

125 GeV Higgs decay with lepton flavor violation in the $\mu\nu$ SSM *

Hai-Bin Zhang(张海斌)^{1;1)} Tai-Fu Feng(冯太傅)^{1,2;2)} Shu-Min Zhao(赵树民)¹,
Yu-Li Yan(阎玉立)¹ Fei Sun(孙飞)³

¹ Department of Physics, Hebei University, Baoding 071002, China

² State Key Laboratory of Theoretical Physics (KLTP),

Institute of Theoretical Physics, Chinese Academy of Sciences, Beijing 100190, China

³ College of Science, China Three Gorges University, Yichang 443002, China

Abstract: Recently, the CMS and ATLAS collaborations have reported direct searches for the 125 GeV Higgs decay with lepton flavor violation, $h \rightarrow \mu\tau$. In this work, we analyze the signal of the lepton flavour violating (LFV) Higgs decay $h \rightarrow \mu\tau$ in the μ from ν Supersymmetric Standard Model ($\mu\nu$ SSM) with slepton flavor mixing. Simultaneously, we consider the constraints from the LFV decay $\tau \rightarrow \mu\gamma$, the muon anomalous magnetic dipole moment and the lightest Higgs mass around 125 GeV.

Keywords: supersymmetry, Higgs decay, lepton flavor violation

PACS: 12.60.Jv, 14.80.Da, 11.30.Fs **DOI:** 10.1088/1674-1137/41/4/043106

1 Introduction

The discovery of the Higgs boson by the ATLAS and CMS Collaborations [1, 2] is a great success of the Large Hadron Collider (LHC). Combining the updated data of the ATLAS and CMS Collaborations, the measured mass of the Higgs boson now is [3]

$$m_h = 125.09 \pm 0.24 \text{ GeV}. \quad (1)$$

The next step is focusing on searching for its properties. In the Standard Model (SM), which is renormalizable, lepton flavour violating (LFV) Higgs decays are forbidden [4]. But recently, a direct search for the 125 GeV Higgs decay with lepton flavor violation, $h \rightarrow \mu\tau$, has been described by the CMS collaboration [5, 6]. The upper limit on the branching ratio of $h \rightarrow \mu\tau$ at 95% confidence level (CL) is [6]

$$Br(h \rightarrow \mu\tau) < 1.20 \times 10^{-2}. \quad (2)$$

Here, interpreted as a signal, $\mu\tau$ means the final state consisting of $\bar{\mu}\tau$ and $\mu\bar{\tau}$.

The ATLAS Collaboration gives the constraint on the branching ratio of $h \rightarrow \mu\tau$ at 95% CL to be [7, 8]

$$Br(h \rightarrow \mu\tau) < 1.43 \times 10^{-2}. \quad (3)$$

The ATLAS and CMS experiments do not currently show a significant deviation from the SM. Therefore, the experiments still need to make more precise measurements in the future.

LFV Higgs decays can occur naturally in models beyond the SM, such as supersymmetric models [9–17], composite Higgs boson models [18, 19], Randall-Sundrum models [20–22], and many others [23–40]. Due to the running of the LHC, LFV Higgs decays have recently been discussed within various theoretical frameworks [41–105]. In this paper, we will study the LFV Higgs decay $h \rightarrow \mu\tau$ in the “ μ from ν Supersymmetric Standard Model” ($\mu\nu$ SSM) [106–108]. As an extension of the Minimal Supersymmetric Standard Model (MSSM) [109–113], the $\mu\nu$ SSM solves the μ problem [114] of the MSSM, through the R-parity breaking couplings $\lambda_i \hat{\nu}_i^c \hat{H}_d^a \hat{H}_u^b$ in the superpotential. The μ term is generated spontaneously via the nonzero vacuum expectative values (VEVs) of right-handed sneutrinos, $\mu = \lambda_i \langle \hat{\nu}_i^c \rangle$, when the electroweak symmetry is broken (EWSB). In addition, nonzero VEVs of sneutrinos in the $\mu\nu$ SSM can generate three tiny massive Majorana

Received 05 September 2016, Revised 21 November 2016

* Supported by Major Project of NNSFC (11535002) and NNSFC (11275036, 11647120), the Open Project Program of State Key Laboratory of Theoretical Physics, Institute of Theoretical Physics, Chinese Academy of Sciences, China (Y5KF131CJ1), the Natural Science Foundation of Hebei province (A2013201277, A2016201010, A2016201069), Hebei Key Lab of Optic-Electronic Information and Materials, and Midwest Universities Comprehensive Strength Promotion Project

