phi\right) \\
& +\lambda_{\mathrm{HB}}\left(H^{\dagger} H\right)\left(S_{\mathrm{B}}^{*} S_{\mathrm{B}}\right)+\lambda_{\mathrm{HL}}\left(H^{\dagger} H\right)\left(S_{\mathrm{L}}^{*} S_{\mathrm{L}}\right) \\
& +\lambda_{\mathrm{HS}}\left(H^{\dagger} H\right)\left(S^{*} S\right)+\lambda_{\phi \mathrm{B}}\left(\phi^{\dagger} \phi\right)\left(S_{\mathrm{B}}^{*} S_{\mathrm{B}}\right) \\
& +\lambda_{\phi \mathrm{L}}\left(\phi^{\dagger} \phi\right)\left(S_{\mathrm{L}}^{*} S_{\mathrm{L}}\right)+\lambda_{\phi \mathrm{S}}\left(\phi^{\dagger} \phi\right)\left(S^{*} S\right) \\
& +\lambda_{\mathrm{BL}}\left(S_{\mathrm{B}}^{*} S_{\mathrm{B}}\right)\left(S_{\mathrm{L}}^{*} S_{\mathrm{L}}\right)+\lambda_{\mathrm{BS}}\left(S_{\mathrm{B}}^{*} S_{\mathrm{B}}\right)\left(S^{*} S\right) \\
& +\lambda_{\mathrm{LS}}\left(S_{\mathrm{L}}^{*} S_{\mathrm{L}}\right)\left(S^{*} S\right)+\lambda_{\mathrm{H} \phi}^{\prime}\left(H^{\dagger} \phi\right)\left(\phi^{\dagger} H\right) \\
& +\left[\mu_{1}\left(H^{\dagger} \phi\right) S+\mu_{2} S_{\mathrm{B}}^{*} S^{2}+\text { h.c. }\right] . \tag{2}
\end{align*}
$$

When the $S U(2)_{\mathrm{L}}$ doublet $H$ and $S U(2)_{\mathrm{L}}$ singlets $S_{\mathrm{B}}, S_{\mathrm{L}}$ acquire the nonzero vacuum expectation values (VEVs)

## $v, v_{\mathrm{B}, \mathrm{L}}$,

$$
\begin{align*}
H & =\binom{G^{+}}{\frac{1}{\sqrt{2}}\left(v+H_{0}+\mathrm{i} G^{0}\right)} \\
S_{\mathrm{B}} & =\frac{1}{\sqrt{2}}\left(v_{\mathrm{B}}+S_{\mathrm{B}}^{0}+\mathrm{i} G_{\mathrm{B}}^{0}\right) \\
S_{\mathrm{L}} & =\frac{1}{\sqrt{2}}\left(v_{\mathrm{L}}+S_{\mathrm{L}}^{0}+\mathrm{i} G_{\mathrm{L}}^{0}\right) \tag{3}
\end{align*}
$$

the local gauge symmetry $S U(2)_{\mathrm{L}} \otimes U(1)_{\mathrm{Y}} \otimes U(1)_{\mathrm{B}} \otimes U(1)_{\mathrm{L}}$ is broken down to the electromagnetic symmetry $U(1)_{\mathrm{e}}$, where $G^{+}, G^{0}, G_{\mathrm{B}}^{0}$ and $G_{\mathrm{L}}^{0}$ denote massless Goldstone bosons. Correspondingly, the mass terms for the neutral Higgs are formulated as

$$
-\mathcal{L}_{\text {mass }}^{\mathrm{H}}=\frac{1}{2}\left(H_{0}, S_{\mathrm{B}}^{0}, S_{\mathrm{L}}^{0}\right) m_{\mathrm{CPE}}^{2}\left(\begin{array}{c}
H_{0}  \tag{4}\\
S_{\mathrm{B}}^{0} \\
S_{\mathrm{L}}^{0}
\end{array}\right)
$$

where the symmetric $3 \times 3$ mass squared matrix $m_{\text {CPE }}^{2}$ is

$$
m_{\mathrm{CPE}}^{2}=\left(\begin{array}{ccc}
2 \lambda_{\mathrm{HH}} v^{2} & \lambda_{\mathrm{HB}} v v_{\mathrm{B}} & \lambda_{\mathrm{HL}} v v_{\mathrm{L}}  \tag{5}\\
\lambda_{\mathrm{HB}} v v_{\mathrm{B}} & 2 \lambda_{\mathrm{BB}} v_{\mathrm{B}}^{2} & \lambda_{\mathrm{BL}} v_{\mathrm{B}} v_{\mathrm{L}} \\
\lambda_{\mathrm{HL}} v v_{\mathrm{L}} & \lambda_{\mathrm{BL}} v_{\mathrm{B}} v_{\mathrm{L}} & 2 \lambda_{\mathrm{LL}} v_{\mathrm{L}}^{2}
\end{array}\right) .
$$

Through the orthogonal $3 \times 3$ transformation matrix $Z_{\text {CPE }}$, the mass squared matrix $m_{\text {CPE }}^{2}$ can be diagonalized as

$$
\begin{equation*}
Z_{\mathrm{CPE}}^{\mathrm{T}} m_{\mathrm{CPE}}^{2} Z_{\mathrm{CPE}}=\left(m_{\mathrm{H}_{1}^{0}}^{2}, m_{\mathrm{H}_{2}^{0}}^{2}, m_{\mathrm{H}_{3}^{0}}^{2}\right) \tag{6}
\end{equation*}
$$

where $m_{\mathrm{H}_{1}^{0}}=m_{\mathrm{h}_{0}} \approx 125 \mathrm{GeV}$.
In a similar way, we can write the $S U(2)_{\mathrm{L}}$ doublet $\phi$ and the $S U(2)_{\mathrm{L}}$ singlet $S$ as

$$
\begin{align*}
& \phi=\binom{\phi^{+}}{\frac{1}{\sqrt{2}}\left(\phi_{\mathrm{R}}^{0}+\mathrm{i} \phi_{\mathrm{I}}^{0}\right)},  \tag{7}\\
& S=\frac{1}{\sqrt{2}}\left(S_{\mathrm{R}}^{0}+\mathrm{i} S_{\mathrm{I}}^{0}\right)
\end{align*}
$$

Since the local gauge symmetry $S U(2)_{\mathrm{L}} \otimes U(1)_{\mathrm{Y}} \otimes U(1)_{\mathrm{B}} \otimes$ $U(1)_{\mathrm{L}}$ is broken down to the electromagnetic symmetry $U(1)_{\mathrm{e}}$, the terms in square brackets in Eq. (2) induce mixing among the neutral scalar particles $\phi_{\mathrm{R}}^{0}, \phi_{\mathrm{I}}^{0}, S_{\mathrm{R}}^{0}$, $S_{\mathrm{I}}^{0}$, and the mass terms are written as

$$
-\mathcal{L}_{\text {mass }}^{\Phi}=\frac{1}{2}\left(\phi_{\mathrm{R}}^{0}, S_{\mathrm{R}}^{0}, \phi_{\mathrm{I}}^{0}, S_{\mathrm{I}}^{0}\right) m_{\mathrm{CPM}}^{2}\left(\begin{array}{c}
\phi_{\mathrm{R}}^{0}  \tag{8}\\
S_{\mathrm{R}}^{0} \\
\phi_{\mathrm{I}}^{0} \\
S_{\mathrm{I}}^{0}
\end{array}\right),
$$

with the symmetric $4 \times 4$ mass squared matrix $m_{\text {CPM }}^{2}$ being

$$
m_{\mathrm{CPM}}^{2}=\left(\begin{array}{cccc}
m_{\Phi}^{2} & \sqrt{2} v \Re\left(\mu_{1}\right) & 0 & -\sqrt{2} v \Im\left(\mu_{1}\right)  \tag{9}\\
\sqrt{2} v \Re\left(\mu_{1}\right) & m_{\mathrm{S}}^{2}+2 \sqrt{2} v_{\mathrm{B}} \Re\left(\mu_{2}\right) & 0 & -2 \sqrt{2} v_{\mathrm{B}} \Im\left(\mu_{2}\right) \\
0 & 0 & m_{\Phi}^{2} & -\sqrt{2} v \Re\left(\mu_{1}\right) \\
-\sqrt{2} v \Im\left(\mu_{1}\right) & -2 \sqrt{2} v_{\mathrm{B}} \Im\left(\mu_{2}\right) & -\sqrt{2} v \Re\left(\mu_{1}\right) & m_{\mathrm{S}}^{2}+2 \sqrt{2} v_{\mathrm{B}} \Re\left(\mu_{2}\right)
\end{array}\right) .
$$

We also diagonalize the mass squared matrix $m_{\text {CPM }}^{2}$ through the $4 \times 4$ orthogonal rotation $Z_{\text {СРм }}$ :

$$
\begin{equation*}
Z_{\mathrm{CPM}}^{\mathrm{T}} m_{\mathrm{CPM}}^{2} Z_{\mathrm{CPM}}=\left(m_{\Phi_{1}^{0}}^{2}, m_{\Phi_{2}^{0}}^{2}, m_{\Phi_{3}^{0}}^{2}, m_{\Phi_{4}^{0}}^{2}\right) . \tag{10}
\end{equation*}
$$

Similarly, the mass for the charged scalar $\phi^{ \pm}$is expressed by

$$
\begin{equation*}
m_{\phi^{ \pm}}^{2}=\frac{1}{2} m_{\phi}^{2}-\frac{1}{2} \lambda_{\mathrm{H} \phi}^{\prime} v^{2} \tag{11}
\end{equation*}
$$

Since the field $\phi$ does not get a nonzero VEV after the electroweak symmetry is broken down, there is no mass mixing between the exotic quarks and the SM quarks.

In the left-handed basis $\left(\nu_{\mathrm{L}}^{I}, \nu_{\mathrm{L}}^{\prime}, \nu_{\mathrm{R}}^{\prime c}, \nu_{\mathrm{R}}^{I, c}\right),(I=1,2$, 3), the mass matrix for neutrinos is given by the $8 \times 8$ matrix

$$
\mathcal{M}_{\mathrm{n}}=\left(\begin{array}{cc}
0_{3 \times 3} & \left(M_{\mathrm{D}}\right)_{3 \times 5}  \tag{12}\\
\left(M_{\mathrm{D}}^{\mathrm{T}}\right)_{5 \times 3} & \left(M_{\mathrm{N}}\right)_{5 \times 5}
\end{array}\right)
$$

Here, the $3 \times 5$ matrix $M_{\mathrm{D}}$ is written as

$$
\begin{equation*}
M_{\mathrm{D}}=\left(0_{3 \times 2}, \frac{v}{\sqrt{2}}\left(Y_{v}^{*}\right)_{3 \times 3}\right) \tag{13}
\end{equation*}
$$

and the $5 \times 5$ matrix $M_{\mathrm{N}}$ is

$$
M_{\mathrm{N}}=\left(\begin{array}{ccc}
0, & \frac{v}{\sqrt{2}} Y_{\nu}^{\prime *}, & 0_{1 \times 3}  \tag{14}\\
\frac{v}{\sqrt{2}} Y_{\nu}^{\prime *}, & 0, & \frac{v_{\mathrm{L}}}{\sqrt{2}}\left(\lambda_{\mathrm{b}}^{*}\right)_{1 \times 3} \\
0_{3 \times 1}, & \frac{v_{\mathrm{L}}}{\sqrt{2}}\left(\lambda_{\mathrm{b}}^{\dagger}\right)_{3 \times 1}, & \frac{v_{\mathrm{L}}}{\sqrt{2}}\left(\lambda_{\mathrm{a}}^{\dagger}\right)_{3 \times 3}
\end{array}\right)
$$

By integrating the heavy freedoms out, we get the following mass matrix for three light neutrinos:

$$
\begin{equation*}
\mathcal{M}_{v}=-M_{\mathrm{D}} M_{\mathrm{N}}^{-1} M_{\mathrm{D}}^{\mathrm{T}} \tag{15}
\end{equation*}
$$

which is diagonalized by the Pontecorvo-Maki-Nakagawa-Sakata matrix $U_{\text {PMNS }}$

$$
\begin{equation*}
U_{\mathrm{PMNS}}^{\mathrm{T}} \mathcal{M}_{\nu} U_{\mathrm{PMNS}}=\operatorname{diag}\left(m_{v_{1}}, m_{v_{2}}, m_{v_{3}}\right) \tag{16}
\end{equation*}
$$

Meanwhile, the Majorana mass matrix $M_{\mathrm{N}}$ is similarly diagonalized by a $5 \times 5$ matrix $Z_{\mathrm{N}}$

$$
\begin{equation*}
Z_{\mathrm{N}}^{\mathrm{T}} M_{\mathrm{N}} Z_{\mathrm{N}}=\operatorname{diag}\left(m_{\mathrm{N}_{1}}, m_{\mathrm{N}_{2}}, m_{\mathrm{N}_{3}}, m_{\mathrm{N}_{4}}, m_{\mathrm{N}_{5}}\right) \tag{17}
\end{equation*}
$$

## 3 The $g g \rightarrow h_{0} \rightarrow \gamma \gamma$ process in gauged baryon and lepton numbers

At the LHC, the Higgs is produced chiefly through gluon fusion. In the SM, the leading order (LO) contributions originate from the one-loop diagram, which involves virtual top quarks. The cross section for this process is known as the next-to-next-to-leading order (NNLO) [6], which can enhance the LO result by $80 \%-100 \%$. Furthermore, any new particle that couples strongly with the Higgs can significantly modify this cross section. In the extension of the SM considered here, the LO decay width for the process $\mathrm{h}_{0} \rightarrow \mathrm{gg}$ is given by (see Ref. [7] and the references therein)

$$
\begin{align*}
\Gamma_{\mathrm{NP}}\left(\mathrm{~h}_{0} \rightarrow \mathrm{gg}\right)= & \left.\frac{G_{\mathrm{F}} \alpha_{s}^{2} m_{\mathrm{h}_{0}}^{3}\left|\left(Z_{\mathrm{CPE}}\right)_{11}\right|^{2}}{64 \sqrt{2} \pi^{3}} \right\rvert\, A_{1 / 2}\left(x_{\mathrm{t}}\right) \\
& +A_{1 / 2}\left(x_{t^{\prime}}\right)+\left.A_{1 / 2}\left(x_{\mathrm{b}^{\prime}}\right)\right|^{2} \tag{18}
\end{align*}
$$

where $x_{\mathrm{a}}=m_{\mathrm{h}_{0}}^{2} /\left(4 m_{\mathrm{a}}^{2}\right), \mathrm{a}=\mathrm{t}, \mathrm{t}^{\prime}, \mathrm{b}^{\prime}$, and the loop function $A_{1 / 2}$ is defined as given in the Appendix.

The Higgs to diphoton decay is also obtained from loop diagrams. The LO contributions are derived from the one-loop diagrams containing virtual charged gauge bosons $\mathrm{W}^{ \pm}$or virtual top quarks in the SM. In this model, the additional charged scalar $\phi^{ \pm}$and exotic fermions $\mathrm{t}^{\prime}, \mathrm{b}^{\prime}, \tau^{\prime}$ contribute corrections to the decay width of the Higgs to diphoton at LO. The corresponding expression is written as

$$
\begin{align*}
\Gamma_{\mathrm{NP}}\left(\mathrm{~h}_{0} \rightarrow \gamma \gamma\right)= & \left.\frac{G_{\mathrm{F}} \alpha^{2} m_{\mathrm{h}_{0}}^{3}}{128 \sqrt{2} \pi^{3}} \right\rvert\,\left(Z_{\mathrm{CPE}}\right)_{11}\left(\frac{4}{3} A_{1 / 2}\left(x_{\mathrm{t}}\right)\right. \\
& +\frac{4}{3} A_{1 / 2}\left(x_{\mathrm{t}^{\prime}}\right)+\frac{1}{3} A_{1 / 2}\left(x_{\mathrm{b}^{\prime}}\right) \\
& \left.+A_{1 / 2}\left(x_{\tau^{\prime}}\right)+A_{1}\left(x_{\mathrm{W}}\right)\right) \\
& +\frac{8 m_{\mathrm{W}}^{2} s_{\mathrm{W}}^{2}}{\mathrm{e}^{2} m_{\phi^{ \pm}}^{2}}\left(\lambda_{\mathrm{H} \phi}\left(Z_{\mathrm{CPE}}\right)_{11}\right. \\
& +\frac{v_{\mathrm{B}}}{v} \lambda_{\phi \mathrm{B}}\left(Z_{\mathrm{CPE}}\right)_{21} \\
& \left.+\frac{v_{\mathrm{L}}}{v} \lambda_{\phi \mathrm{L}}\left(Z_{\mathrm{CPE}}\right)_{31}\right)\left.A_{0}\left(x_{\phi^{ \pm}}\right)\right|^{2} \tag{19}
\end{align*}
$$

the concrete expressions for the loop functions $A_{0}, A_{1}$
are given in the Appendix.
The Higgs discoveries from both the ATLAS and CMS experiments have observed an excess in Higgs production and decay into the diphoton channel, which is a factor of 1.4-2 times larger than the SM expectations. The observed signal for the diphoton channels is quantified by the ratio

$$
\begin{equation*}
R_{\gamma \gamma}=\frac{\Gamma_{\mathrm{NP}}\left(\mathrm{~h}_{0} \rightarrow \mathrm{gg}\right) \Gamma_{\mathrm{NP}}\left(\mathrm{~h}_{0} \rightarrow \gamma \gamma\right)}{\Gamma_{\mathrm{SM}}\left(\mathrm{~h}_{0} \rightarrow \mathrm{gg}\right) \Gamma_{\mathrm{SM}}\left(\mathrm{~h}_{0} \rightarrow \gamma \gamma\right)}, \tag{20}
\end{equation*}
$$

where we assume that all exotic fields are heavier than the lightest Higgs $h_{0}$. The current value of this ratio is as follows $[8,9]$ :

$$
\begin{align*}
& \text { ATLAS: } R_{\gamma \gamma}=1.90 \pm 0.5 \\
& \text { CMS: } R_{\gamma \gamma}=1.56 \pm 0.43 \\
& \text { ATLAS+CMS: } R_{\gamma \gamma}=1.71 \pm 0.33 \tag{21}
\end{align*}
$$

Note that the combination of the ATLAS and CMS results is taken from Ref. [10].

## 4 Corrections to the oblique parameters

A common approach to constrain physics beyond the SM is to use global electroweak fitting through the oblique parameters $S, T, U[1]$. In the SM, electroweak precision tests imply a relationship between $m_{\mathrm{h}_{0}}$ and $m_{\mathrm{Z}}$. In the model considered here, the electroweak precision
tests also strongly constrain the mass spectrum and relevant couplings.

Here, we adopt the definitions of the oblique parameters $S, T, U$ given in $[1,11]$ :

$$
\begin{align*}
& S=16 \pi\left\{\frac{\Pi_{33}\left(m_{\mathrm{Z}}^{2}\right)-\Pi_{33}(0)}{m_{\mathrm{Z}}^{2}}-\frac{\Pi_{3 \mathrm{Q}}\left(m_{\mathrm{Z}}^{2}\right)-\Pi_{3 \mathrm{Q}}(0)}{m_{\mathrm{Z}}^{2}}\right\} \\
& T=4 \pi \frac{\Pi_{11}(0)-\Pi_{33}(0)}{m_{\mathrm{Z}}^{2} s_{\mathrm{W}}^{2} c_{\mathrm{W}}^{2}},  \tag{22}\\
& U=16 \pi\left\{\frac{\Pi_{11}\left(m_{\mathrm{Z}}^{2}\right)-\Pi_{11}(0)}{m_{\mathrm{Z}}^{2}}-\frac{\Pi_{33}\left(m_{\mathrm{Z}}^{2}\right)-\Pi_{33}(0)}{m_{\mathrm{Z}}^{2}}\right\}
\end{align*}
$$

where $s_{\mathrm{W}}=\sin \theta_{\mathrm{W}}$ and $c_{\mathrm{W}}=\cos \theta_{\mathrm{W}}$ with the Weinberg angle $\theta_{\mathrm{W}}$ defined at the energy scale $\mu=m_{\mathrm{Z}}$. In the above definitions, $\Pi_{11}$ and $\Pi_{33}$ are the vacuum polarizations of isospin currents, and $\Pi_{3 Q}$ is the vacuum polarization of one isospin and one hypercharge current.