1) E-mail: hbzhang@hbu.edu.cn

2) E-mail: fengtf@hbu.edu.cn

 Content from this work may be used under the terms of the Creative Commons Attribution 3.0 licence. Any further distribution of this work must maintain attribution to the author(s) and the title of the work, journal citation and DOI. Article funded by SCOAP³ and published under licence by Chinese Physical Society and the Institute of High Energy Physics of the Chinese Academy of Sciences and the Institute of Modern Physics of the Chinese Academy of Sciences and IOP Publishing Ltd

neutrinos at tree level through the seesaw mechanism [106–108, 115–119].

Within the $\mu\nu$ SUSY, we have studied some LFV processes, $l_j^- \rightarrow l_i^- \gamma$, $l_j^- \rightarrow l_i^- l_i^- l_i^+$, muon conversion to electrons in nuclei and $Z \rightarrow l_i^\pm l_j^\mp$ in our previous work [120–122]. The numerical results show that the LFV rates for $l_j - l_i$ transitions in the $\mu\nu$ SUSY depend on the slepton flavor mixing, and the present experimental limits for the branching ratio of $l_j^- \rightarrow l_i^- \gamma$ constrain the slepton mixing parameters most strictly [122]. In this work, considering the constraint of $\tau \rightarrow \mu\gamma$, we continue to analyze the LFV Higgs decay $h \rightarrow \mu\tau$ in the $\mu\nu$ SUSY with slepton flavor mixing.

The paper is organized as follows. In Section 2, we briefly present the $\mu\nu$ SUSY, including its superpotential and the general soft SUSY-breaking terms. Section 3 contains the analytical expressions of the 125 GeV Higgs decay with lepton flavor violation in the $\mu\nu$ SUSY. The numerical analysis and the summary are given in Section 4 and Section 5, respectively. Some formulae are collected in Appendix and Appendix .

2 The $\mu\nu$ SUSY

In addition to the superfields of the MSSM, the $\mu\nu$ SUSY introduces right-handed neutrino superfields $\tilde{\nu}_i^c$ ($i = 1, 2, 3$). Besides the MSSM Yukawa couplings for quarks and charged leptons, the superpotential of the $\mu\nu$ SUSY contains Yukawa couplings for neutrinos, two additional types of terms involving the Higgs doublet superfields \hat{H}_u and \hat{H}_d , and the right-handed neutrino superfields $\tilde{\nu}_i^c$, [106]

$$W = \epsilon_{ab} \left(Y_{u_{ij}} \hat{H}_u^b \hat{Q}_i^a \hat{u}_j^c + Y_{d_{ij}} \hat{H}_d^a \hat{Q}_i^b \hat{d}_j^c + Y_{e_{ij}} \hat{H}_d^a \hat{L}_i^b \hat{e}_j^c \right) + \epsilon_{ab} Y_{\nu_{ij}} \hat{H}_u^b \hat{L}_i^a \tilde{\nu}_j^c - \epsilon_{ab} \lambda_i \tilde{\nu}_i^c \hat{H}_d^a \hat{H}_u^b + \frac{1}{3} \kappa_{ijk} \tilde{\nu}_i^c \tilde{\nu}_j^c \tilde{\nu}_k^c, \quad (4)$$

where $\hat{H}_u^T = (\hat{H}_u^+, \hat{H}_u^0)$, $\hat{H}_d^T = (\hat{H}_d^0, \hat{H}_d^-)$, $\hat{Q}_i^T = (\hat{u}_i, \hat{d}_i)$, $\hat{L}_i^T = (\hat{\nu}_i, \hat{e}_i)$ are $SU(2)$ doublet superfields, and \hat{u}_i^c , \hat{d}_i^c , and \hat{e}_i^c denote the singlet up-type quark, down-type quark and charged lepton superfields, respectively. Here, Y , λ , and κ are dimensionless matrices, a vector, and a totally symmetric tensor. $i, j, k = 1, 2, 3$ are the generation indices, $a, b = 1, 2$ are the $SU(2)$ indices with antisymmetric tensor $\epsilon_{12} = 1$. In the superpotential, the last two terms explicitly violate lepton number and R-parity. Note that the summation convention is implied on repeated indices in this paper.