By comparing the measurable electroweak observables with the theoretical predictions, one finds the fitted values [12]

$$
\begin{align*}
& \Delta S=S-S_{\mathrm{SM}}=0.04 \pm 0.10 \\
& \Delta T=T-T_{\mathrm{SM}}=0.05 \pm 0.11  \tag{23}\\
& \Delta U=U-U_{\mathrm{SM}}=0.08 \pm 0.11
\end{align*}
$$

As mentioned above, there is no mass mixing between the exotic quarks and the SM quarks. The corresponding corrections to the oblique parameters from exotic quarks are

$$
\begin{align*}
\Delta S_{\mathrm{Q}^{\prime}}= & \frac{1}{\pi}\left\{\int_{0}^{1} \mathrm{~d} x x(1-x) \ln \frac{m_{\mathrm{b}^{\prime}}^{2}-x(1-x) m_{\mathrm{Z}}^{2}}{m_{t^{\prime}}^{2}-x(1-x) m_{\mathrm{Z}}^{2}}-\frac{3 m_{\mathrm{t}^{\prime}}^{2}}{2 m_{\mathrm{Z}}^{2}} \int_{0}^{1} \mathrm{~d} x \ln \frac{m_{\mathrm{t}^{\prime}}^{2}-x(1-x) m_{\mathrm{Z}}^{2}}{m_{\mathrm{t}^{\prime}}^{2}}-\frac{3 m_{\mathrm{b}^{\prime}}^{2}}{2 m_{\mathrm{Z}}^{2}} \int_{0}^{1} \mathrm{~d} x \ln \frac{m_{\mathrm{b}^{\prime}}^{2}-x(1-x) m_{\mathrm{Z}}^{2}}{m_{\mathrm{b}^{\prime}}^{2}}\right\} \\
\Delta T_{\mathrm{Q}^{\prime}}= & -\frac{3}{4 \pi s_{\mathrm{W}}^{2} c_{\mathrm{W}}^{2}}\left\{\frac{m_{\mathrm{t}^{\prime}}^{2}}{m_{\mathrm{Z}}^{2}} \int_{0}^{1} \mathrm{~d} x x \ln \frac{x m_{\mathrm{t}^{\prime}}^{2}+(1-x) m_{\mathrm{b}^{\prime}}^{2}}{m_{\mathrm{t}^{\prime}}^{2}}+\frac{m_{\mathrm{b}^{\prime}}^{2}}{m_{\mathrm{Z}}^{2}} \int_{0}^{1} \mathrm{~d} x(1-x) \ln \frac{x m_{\mathrm{t}^{\prime}}^{2}+(1-x) m_{\mathrm{b}^{\prime}}^{2}}{m_{\mathrm{b}^{\prime}}^{2}}\right\} \\
\Delta U_{\mathrm{Q}^{\prime}}= & \frac{1}{\pi}\left\{3 \int_{0}^{1} \mathrm{~d} x x(1-x) \ln \frac{\left[x m_{\mathrm{t}^{\prime}}^{2}+(1-x) m_{\mathrm{b}^{\prime}}^{2}-x(1-x) m_{\mathrm{Z}}^{2}\right]^{2}}{\left[m_{\mathrm{t}^{\prime}}^{2}-x(1-x) m_{\mathrm{Z}}^{2}\right]\left[m_{\mathrm{b}^{\prime}}^{2}-x(1-x) m_{\mathrm{Z}}^{2}\right]}\right. \\
& -3 \int_{0}^{1} \mathrm{~d} x\left(x \frac{m_{\mathrm{t}^{\prime}}^{2}}{m_{\mathrm{Z}}^{2}}+(1-x) \frac{m_{\mathrm{b}^{\prime}}^{2}}{m_{\mathrm{Z}}^{2}}\right) \ln \frac{x m_{\mathrm{t}^{\prime}}^{2}+(1-x) m_{\mathrm{b}^{\prime}}^{2}-x(1-x) m_{\mathrm{Z}}^{2}}{x m_{\mathrm{t}^{\prime}}^{2}+(1-x) m_{\mathrm{b}^{\prime}}^{2}} \\
& \left.-\frac{3 m_{\mathrm{t}^{\prime}}^{2}}{2 m_{\mathrm{Z}}^{2}} \int_{0}^{1} \mathrm{~d} x \ln \frac{m_{\mathrm{t}^{\prime}}^{2}-x(1-x) m_{\mathrm{Z}}^{2}}{m_{\mathrm{t}^{\prime}}^{2}}-\frac{3 m_{\mathrm{b}^{\prime}}^{2}}{2 m_{\mathrm{Z}}^{2}} \int_{0}^{1} \mathrm{~d} x \ln \frac{m_{\mathrm{b}^{\prime}}^{2}-x(1-x) m_{\mathrm{Z}}^{2}}{m_{\mathrm{b}^{\prime}}^{2}}\right\} . \tag{24}
\end{align*}
$$

Here, $m_{\mathrm{b}^{\prime}}$ and $m_{\mathrm{t}^{\prime}}$ denote the masses of the charged $-1 / 3 \quad$ the exotic charged leptons and the SM leptons. Ignoring exotic quark $b^{\prime}$ and the charged $2 / 3$ exotic quark $\mathrm{t}^{\prime}$, respectively.

In a similar way, there is no mass mixing between
the tiny mixing between the left-handed neutrinos and heavy majorana neutrinos, we write the corrections to the oblique parameters from exotic leptons as

$$
\begin{align*}
& \Delta S_{\mathrm{L}^{\prime}}=\frac{1}{\pi} \sum_{i, j=1}^{5}\left(Z_{\mathrm{N}}\right)_{1 i}\left(Z_{\mathrm{N}}^{\dagger}\right)_{i 1}\left(Z_{\mathrm{N}}\right)_{1 j}\left(Z_{\mathrm{N}}^{\dagger}\right)_{j 1}\left\{-\frac{1}{2} \int_{0}^{1} \mathrm{~d} x\left(x \frac{m_{\mathrm{N}_{\mathrm{i}}}^{2}}{m_{\mathrm{z}}^{2}}+(1-x) \frac{m_{\mathrm{N}_{j}}^{2}}{m_{\mathrm{Z}}^{2}}\right)\right. \\
& \times \ln \frac{x m_{\mathrm{N}_{i}}^{2}+(1-x) m_{\mathrm{N}_{j}}^{2}-x(1-x) m_{\mathrm{Z}}^{2}}{x m_{\mathrm{N}_{i}}^{2}+(1-x) m_{\mathrm{N}_{j}}^{2}}-\frac{m_{\tau^{\prime}}^{2}}{2 m_{\mathrm{Z}}^{2}} \int_{0}^{1} \mathrm{~d} x \ln \frac{m_{\tau^{\prime}}^{2}-x(1-x) m_{\mathrm{Z}}^{2}}{m_{\tau^{\prime}}^{2}} \\
& \left.+\int_{0}^{1} \mathrm{~d} x x(1-x) \ln \frac{x m_{\mathrm{N}_{i}}^{2}+(1-x) m_{\mathrm{N}_{j}}^{2}-x(1-x) m_{\mathrm{Z}}^{2}}{m_{\tau^{\prime}}^{2}-x(1-x) m_{\mathrm{Z}}^{2}}\right\}, \\
& \Delta T_{\mathrm{L}^{\prime}}=-\frac{1}{4 \pi s_{\mathrm{W}}^{2} c_{\mathrm{W}}^{2}} \sum_{i, j=1}^{5}\left(Z_{\mathrm{N}}\right)_{1 i}\left(Z_{\mathrm{N}}^{\dagger}\right)_{i 1}\left(Z_{\mathrm{N}}\right)_{1 j}\left(Z_{\mathrm{N}}^{\dagger}\right)_{j 1}\left\{\frac{m_{\mathrm{N}_{i}}^{2}}{m_{\mathrm{Z}}^{2}} \int_{0}^{1} \mathrm{~d} x \ln \frac{x m_{\mathrm{N}_{i}}^{2}+(1-x) m_{\tau^{\prime}}^{2}}{x m_{\mathrm{N}_{i}}^{2}+(1-x) m_{\mathrm{N}_{j}}^{2}}\right. \\
& \left.+\frac{m_{\tau^{\prime}}^{2}}{m_{\mathrm{Z}}^{2}} \int_{0}^{1} \mathrm{~d} x(1-x) \ln \frac{x m_{\mathrm{N}_{i}}^{2}+(1-x) m_{\tau^{\prime}}^{2}}{m_{\tau^{\prime}}^{2}}\right\}, \\
& \Delta U_{\mathrm{L}^{\prime}}=\frac{1}{\pi} \sum_{i, j=1}^{5}\left(Z_{\mathrm{N}}\right)_{1 i}\left(Z_{\mathrm{N}}^{\dagger}\right)_{i 1}\left(Z_{\mathrm{N}}\right)_{1 j}\left(Z_{\mathrm{N}}^{\dagger}\right)_{j 1}\left\{\frac{m_{\tau^{\prime}}^{2}}{2 m_{\mathrm{Z}}^{2}} \int_{0}^{1} \mathrm{~d} x \ln \frac{m_{\tau^{\prime}}^{2}-x(1-x) m_{\mathrm{Z}}^{2}}{m_{\tau^{\prime}}^{2}}\right. \\
& +\int_{0}^{1} \mathrm{~d} x x(1-x) \ln \frac{\left[x m_{\mathrm{N}_{i}}^{2}+(1-x) m_{\tau^{\prime}}^{2}-x(1-x) m_{\mathrm{Z}}^{2}\right]^{2}}{\left(x m_{\mathrm{N}_{i}}^{2}+(1-x) m_{\mathrm{N}_{j}}^{2}-x(1-x) m_{\mathrm{Z}}^{2}\right)\left(m_{\tau^{\prime}}^{2}-x(1-x) m_{\mathrm{Z}}^{2}\right)} \\
& -\int_{0}^{1} \mathrm{~d} x\left(x \frac{m_{\mathrm{N}_{i}}^{2}}{m_{\mathrm{Z}}^{2}}+(1-x) \frac{m_{\tau^{\prime}}^{2}}{m_{\mathrm{Z}}^{2}}\right) \ln \frac{x m_{\mathrm{N}_{i}}^{2}+(1-x) m_{\tau^{\prime}}^{2}-x(1-x) m_{\mathrm{Z}}^{2}}{x m_{\mathrm{N}_{i}}^{2}+(1-x) m_{\tau^{\prime}}^{2}} \\
& \left.+\frac{1}{2} \int_{0}^{1} \mathrm{~d} x\left(x \frac{m_{\mathrm{N}_{i}}^{2}}{m_{\mathrm{Z}}^{2}}+(1-x) \frac{m_{\mathrm{N}_{j}}^{2}}{m_{\mathrm{Z}}^{2}}\right) \ln \frac{x m_{\mathrm{N}_{i}}^{2}+(1-x) m_{\mathrm{N}_{j}}^{2}-x(1-x) m_{\mathrm{Z}}^{2}}{x m_{\mathrm{N}_{i}}^{2}+(1-x) m_{\mathrm{N}_{j}}^{2}}\right\} . \tag{25}
\end{align*}
$$

Here, the $5 \times 5$ unitary matrix $Z_{\mathrm{N}}$ is the mixing matrix for heavy majorana neutrinos, $m_{\mathrm{N}_{i}}(i=1,2, \cdots, 5)$ are the corresponding masses of the heavy neutrinos, and $m_{\tau^{\prime}}$ is the mass of the charged exotic lepton $\tau^{\prime}$.

Since the radiative corrections to the self energy of gauge bosons originate from three $C P$-even $\operatorname{Higgs}\left(\mathrm{h}_{0}, \mathrm{H}_{2}^{0}\right.$, $\mathrm{H}_{3}^{0}$ ), the corresponding contributions to the oblique parameters are given by

$$
\begin{aligned}
\Delta S_{\mathrm{H}}= & \frac{1}{\pi} \sum_{i=1}^{3}\left(Z_{\mathrm{CPE}}\right)_{1 i}^{2}\left\{\frac{1}{2} \int_{0}^{1} \mathrm{~d} x x(1-x) \ln \frac{x^{2} m_{\mathrm{Z}}^{2}+(1-x) m_{\mathrm{H}_{i}^{0}}^{2}}{m_{\mathrm{Z}}^{2}}+\int_{0}^{1} \mathrm{~d} x\left(1-\frac{x}{2}-(1-x) \frac{m_{\mathrm{H}_{i}^{0}}^{2}}{2 m_{\mathrm{Z}}^{2}}\right) \ln \frac{x^{2} m_{\mathrm{Z}}^{2}+(1-x) m_{\mathrm{H}_{i}^{0}}^{2}}{x m_{\mathrm{Z}}^{2}+(1-x) m_{\mathrm{H}_{i}^{0}}^{2}}\right\} \\
\Delta T_{\mathrm{H}}= & \frac{1}{4 \pi s_{\mathrm{W}}^{2} c_{\mathrm{W}}^{2}} \sum_{i=1}^{3}\left(Z_{\mathrm{CPE}}\right)_{1 i}^{2}\left\{-\int_{0}^{1} \mathrm{~d} x\left(1-\frac{x}{2}-(1-x) \frac{m_{\mathrm{H}_{i}^{0}}^{2}}{2 m_{\mathrm{Z}}^{2}}\right) \ln \frac{x m_{\mathrm{Z}}^{2}+(1-x) m_{\mathrm{H}_{i}^{0}}^{2}}{m_{\mathrm{Z}}^{2}}\right. \\
& \left.+\int_{0}^{1} \mathrm{~d} x\left(\left(1-\frac{x}{2}\right) c_{\mathrm{W}}^{2}-(1-x) \frac{m_{\mathrm{H}_{i}^{0}}^{2}}{2 m_{\mathrm{Z}}^{2}}\right) \ln \frac{x m_{\mathrm{W}}^{2}+(1-x) m_{\mathrm{H}_{i}^{0}}^{2}}{m_{\mathrm{Z}}^{2}}\right\}, \\
\Delta U_{\mathrm{H}}= & \frac{1}{\pi} \sum_{i=1}^{3}\left(Z_{\mathrm{CPE}}\right)_{1 i}^{2}\left\{\frac{1}{2} \int_{0}^{1} \mathrm{~d} x\left(x^{2}+(1-x) \frac{m_{\mathrm{H}_{i}^{0}}^{2}}{m_{\mathrm{Z}}^{2}}\right) \ln \frac{x^{2} m_{\mathrm{Z}}^{2}+(1-x) m_{\mathrm{H}_{i}^{0}}^{2}}{m_{\mathrm{Z}}^{2}}-\int_{0}^{1} \mathrm{~d} x \ln \frac{x^{2} m_{\mathrm{Z}}^{2}+(1-x) m_{\mathrm{H}_{i}^{0}}^{2}}{x m_{\mathrm{Z}}^{2}+(1-x) m_{\mathrm{H}_{i}^{0}}^{2}}\right. \\
& -\frac{1}{2} \int_{0}^{1} \mathrm{~d} x\left(x+(1-x) \frac{m_{\mathrm{H}_{i}^{0}}^{2}}{m_{\mathrm{Z}}^{2}}\right) \ln \frac{x m_{\mathrm{Z}}^{2}+(1-x) m_{\mathrm{H}_{i}^{0}}^{2}}{m_{\mathrm{Z}}^{2}}+\int_{0}^{1} \mathrm{~d} x\left(\left(1-\frac{x}{2}\right) c_{\mathrm{W}}^{2}-(1-x) \frac{m_{\mathrm{H}_{i}^{0}}^{2}}{2 m_{\mathrm{Z}}^{2}}\right) \ln \frac{x m_{\mathrm{W}}^{2}+(1-x) m_{\mathrm{H}_{i}^{0}}^{2}-x(1-x) m_{\mathrm{Z}}^{2}}{x m_{\mathrm{W}}^{2}+(1-x) m_{\mathrm{H}_{i}^{0}}^{2}}
\end{aligned}
$$

$$
\begin{equation*}
\left.+\frac{1}{2} \int_{0}^{1} \mathrm{~d} x x(1-x) \ln \frac{x m_{\mathrm{W}}^{2}+(1-x) m_{\mathrm{H}_{i}^{0}}^{2}-x(1-x) m_{\mathrm{Z}}^{2}}{m_{\mathrm{Z}}^{2}}\right\} . \tag{26}
\end{equation*}
$$

Here, we adopt the notation $H_{1}^{0}$ to represent the lightest neutral Higgs $h_{0}$. In addition, the contributions from $\Phi_{i}^{0}$ and $\phi^{ \pm}$to the oblique parameters are formulated as follows

$$
\begin{align*}
\Delta S_{\phi}= & \frac{1}{2 \pi}\left\{\sum _ { i , j } ^ { 4 } ( Z _ { \mathrm { CPM } } ) _ { 1 _ { i } } ^ { 2 } ( Z _ { \mathrm { CPM } } ) _ { 3 j } ^ { 2 } \left[-\int_{0}^{1} \mathrm{~d} x\left(x \frac{m_{\Phi_{i}^{0}}^{2}}{m_{\mathrm{Z}}^{2}}+(1-x) \frac{m_{\Phi_{j}^{0}}^{2}}{m_{\mathrm{Z}}^{2}}\right) \ln \frac{x m_{\Phi_{i}^{0}}^{2}+(1-x) m_{\Phi_{j}^{0}}^{2}-x(1-x) m_{\mathrm{Z}}^{2}}{x m_{\Phi_{i}^{0}}^{2}+(1-x) m_{\Phi_{j}^{0}}^{2}}\right.\right. \\
& \left.\left.-\int_{0}^{1} \mathrm{~d} x x(1-x) \ln \frac{x m_{\Phi_{i}^{0}}^{2}+(1-x) m_{\Phi_{j}^{0}}^{2}-x(1-x) m_{\mathrm{Z}}^{2}}{m_{\Phi^{ \pm}}^{2}-x(1-x) m_{\mathrm{Z}}^{2}}\right]+\frac{m_{\phi^{ \pm}}^{2}}{m_{\mathrm{Z}}^{2}} \int_{0}^{1} \mathrm{~d} x \ln \frac{m_{\Phi^{ \pm}}^{2}-x(1-x) m_{\mathrm{Z}}^{2}}{m_{\Phi^{ \pm}}^{2}}\right\}, \\
\Delta T_{\Phi}= & \frac{1}{8 \pi s_{\mathrm{W}}^{2} c_{\mathrm{W}}^{2}}\left\{-\sum_{i=1}^{4}\left[\left(Z_{\mathrm{CPM}}\right)_{1 i}^{2}+\left(Z_{\mathrm{CPM}}\right)_{3 i}^{2}\right]\left[\frac{m_{\phi^{ \pm}}^{2}}{m_{\mathrm{Z}}^{2}} \int_{0}^{1} \mathrm{~d} x x \ln \frac{x m_{\phi^{ \pm}}^{2}+(1-x) m_{\Phi_{i}^{0}}^{2}}{m_{\phi^{ \pm}}^{2}}+\frac{m_{\Phi_{i}^{0}}^{2}}{m_{\mathrm{Z}}^{2}} \int_{0}^{1} \mathrm{~d} x x \ln \frac{x m_{\Phi_{i}^{0}}^{2}+(1-x) m_{\phi^{ \pm}}^{2}}{m_{\Phi_{i}^{0}}^{2}}\right]\right. \\
& \left.+\sum_{i, j}^{4}\left(Z_{\mathrm{CPM}}\right)_{1 i}^{2}\left(Z_{\mathrm{CPM}}\right)_{3 j}^{2} \int_{0}^{1} \mathrm{~d} x\left(x \frac{m_{\Phi_{i}^{0}}^{2}}{m_{\mathrm{Z}}^{2}}+(1-x) \frac{m_{\Phi_{j}^{0}}^{2}}{m_{\mathrm{Z}}^{2}}\right) \ln \frac{x m_{\Phi_{i}^{0}}^{2}+(1-x) m_{\Phi_{j}^{0}}^{2}}{m_{\Phi_{i}^{0}}^{2}}\right\}, \\
\Delta U_{\Phi}= & \frac{1}{2 \pi}\left\{-\sum_{i=1}^{4}\left[\left(Z_{\mathrm{CPM}}\right)_{1 i}^{2}+\left(Z_{\mathrm{CPM}}\right)_{3 i}^{2}\right]\left[\int_{0}^{1} \mathrm{~d} x\left(x \frac{m_{\phi^{ \pm}}^{2}}{m_{\mathrm{Z}}^{2}}+(1-x) \frac{m_{\Phi_{i}^{0}}^{2}}{m_{\mathrm{Z}}^{2}}\right) \ln \frac{x m_{\phi^{ \pm}}^{2}+(1-x) m_{\Phi_{i}^{0}}^{2}-x(1-x) m_{\mathrm{Z}}^{2}}{x m_{\phi^{ \pm}}^{2}+(1-x) m_{\Phi_{i}^{0}}^{2}}\right.\right. \\
& \left.+\frac{1}{2} \int_{0}^{1} \mathrm{~d} x x(1-x) \ln \frac{x m_{\phi^{ \pm}}^{2}+(1-x) m_{\Phi_{i}^{0}}^{2}-x(1-x) m_{\mathrm{Z}}^{2}}{m_{\phi^{ \pm}}^{2}-x(1-x) m_{\mathrm{Z}}^{2}}\right]+\sum_{i, j}^{4}\left(Z_{\mathrm{CPM}}\right)_{1 i}^{2}\left(Z_{\mathrm{CPM}^{2}}\right)_{3 j}^{2}\left[\int_{0}^{1} \mathrm{~d} x\left(x \frac{m_{\Phi_{i}^{0}}^{2}}{m_{\mathrm{Z}}^{2}}+(1-x) \frac{m_{\Phi_{j}^{0}}^{2}}{m_{\mathrm{Z}}^{2}}\right)\right. \\
& \times \ln \frac{x m_{\Phi_{i}^{0}}^{2}+(1-x) m_{\Phi_{j}^{0}}^{2}-x(1-x) m_{\mathrm{Z}}^{2}}{x m_{\Phi_{i}^{0}}^{2}+(1-x) m_{\Phi_{j}^{0}}^{2}}+\int_{0}^{1} \mathrm{~d} x x(1-x) \ln \frac{x m_{\Phi_{i}^{0}}^{2}+(1-x) m_{\Phi_{j}^{0}}^{2}-x(1-x) m_{\mathrm{Z}}^{2}}{x m_{\Phi_{i}^{0}}^{2}+(1-x) m_{\phi^{ \pm}}^{2}-x(1-x) m_{\mathrm{Z}}^{2}} \\
& \left.\left.+\int_{0}^{1} \mathrm{~d} x x(1-x) \ln \frac{x m_{\Phi_{i}^{0}}^{2}+(1-x) m_{\Phi_{j}^{0}}^{2}-x(1-x) m_{\mathrm{Z}}^{2}}{x m_{\phi^{ \pm}}^{2}+(1-x) m_{\Phi_{j}^{0}}^{2}-x(1-x) m_{\mathrm{Z}}^{2}}\right]+\frac{m_{\phi^{ \pm}}^{2}}{m_{\mathrm{Z}}^{2}} \int_{0}^{1} \mathrm{~d} x \ln \frac{m_{\phi^{ \pm}}^{2}-x(1-x) m_{\mathrm{Z}}^{2}}{m_{\Phi^{ \pm}}^{2}}\right\} . \tag{27}
\end{align*}
$$

## 5 Numerical analysis

As mentioned above, the most stringent constraint on the parameter space is that the $3 \times 3$ mass squared matrix in Eq. (5) should produce the lightest eigenvector with a mass $m_{\mathrm{h}_{0}}=125 \mathrm{GeV}$.

In order to make the final results consistent with this condition, we require the self coupling of the Higgs doublet to satisfy

$$
\begin{equation*}
\lambda_{\mathrm{HH}}=\frac{A}{B} \tag{28}
\end{equation*}
$$

where

$$
\begin{align*}
A= & m_{\mathrm{h}_{0}}^{6}-2\left[\lambda_{\mathrm{BB}} v_{\mathrm{B}}^{2}+\lambda_{\mathrm{LL}} v_{\mathrm{L}}^{2}\right] m_{\mathrm{h}_{0}}^{4}+\left[\left(4 \lambda_{\mathrm{BB}} \lambda_{\mathrm{LL}}-\lambda_{\mathrm{BL}}^{2}\right) v_{\mathrm{B}}^{2} v_{\mathrm{L}}^{2}\right. \\
& \left.-\lambda_{\mathrm{HB}}^{2} v^{2} v_{\mathrm{B}}^{2}-\lambda_{\mathrm{HL}}^{2} v^{2} v_{\mathrm{L}}^{2}\right] m_{\mathrm{h}_{0}}^{2}+2\left[\lambda_{\mathrm{BB}} \lambda_{\mathrm{HL}}^{2}+\lambda_{\mathrm{LL}} \lambda_{\mathrm{HB}}^{2}\right. \\
& \left.-\lambda_{\mathrm{HB}} \lambda_{\mathrm{HL}} \lambda_{\mathrm{BL}}\right] v^{2} v_{\mathrm{B}}^{2} v_{\mathrm{L}}^{2} \\
B= & 2 v^{2}\left[m_{\mathrm{h}_{0}}^{4}-2\left(\lambda_{\mathrm{BB}} v_{\mathrm{B}}^{2}+\lambda_{\mathrm{LL}} v_{\mathrm{L}}^{2}\right) m_{\mathrm{h}_{0}}^{2}\right. \\
& \left.+\left(4 \lambda_{\mathrm{BB}} \lambda_{\mathrm{LL}}-\lambda_{\mathrm{BL}}^{2}\right) v_{\mathrm{B}}^{2} v_{\mathrm{L}}^{2}\right] . \tag{29}
\end{align*}
$$

The present experimental lower bounds on the fourth generation charged lepton $\tau^{\prime}$, up-type and down-type quarks $\mathrm{t}^{\prime}$ and $\mathrm{b}^{\prime}$ at $95 \%$ C.L. are $m_{\tau^{\prime}}>100.8 \mathrm{GeV}$, $m_{\mathrm{t}^{\prime}}>420 \mathrm{GeV}$ and $m_{\mathrm{b}^{\prime}}>372 \mathrm{GeV}$, respectively. The fourth generation quarks $\mathrm{t}^{\prime}$ and $\mathrm{b}^{\prime}$ acquire nonzero masses $m_{\mathrm{t}^{\prime}}=m_{\mathrm{b}^{\prime}}=\frac{Y_{\mathrm{Q}}^{\prime}}{\sqrt{2}} v_{\mathrm{B}}$ when local $U(1)_{\mathrm{B}}$ symmetry is broken. In addition, the charged leptons of the fourth generation $\tau^{\prime}$ obtain nonzero masses $m_{\tau^{\prime}}=\frac{Y_{\mathrm{E}}^{\prime}}{\sqrt{2}} v_{\mathrm{L}}$ when local $U(1)_{\mathrm{L}}$ symmetry is broken.

However, there are too many free parameters in the model considered here. In our numerical analysis, we adopt the assumption on the parameter space

$$
\begin{align*}
m_{\mathrm{t}^{\prime}} & =m_{\mathrm{b}^{\prime}}=m_{\tau^{\prime}}=m_{\mathrm{F}} \\
\lambda_{\mathrm{BB}} & =\lambda_{\mathrm{LL}}=0.5, \lambda_{\mathrm{HL}}=\lambda_{\mathrm{BL}}=\lambda_{\mathrm{HB}}=\lambda_{\mathrm{NP}}  \tag{30}\\
\lambda_{\phi \mathrm{H}} & =\lambda_{\phi \mathrm{B}}=\lambda_{\phi \mathrm{L}}=\lambda_{\mathrm{NP}}^{\prime}
\end{align*}
$$

to decrease the number of free parameters in the concerned model. Furthermore, we assume $v \ll v_{\mathrm{B}, \mathrm{L}}$,
$\left(\lambda_{\mathrm{a}}\right)_{3 \times 3}=\operatorname{diag}\left(\lambda_{\mathrm{a}}, \lambda_{\mathrm{a}}, \lambda_{\mathrm{a}}\right)$, and choose the hierarchical assumption on Yukawa couplings $\left|\left(Y_{v}\right)_{I J}\right| \ll\left|Y_{v}^{\prime}\right| \sim \lambda_{\mathrm{a}} \sim \lambda_{\mathrm{b}_{I}}$, $(I, J=1,2,3)$ to obtain our final results. Applying the assumptions above, we obtain the majorana mass for the lightest exotic neutrino $\mathrm{N}_{1}$ to be

$$
\begin{equation*}
m_{\mathrm{N}_{1}} \approx \frac{v^{2}}{\sqrt{2} v_{\mathrm{L}}} \frac{\lambda_{\mathrm{a}}\left|Y_{v}^{\prime}\right|}{\lambda_{\mathrm{b}}^{2}} \tag{31}
\end{equation*}
$$

with $\lambda_{\mathrm{b}}^{2}=\lambda_{\mathrm{b}_{1}}^{2}+\lambda_{\mathrm{b}_{2}}^{2}+\lambda_{\mathrm{b}_{3}}^{2}$. Of course, we need this mass
to be greater than $m_{\mathrm{Z}} / 2$ in order to be consistent with the measured Z-boson decay width. The masses of other heavy majorana neutrinos are

$$
\begin{equation*}
m_{\mathrm{N}_{i}} \approx\left(\frac{v_{\mathrm{L}}}{\sqrt{2}} \lambda_{\mathrm{a}}, \frac{v_{\mathrm{L}}}{\sqrt{2}} \lambda_{\mathrm{a}}, \frac{v_{\mathrm{L}}}{2 \sqrt{2}}\left(\Delta-\lambda_{\mathrm{a}}\right), \frac{v_{\mathrm{L}}}{2 \sqrt{2}}\left(\Delta+\lambda_{\mathrm{a}}\right)\right), \tag{32}
\end{equation*}
$$

for $i=2,3,4,5$ and $\Delta=\sqrt{\lambda_{\mathrm{a}}^{2}+4 \lambda_{\mathrm{b}}^{2}}$.
Correspondingly, the $5 \times 5$ mixing matrix $Z_{\mathrm{N}}$ is approximated as

$$
Z_{\mathrm{N}} \approx\left(\begin{array}{ccccc}
1, & \frac{\lambda_{\mathrm{a}} Y_{v}^{*} v}{\lambda_{\mathrm{b}}^{2} v_{\mathrm{L}}}-, & \frac{\lambda_{\mathrm{b}_{1}} Y_{v}^{*} v}{\lambda_{\mathrm{b}}^{2} v_{\mathrm{L}}}, & \frac{\lambda_{\mathrm{b}_{2}} Y_{v}^{*} v}{\lambda_{\mathrm{b}}^{2} v_{\mathrm{L}}}, & \frac{\lambda_{\mathrm{b}_{3}} Y_{v}^{*} v}{\lambda_{\mathrm{b}}^{2} v_{\mathrm{L}}}  \tag{33}\\
\frac{\lambda_{\mathrm{a}} Y_{v}^{*} v}{\lambda_{\mathrm{b}}^{2} v_{\mathrm{L}}}, & 0, & 0, & -\mathrm{i} \sqrt{\frac{\Delta+\lambda_{\mathrm{a}}}{2 \Delta}}, & \sqrt{\frac{\Delta-\lambda_{\mathrm{a}}}{2 \Delta}} \\
-\frac{\lambda_{\mathrm{b}_{1}} Y_{v}^{*} v}{\lambda_{\mathrm{b}}^{2} v_{\mathrm{L}}}, & -\frac{\lambda_{\mathrm{b}_{3}}}{\sqrt{\lambda_{\mathrm{b}_{1}}^{2}+\lambda_{\mathrm{b}_{3}}^{2}}}, & -\frac{\lambda_{\mathrm{b}_{2}}}{\sqrt{\lambda_{\mathrm{b}_{1}}^{2}+\lambda_{\mathrm{b}_{2}}^{2}}}, & \frac{\mathrm{i} \sqrt{2} \lambda_{\mathrm{b}_{1}}}{\sqrt{\Delta^{2}+\lambda_{\mathrm{a}} \Delta}}, & \frac{\sqrt{2} \lambda_{\mathrm{b}_{1}}}{\sqrt{\Delta^{2}-\lambda_{\mathrm{a}} \Delta}} \\
-\frac{\lambda_{\mathrm{b}_{2}} Y_{v}^{*} v}{\lambda_{\mathrm{b}}^{2} v_{\mathrm{L}}}, & 0, & 0, & \frac{\mathrm{i} \sqrt{2} \lambda_{\mathrm{b}_{2}}}{\sqrt{\Delta^{2}+\lambda_{\mathrm{a}} \Delta}}, & \frac{\sqrt{2} \lambda_{\mathrm{b}_{2}}}{\sqrt{\Delta^{2}-\lambda_{\mathrm{a}} \Delta}} \\
-\frac{\lambda_{\mathrm{b}_{3}} Y_{v}^{*} v}{\lambda_{\mathrm{b}}^{2} v_{\mathrm{L}}}, & \frac{\lambda_{\mathrm{b}_{1}}}{\sqrt{\lambda_{\mathrm{b}_{1}}^{2}+\lambda_{\mathrm{b}_{3}}^{2}}}, & \frac{\lambda_{\mathrm{b}_{1}}}{\sqrt{\lambda_{\mathrm{b}_{1}}^{2}+\lambda_{\mathrm{b}_{2}}^{2}}}, & \frac{\mathrm{i} \sqrt{2} \lambda_{\mathrm{b}_{3}}}{\sqrt{\Delta^{2}+\lambda_{\mathrm{a}} \Delta}}, & \frac{\sqrt{2} \lambda_{\mathrm{b}_{3}}}{\sqrt{\Delta^{2}-\lambda_{\mathrm{a}} \Delta}}
\end{array}\right) .
$$

For the relevant parameters in the SM, we take [13]

$$
\begin{align*}
\alpha_{\mathrm{s}}\left(m_{\mathrm{Z}}\right) & =0.118, \alpha\left(m_{\mathrm{z}}\right)=1 / 128 \\
s_{\mathrm{W}}^{2}\left(m_{\mathrm{Z}}\right) & =0.23, m_{\mathrm{t}}=174.2 \mathrm{GeV}, m_{\mathrm{W}}=80.4 \mathrm{GeV} \tag{34}
\end{align*}
$$

### 5.1 Numerical results of $\boldsymbol{R}_{\gamma \gamma}$

Under our assumptions of the parameter space, the theoretical prediction of $R_{\gamma \gamma}$ depends on six parameters in the model: $m_{\mathrm{F}}, m_{\phi^{ \pm}}, \lambda_{\mathrm{NP}}, \lambda_{\mathrm{NP}}^{\prime}, v_{\mathrm{B}}$ and $v_{\mathrm{L}}$. Taking $m_{\phi^{ \pm}}=500 \mathrm{GeV}, \lambda_{\mathrm{NP}}=0.5$ and $\lambda_{\mathrm{NP}}^{\prime}=-0.5$, we plot the variation of $R_{\gamma \gamma}$ with the mass scalar of exotic fermions $M_{\mathrm{F}}$, as shown in Fig. 1. The dotted line corresponds to $v_{\mathrm{B}}=v_{\mathrm{L}}=500 \mathrm{GeV}$, the dashed line corresponds to $v_{\mathrm{B}}=v_{\mathrm{L}}=1000 \mathrm{GeV}$, and the solid line corresponds to $v_{\mathrm{B}}=v_{\mathrm{L}}=1500 \mathrm{GeV}$. In general, the ratio $R_{\gamma \gamma}$ depends very weakly on the mass scale $m_{\mathrm{F}}$, and the value of $R_{\gamma \gamma}$ is about $1.8-1.9$ when $500 \mathrm{GeV} \leqslant v_{\mathrm{B}}=v_{\mathrm{L}} \leqslant 1500 \mathrm{GeV}$.

In Fig. 2(a), we plot the variation of $R_{\gamma \gamma}$ with the VEV $v_{\mathrm{L}}$ when $m_{\phi^{ \pm}}=v_{\mathrm{B}}=500 \mathrm{GeV}, \lambda_{\mathrm{NP}}^{\prime}=-0.5$ and $\lambda_{\mathrm{NP}}=0.5$. The dotted line corresponds to $m_{\mathrm{F}}=500 \mathrm{GeV}$, the dashed line corresponds to $m_{\mathrm{F}}=550 \mathrm{GeV}$, and the solid line corresponds to $m_{\mathrm{F}}=600 \mathrm{GeV}$. The dependence of $R_{\gamma \gamma}$ on $v_{\mathrm{L}}$ is relatively sensitive for $v_{\mathrm{L}} \leqslant 600 \mathrm{GeV}$, and is weak for $v_{\mathrm{L}}>600 \mathrm{GeV}$. Since the dependence of $R_{\gamma \gamma}$ on $m_{\mathrm{F}}$ and $v_{\mathrm{B}}$ is very weak, the three lines almost coincide with each other. In Fig. 2(b), we plot the variation of $R_{\gamma \gamma}$ with the VEV $v_{\mathrm{L}}$ when $m_{\mathrm{F}}=m_{\phi^{ \pm}}=v_{\mathrm{B}}=500 \mathrm{GeV}$, $\lambda_{\mathrm{NP}}^{\prime}=-0.5$. The dotted line corresponds to $\lambda_{\mathrm{NP}}=0.5$, the dashed line corresponds to $\lambda_{\mathrm{NP}}=0$, and the solid line corresponds to $\lambda_{\mathrm{NP}}=-0.5$. Generally, there is a weak dependence of the ratio $R_{\gamma \gamma}$ on $v_{\mathrm{L}}$.


Fig. 1. Variation of $R_{\gamma \gamma}$ with the mass scale of exotic fermions $m_{\mathrm{F}}$ when $m_{\phi^{ \pm}}=500 \mathrm{GeV}, \lambda_{\mathrm{NP}}=0.5$, $\lambda_{\mathrm{NP}}^{\prime}=-0.5$. The dotted line represents $v_{\mathrm{B}}=$ $v_{\mathrm{L}}=500 \mathrm{GeV}$, the dashed line represents $v_{\mathrm{B}}=$ $v_{\mathrm{L}}=1000 \mathrm{GeV}$, and the solid line represents $v_{\mathrm{B}}=$ $v_{\mathrm{L}}=1500 \mathrm{GeV}$.

In Fig. 3(a), we show the variation of $R_{\gamma \gamma}$ with the VEV $v_{\mathrm{B}}$ when $m_{\mathrm{F}}=v_{\mathrm{L}}=500 \mathrm{GeV}, \lambda_{\mathrm{NP}}=0.5$. The dotted line corresponds to $\lambda_{\mathrm{NP}}^{\prime}=0.5$, the dashed line corresponds to $\lambda_{\mathrm{NP}}^{\prime}=0$, and the solid line corresponds to $\lambda_{\mathrm{NP}}^{\prime}=-0.5$. The dependence of $R_{\gamma \gamma}$ on $v_{\mathrm{B}}$ is relatively sensitive for $v_{\mathrm{B}} \leqslant 600 \mathrm{GeV}$, and is weak for $v_{\mathrm{B}}>600 \mathrm{GeV}$. In Fig. 3(b), we show the variation of $R_{\gamma \gamma}$ with $v_{\mathrm{B}}$ when $m_{\mathrm{F}}=v_{\mathrm{L}}=500 \mathrm{GeV}, \lambda_{\mathrm{NP}}^{\prime}=-0.5$ and $\lambda_{\mathrm{NP}}=0.5$. The dotted line corresponds to $m_{\phi^{ \pm}}=1500 \mathrm{GeV}$, the dashed line


Fig. 2. Variation of $R_{\gamma \gamma}$ with the VEV $v_{\mathrm{L}}$ when $m_{\phi^{ \pm}}=v_{\mathrm{B}}=500 \mathrm{GeV}, \lambda_{\mathrm{NP}}^{\prime}=-0.5$. In (a), $\lambda_{\mathrm{NP}}=$ 0.5 , the dotted line corresponds to $m_{\mathrm{F}}=500 \mathrm{GeV}$, the dashed line corresponds to $m_{\mathrm{F}}=550 \mathrm{GeV}$, and the solid line corresponds to $m_{\mathrm{F}}=600 \mathrm{GeV}$. In (b), $m_{\mathrm{F}}=500 \mathrm{GeV}$, the dotted line corresponds to $\lambda_{\mathrm{NP}}=0.5$, the dashed line corresponds to $\lambda_{\mathrm{NP}}=0$, and the solid line corresponds to $\lambda_{\mathrm{NP}}=-0.5$.
corresponds to $m_{\phi^{ \pm}}=1000 \mathrm{GeV}$, and the solid line corresponds to $m_{\phi^{ \pm}}=500 \mathrm{GeV}$. Generally, there is a very weak dependence of the ratio $R_{\gamma \gamma}$ on $v_{\mathrm{B}}$.

Choosing $v_{\mathrm{B}}=v_{\mathrm{L}}=500 \mathrm{GeV}, \lambda_{\mathrm{NP}}^{\prime}=-0.5$, Fig. 4 presents the variation of the ratio $R_{\gamma \gamma}$ with $\lambda_{\mathrm{NP}}$. The dotted line represents $m_{\mathrm{F}}=500 \mathrm{GeV}, m_{\phi^{ \pm}}=$ 1500 GeV , the dashed line represents $m_{\mathrm{F}}=550 \mathrm{GeV}$, $m_{\phi^{ \pm}}=1000 \mathrm{GeV}$, and the solid line represents $m_{\mathrm{F}}=$ $m_{\phi^{ \pm}}=500 \mathrm{GeV}$. As $\Lambda_{\mathrm{NP}}$ increases, $R_{\gamma \gamma}$ changes drastically and can easily coincide with the present experimental data, as $-0.5 \leqslant \lambda_{\mathrm{NP}} \leqslant 1.0$. Choosing $m_{\mathrm{F}}=v_{\mathrm{B}}=$ $v_{\mathrm{L}}=500 \mathrm{GeV}$, and $\lambda_{\mathrm{NP}}=-0.5$, Fig. 5 shows the ratio $R_{\gamma \gamma}$ versus $m_{\phi^{ \pm}}$. The dotted line represents $\lambda_{\mathrm{NP}}^{\prime}=0.5$, the dashed line represents $\lambda_{\mathrm{NP}}^{\prime}=0$, and the solid line represents $\lambda_{\mathrm{NP}}^{\prime}=-0.5$. For $\lambda_{\mathrm{NP}}^{\prime}=0$, there is a slight dependence of $R_{\gamma \gamma}$ on the mass $m_{\phi \pm}$. When $\lambda_{\mathrm{NP}}^{\prime}= \pm 0.5$,


Fig. 3. Variation of $R_{\gamma \gamma}$ with the VEV $v_{\mathrm{B}}$ when $m_{\mathrm{F}}=v_{\mathrm{L}}=500 \mathrm{GeV}, \lambda_{\mathrm{NP}}=0.5$ for: (a) $m_{\phi^{ \pm}}=500 \mathrm{GeV}$, where the dotted line corresponds to $\lambda_{\mathrm{NP}}^{\prime}=0.5$, the dashed line corresponds to $\lambda_{\mathrm{NP}}^{\prime}=0$, and the solid line corresponds to $\lambda_{\mathrm{NP}}^{\prime}=-0.5 ;(\mathrm{b}) \lambda_{\mathrm{NP}}^{\prime}=-0.5$, where the dotted line corresponds to $m_{\phi^{ \pm}}=1500 \mathrm{GeV}$, the dashed line corresponds to $m_{\phi^{ \pm}}=1000 \mathrm{GeV}$, and the solid line corresponds to $m_{\phi^{ \pm}}=500 \mathrm{GeV}$.
$R_{\gamma \gamma}$ decreases steeply as $m_{\phi^{ \pm}}$increases.
In Fig. 6, we plot the variation of the ratio $R_{\gamma \gamma}$ with $\lambda_{\mathrm{NP}}^{\prime}$ when $m_{\mathrm{F}}=v_{\mathrm{B}}=v_{\mathrm{L}}=500 \mathrm{GeV}$ and $\lambda_{\mathrm{NP}}=-0.5$. The dotted line represents $m_{\phi \pm}=1500 \mathrm{GeV}$, the dashed line represents $m_{\phi^{ \pm}}=1000 \mathrm{GeV}$, and the solid line represents $m_{\phi^{ \pm}}=500 \mathrm{GeV}$. The dependence of $R_{\gamma \gamma}$ on $\lambda_{\mathrm{NP}}^{\prime}$ is strong when $m_{\phi^{ \pm}}=500 \mathrm{GeV}$ but weaker for higher values of $m_{\phi \pm}$.

Generally, the ratio $R_{\gamma \gamma}$ depends strongly on the parameters $\lambda_{\mathrm{NP}}, \lambda_{\mathrm{NP}}^{\prime}$ and $m_{\phi^{ \pm}}$, and depends weakly on $v_{\mathrm{B}}$, $v_{\mathrm{L}}$ and $m_{\mathrm{F}}$. These numerical results can be reasonably explained from Eq. (18) and Eq. (19), where $\lambda_{\mathrm{NP}}$ affects theoretical predictions of $R_{\gamma \gamma}$ through the $3 \times 3$ mixing matrix $Z_{\mathrm{CPE}}$, while $\lambda_{\mathrm{NP}}^{\prime}$ and $m_{\phi^{ \pm}}$affect theoretical predictions of $R_{\gamma \gamma}$ through the last term in Eq. (19).


Fig. 4. Variation of $R_{\gamma \gamma}$ with $\lambda_{\mathrm{NP}}$ when $v_{\mathrm{B}}=$ $v_{\mathrm{L}}=500 \mathrm{GeV}, \lambda_{\mathrm{NP}}^{\prime}=-0.5$, where the dotted line represents $m_{\mathrm{F}}=500 \mathrm{GeV}, m_{\phi^{ \pm}}=1500 \mathrm{GeV}$, the dashed line represents $m_{\mathrm{F}}=550 \mathrm{GeV}$, $m_{\phi^{ \pm}}=1000 \mathrm{GeV}$, and the solid line represents $m_{\mathrm{F}}=m_{\phi^{ \pm}}=500 \mathrm{GeV}$.


Fig. 5. Variation of $R_{\gamma \gamma}$ with $m_{\phi^{ \pm}}$when $m_{\mathrm{F}}=v_{\mathrm{B}}=$ $v_{\mathrm{L}}=500 \mathrm{GeV}, \lambda_{\mathrm{NP}}=-0.5$, where the dotted line represents $\lambda_{\mathrm{NP}}^{\prime}=0.5$, the dashed line represents $\lambda_{\mathrm{NP}}^{\prime}=0$, and the solid line represents $\lambda_{\mathrm{NP}}^{\prime}=-0.5$.

The important point is that the parameters $\lambda_{\mathrm{a}}$, $\lambda_{\mathrm{b}_{i}}(i=1,2,3)$ do not affect the theoretical predictions of $R_{\gamma \gamma}$ since there is no correction to the decay widths of $\mathrm{h}_{0} \rightarrow \gamma \gamma$ and $\mathrm{h}_{0} \rightarrow \mathrm{gg}$ from the neutrino sector at one-loop level. Similarly, the parameters $M_{\mathrm{S}}, \mu_{1}, \mu_{2}$ also do not affect theoretical evaluations of $R_{\gamma \gamma}$ because there is no one-loop correction to the decay widths of $\mathrm{h}_{0} \rightarrow \gamma \gamma$ and $\mathrm{h}_{0} \rightarrow \mathrm{gg}$ from virtual $\Phi_{i}^{0}(i=1,2,3,4)$.


Fig. 6. Variation of $R_{\gamma \gamma}$ with $\lambda_{\mathrm{NP}}^{\prime}$ when $m_{\mathrm{F}}=v_{\mathrm{B}}=$ $v_{\mathrm{L}}=500 \mathrm{GeV}, \lambda_{\mathrm{NP}}=-0.5$, where the dotted line represents $m_{\phi^{ \pm}}=1500 \mathrm{GeV}$, the dashed line represents $m_{\phi^{ \pm}}=1000 \mathrm{GeV}$, and the solid line represents $m_{\phi^{ \pm}}=500 \mathrm{GeV}$.


Fig. 7. Adopting the assumptions mentioned in the text and assuming $\theta_{1}=\arg \left(\mu_{1}\right)=\pi, \theta_{2}=\arg \left(\mu_{2}\right)=$ $\pi / 4$, we present the theoretical values for a) $\Delta S$ (solid line), b) $\Delta U$ (dash-dot-dot line), and c) $\Delta T$ (dashed line) versus the mass $m_{\phi^{ \pm}}$.

### 5.2 The constraints on parameter space from oblique corrections

The heavy neutrinos contribute one-loop radiative corrections to the self energies of $\mathrm{ZZ}, \mathrm{W}^{ \pm} \mathrm{W}^{\mp}$ in this model. This results in the theoretical values of the $S$, $T, U$ parameters depending on $\lambda_{\mathrm{a}}, \lambda_{\mathrm{b}_{i}}(i=1,2,3)$ here. Furthermore, the theoretical values of the $S, T, U$ parameters also depend on $m_{\mathrm{S}}, \mu_{1}, \mu_{2}$ through the virtual $\phi^{ \pm}, \Phi_{i}^{0},(i=1,2,3,4)$ radiative corrections to the self energies of $\mathrm{ZZ}, \mathrm{W}^{ \pm} \mathrm{W}^{\mp}$ at one-loop level. So far, fitting
$S, T, U$ within $3 \sigma$ deviation indicates

$$
\begin{align*}
& -0.26 \leqslant \Delta S \leqslant 0.34 \\
& -0.28 \leqslant \Delta T \leqslant 0.38 \\
& -0.25 \leqslant \Delta U \leqslant 0.41 \tag{35}
\end{align*}
$$

In order to obtain theoretical values of $S, T, U$ that satisfy present experimental data, we adopt the following additional assumptions:

$$
\begin{align*}
m_{\mathrm{N}_{1}} & \approx \frac{v^{2}}{\sqrt{2} v_{\mathrm{L}}} \frac{\lambda_{\mathrm{a}}\left|Y_{v}^{\prime}\right|}{\lambda_{\mathrm{b}}^{2}}=m_{\mathrm{F}} \\
\lambda_{\mathrm{b}_{1}} & =\lambda_{b_{2}}=\lambda_{\mathrm{b}_{3}}=\frac{1}{\sqrt{3}} \lambda_{\mathrm{b}} \\
v_{\mathrm{B}} & =v_{\mathrm{L}}=m_{\mathrm{S}}=m_{\mathrm{F}}=500 \mathrm{GeV} \\
\left|\mu_{1}\right| & =20 \mathrm{GeV},\left|\mu_{2}\right|=200 \mathrm{GeV} \\
\lambda_{\mathrm{a}} & =\lambda_{\mathrm{b}}=0.6, \lambda_{\mathrm{BB}}=\lambda_{\mathrm{LL}}=0.5, \lambda_{\mathrm{NP}}=\lambda_{\mathrm{NP}}^{\prime}=0.01 \tag{36}
\end{align*}
$$

Choosing $\theta_{1}=\arg \left(\mu_{1}\right)=\pi, \theta_{2}=\arg \left(\mu_{2}\right)=\pi / 4$, we depict the theoretical values of $\Delta S, \Delta U, \Delta T$ versus the mass of charged scalar $\phi^{ \pm}$in Fig. 7, in which the solid line represents $\Delta S$, the dash-dot-dot line represents $\Delta U$, and the dashed line represents $\Delta T$. For our choices of the relevant parameters, the theoretical value of $\Delta T$ is very sensitive to the mass $m_{\phi^{ \pm}}$, while the theoretical values of $\Delta S$ and $\Delta U$ have a weak dependence on the mass $m_{\phi^{ \pm}}$. When the mass of the charged scalar lies in the range $400 \leqslant m_{\phi^{ \pm}} / \mathrm{GeV} \leqslant 700$, the theoretical predictions of $\Delta S, \Delta T, \Delta U$ simultaneously satisfy the inequalities in Eq. (35). The $C P$ phases $\theta_{1}, \theta_{2}$ also affect the numerical results of $\Delta S, \Delta U, \Delta T$ through the $4 \times 4$ mixing matrix $Z_{\mathrm{CPM}}$. Taking $m_{\phi^{ \pm}}=600 \mathrm{GeV}$ and $\theta_{2}=\pi / 4$, we present the theoretical evaluations on $\Delta S, \Delta U, \Delta T$ versus the $C P$ phase $\theta_{1}$ in Fig. 8. With our assumptions on the parameter space, the theoretical value of $\Delta T$ varies


Fig. 8. Adopting the assumptions mentioned in the text and assuming $m_{\phi^{ \pm}}=600 \mathrm{GeV}, \theta_{2}=\arg \left(\mu_{2}\right)=$ $\pi / 4$, we present the theoretical values of a) $\Delta S$ (solid line), b) $\Delta U$ (dash-dot-dot line), and c) $\Delta T$ (dashed line) versus the $C P$ phase $\theta_{1}=\arg \left(\mu_{1}\right)$.
strongly with the $C P$ phase $\theta_{1}$, while the theoretical values of $\Delta S$ and $\Delta U$ vary weakly with the $C P$ phase $\theta_{1}$. In the neighbourhoods of $\theta_{1}=0, \pm \pi / 2, \pm \pi$, the theoretical predictions on $\Delta S, \Delta T, \Delta U$ simultaneously lie within the ranges presented in Eq. (35).

In Fig. 9, we present the theoretical values of $\Delta S, \Delta T, \Delta U$ varying with the $C P$ phase $\theta_{2}$ when $m_{\phi^{ \pm}}=600 \mathrm{GeV}$ and $\theta_{1}=\pi$. As the $C P$ phase $\theta_{2}$ varies, the theoretical value of $\Delta T$ changes drastically, while the theoretical values of $\Delta S$ and $\Delta U$ change slowly. In the neighbourhoods around $\theta_{2}= \pm \pi / 4, \pm 3 \pi / 4$, the theoretical predictions of $\Delta S, \Delta T, \Delta U$ coincide with the present global EWPD fit within $3 \sigma$ deviations.


Fig. 9. Adopting the mentioned assumptions in text and assuming $m_{\phi^{ \pm}}=600 \mathrm{GeV}, \theta_{1}=\arg \left(\mu_{1}\right)=$ $\pi$, we present the theoretical values of a) $\Delta S$ (solid line), b) $\Delta U$ (dash-dot-dot line), and c) $\Delta T$ (dashed line) versus the $C P$ phase $\theta_{2}=\arg \left(\mu_{2}\right)$.

## 6 Summary

For an extension of the SM with local gauged baryon and lepton numbers, we have discussed the constraints from the oblique parameters $S, T, U$ when the lightest Higgs has a mass around 125 GeV . Considering these constraints, we find that there is parameter space to account for the excess in Higgs production and decay in the diphoton channel observed in the ATLAS and CMS experiments. Of course, our numerical results strongly depend on the assumptions made in the model considered here. In other words, our theoretical prediction cannot be precise because of the theoretical uncertainties. The purpose of our calculation is to show that this extension of the SM may still be right even after the constraints from LHC data on the Higgs and oblique parameters have been taken into account.

## Appendix A

## Higgs masses and relevant couplings

After diagonalizing the mass matrix Eq. (5), we obtain

$$
\begin{align*}
m_{\mathrm{h}_{0}}^{2} & =\operatorname{Min}\left(m_{1}^{2}, m_{2}^{2}, m_{3}^{2}\right) \\
m_{\mathrm{H}_{3}^{0}}^{2} & =\operatorname{Max}\left(m_{1}^{2}, m_{2}^{2}, m_{3}^{2}\right) \tag{A1}
\end{align*}
$$

with

$$
\begin{align*}
m_{1}^{2} & =-\frac{a}{3}+\frac{2}{3} p \cos \phi \\
m_{2}^{2} & =-\frac{a}{3}-\frac{1}{3} p(\cos \phi-\sqrt{3} \sin \phi)  \tag{A2}\\
m_{3}^{2} & =-\frac{a}{3}-\frac{1}{3} p(\cos \phi+\sqrt{3} \sin \phi)
\end{align*}
$$

To formulate the expressions in a concise form, we define the notations

$$
\begin{equation*}
p=\sqrt{a^{2}-3 b}, \phi=\frac{1}{3} \arccos \left(-\frac{1}{p^{3}}\left(a^{3}-\frac{9}{2} a b+\frac{27}{2} c\right)\right) \tag{A3}
\end{equation*}
$$

where

$$
\begin{align*}
a= & -2\left(\lambda_{\mathrm{HH}} v^{2}+\lambda_{\mathrm{BB}} v_{\mathrm{B}}^{2}+\lambda_{\mathrm{LL}} v_{\mathrm{L}}^{2}\right), \\
b= & 4\left(\lambda_{\mathrm{HH}} \lambda_{\mathrm{BB}} v^{2} v_{\mathrm{B}}^{2}+\lambda_{\mathrm{HH}} \lambda_{\mathrm{LL}} v^{2} v_{\mathrm{L}}^{2}+\lambda_{\mathrm{BB}} \lambda_{\mathrm{LL}} v_{\mathrm{B}}^{2} v_{\mathrm{L}}^{2}\right) \\
& -\lambda_{\mathrm{HB}}^{2} v^{2} v_{\mathrm{B}}^{2}-\lambda_{\mathrm{HL}}^{2} v^{2} v_{\mathrm{L}}^{2}-\lambda_{\mathrm{BL}}^{2} v_{\mathrm{B}}^{2} v_{\mathrm{L}}^{2} \\
c= & 2\left(\lambda_{\mathrm{HH}} \lambda_{\mathrm{BL}}^{2}+\lambda_{\mathrm{BB}} \lambda_{\mathrm{HL}}^{2}+\lambda_{\mathrm{LL}} \lambda_{\mathrm{HB}}^{2}-4 \lambda_{\mathrm{HH}} \lambda_{\mathrm{BB}} \lambda_{\mathrm{LL}}\right. \\
& \left.-\lambda_{\mathrm{HB}} \lambda_{\mathrm{HL}} \lambda_{\mathrm{BL}}\right) v^{2} v_{\mathrm{B}}^{2} v_{\mathrm{L}}^{2} . \tag{A4}
\end{align*}
$$

The normalized eigenvectors of the mass squared matrix in Eq. (5) are given by

$$
\left(\begin{array}{c}
\left(Z_{\mathrm{CPE}}\right)_{11} \\
\left(Z_{\mathrm{CPE}}\right)_{21} \\
\left(Z_{\mathrm{CPE}}\right)_{31}
\end{array}\right)=\frac{1}{\sqrt{\left|X_{1}\right|^{2}+\left|Y_{1}\right|^{2}+\left|Z_{1}\right|^{2}}}\left(\begin{array}{c}
X_{1} \\
Y_{1} \\
Z_{1}
\end{array}\right)
$$

$$
\begin{align*}
& \left(\begin{array}{l}
\left(Z_{\mathrm{CPE}}\right)_{12} \\
\left(Z_{\mathrm{CPE}}\right)_{22} \\
\left(Z_{\mathrm{CPE}}\right)_{32}
\end{array}\right)=\frac{1}{\sqrt{\left|X_{2}\right|^{2}+\left|Y_{2}\right|^{2}+\left|Z_{2}\right|^{2}}}\left(\begin{array}{c}
X_{2} \\
Y_{2} \\
Z_{2}
\end{array}\right), \\
& \left(\begin{array}{c}
\left(Z_{\mathrm{CPE}}\right)_{13} \\
\left(Z_{\mathrm{CPE}}\right)_{23} \\
\left(Z_{\mathrm{CPE}}\right)_{33}
\end{array}\right)=\frac{1}{\sqrt{\left|X_{3}\right|^{2}+\left|Y_{3}\right|^{2}+\left|Z_{3}\right|^{2}}}\left(\begin{array}{c}
X_{3} \\
Y_{3} \\
Z_{3}
\end{array}\right), \tag{A5}
\end{align*}
$$

with

$$
\begin{align*}
X_{1} & =\left(2 \lambda_{\mathrm{BB}} v_{\mathrm{B}}^{2}-m_{1}^{2}\right)\left(2 \lambda_{\mathrm{LL}} v_{\mathrm{L}}^{2}-m_{1}^{2}\right)-\lambda_{\mathrm{BL}}^{2} v_{\mathrm{B}}^{2} v_{\mathrm{L}}^{2}, \\
Y_{1} & =\lambda_{\mathrm{HL}} \lambda_{\mathrm{BL}} v v_{\mathrm{B}} v_{\mathrm{L}}^{2}-\lambda_{\mathrm{HB}} v v_{\mathrm{B}}\left(2 \lambda_{\mathrm{LL}} v_{\mathrm{L}}^{2}-m_{1}^{2}\right), \\
Z_{1} & =\lambda_{\mathrm{HB}} \lambda_{\mathrm{BL}} v v_{\mathrm{B}}^{2} v_{\mathrm{L}}-\lambda_{\mathrm{HL}} v v_{\mathrm{L}}\left(2 \lambda_{\mathrm{BB}} v_{\mathrm{B}}^{2}-m_{1}^{2}\right), \\
X_{2} & =\lambda_{\mathrm{HL}} \lambda_{\mathrm{BL}} v v_{\mathrm{B}} v_{\mathrm{L}}^{2}-\lambda_{\mathrm{HB}} v v_{\mathrm{B}}\left(2 \lambda_{\mathrm{LL}} v_{\mathrm{L}}^{2}-m_{2}^{2}\right), \\
Y_{2} & =\left(2 \lambda_{\mathrm{HH}}^{2} v^{2}-m_{2}^{2}\right)\left(2 \lambda_{\mathrm{LL}} v_{\mathrm{L}}^{2}-m_{2}^{2}\right)-\lambda_{\mathrm{HL}}^{2} v^{2} v_{\mathrm{L}}^{2}, \\
Z_{2} & =\lambda_{\mathrm{HB}} \lambda_{\mathrm{HL}} v^{2} v_{\mathrm{B}} v_{\mathrm{L}}-\lambda_{\mathrm{BL}} v_{\mathrm{B}} v_{\mathrm{L}}\left(2 \lambda_{\mathrm{HH}} v^{2}-m_{2}^{2}\right), \\
X_{3} & =\lambda_{\mathrm{HB}} \lambda_{\mathrm{BL}} v v_{\mathrm{B}}^{2} v_{\mathrm{L}}-\lambda_{\mathrm{HL}} v v_{\mathrm{L}}\left(2 \lambda_{\mathrm{BB}} v_{\mathrm{B}}^{2}-m_{3}^{2}\right), \\
Y_{3} & =\lambda_{\mathrm{HB}} \lambda_{\mathrm{HL}} v^{2} v_{\mathrm{B}} v_{\mathrm{L}}-\lambda_{\mathrm{BL}} v_{\mathrm{B}} v_{\mathrm{L}}\left(2 \lambda_{\mathrm{HH}}^{2} v^{2}-m_{3}^{2}\right), \\
Z_{3} & =\left(2 \lambda_{\mathrm{HH}} v^{2}-m_{3}^{2}\right)\left(2 \lambda_{\mathrm{BB}} v_{\mathrm{B}}^{2}-m_{3}^{2}\right)-\lambda_{\mathrm{HB}}^{2} v^{2} v_{\mathrm{B}}^{2} . \tag{A6}
\end{align*}
$$

## Appendix B

## The loop functions

The loop functions in Eq. (18) and Eq. (19) are given as

$$
\begin{align*}
A_{1}(x) & =-\left[2 x^{2}+3 x+3(2 x-1) g(x)\right] / x^{2} \\
A_{1 / 2}(x) & =2[x+(x-1) g(x)] / x^{2} \\
A_{0}(x) & =-(x-g(x)) / x^{2} \tag{B1}
\end{align*}
$$

with

$$
g(x)=\left\{\begin{array}{l}
\arcsin ^{2} \sqrt{x}, x \leqslant 1  \tag{B2}\\
-\frac{1}{4}\left[\ln \frac{1+\sqrt{1-1 / x}}{1-\sqrt{1-1 / x}}-\mathrm{i} \pi\right]^{2}, x>1
\end{array}\right.
$$

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