Once EWSB occurs, the neutral scalars develop in general the VEVs:

$$\langle H_d^0 \rangle = v_d, \quad \langle H_u^0 \rangle = v_u, \quad \langle \tilde{\nu}_i \rangle = v_{\nu_i}, \quad \langle \tilde{\nu}_i^c \rangle = v_{\nu_i^c}. \quad (5)$$

Then, the terms $\epsilon_{ab} Y_{\nu_{ij}} \hat{H}_u^b \hat{L}_i^a \tilde{\nu}_j^c$ and $\epsilon_{ab} \lambda_i \tilde{\nu}_i^c \hat{H}_d^a \hat{H}_u^b$ in the superpotential can generate the effective bilinear terms

$\epsilon_{ab} \varepsilon_i \hat{H}_u^b \hat{L}_i^a$ and $\epsilon_{ab} \mu \hat{H}_d^a \hat{H}_u^b$, with $\varepsilon_i = Y_{\nu_{ij}} \langle \tilde{\nu}_j^c \rangle$ and $\mu = \lambda_i \langle \tilde{\nu}_i^c \rangle$. One can define the neutral scalars as

$$\begin{aligned} H_d^0 &= \frac{h_d + iP_d}{\sqrt{2}} + v_d, & \tilde{\nu}_i &= \frac{(\tilde{\nu}_i)^R + i(\tilde{\nu}_i)^S}{\sqrt{2}} + v_{\nu_i}, \\ H_u^0 &= \frac{h_u + iP_u}{\sqrt{2}} + v_u, & \tilde{\nu}_i^c &= \frac{(\tilde{\nu}_i^c)^R + i(\tilde{\nu}_i^c)^S}{\sqrt{2}} + v_{\nu_i^c}. \end{aligned} \quad (6)$$

In the framework of supergravity-mediated supersymmetry breaking, the general soft SUSY-breaking terms of the $\mu\nu$ SUSY are given by

$$\begin{aligned} &-\mathcal{L}_{\text{soft}} \\ &= m_{\tilde{Q}_{ij}}^2 \tilde{Q}_i^{a*} \tilde{Q}_j^a + m_{\tilde{u}_{ij}^c}^2 \tilde{u}_i^{c*} \tilde{u}_j^c + m_{\tilde{d}_{ij}^c}^2 \tilde{d}_i^{c*} \tilde{d}_j^c + m_{\tilde{L}_{ij}}^2 \tilde{L}_i^{a*} \tilde{L}_j^a \\ &\quad + m_{\tilde{e}_{ij}^c}^2 \tilde{e}_i^{c*} \tilde{e}_j^c + m_{H_d}^2 H_d^{a*} H_d^a + m_{H_u}^2 H_u^{a*} H_u^a + m_{\tilde{\nu}_{ij}^c}^2 \tilde{\nu}_i^{c*} \tilde{\nu}_j^c \\ &\quad + \epsilon_{ab} \left[(A_u Y_u)_{ij} H_u^b \tilde{Q}_i^a \tilde{u}_j^c + (A_d Y_d)_{ij} H_d^a \tilde{Q}_i^b \tilde{d}_j^c \right. \\ &\quad \left. + (A_e Y_e)_{ij} H_d^a \tilde{L}_i^b \tilde{e}_j^c + \text{H.c.} \right] + \left[\epsilon_{ab} (A_\nu Y_\nu)_{ij} H_u^a \tilde{L}_i^b \tilde{\nu}_j^c \right. \\ &\quad \left. - \epsilon_{ab} (A_\lambda \lambda)_{ij} \tilde{\nu}_i^c H_d^a H_u^b + \frac{1}{3} (A_\kappa \kappa)_{ijk} \tilde{\nu}_i^c \tilde{\nu}_j^c \tilde{\nu}_k^c + \text{H.c.} \right] \\ &\quad - \frac{1}{2} \left(M_3 \tilde{\lambda}_3 \tilde{\lambda}_3 + M_2 \tilde{\lambda}_2 \tilde{\lambda}_2 + M_1 \tilde{\lambda}_1 \tilde{\lambda}_1 + \text{H.c.} \right). \end{aligned} \quad (7)$$

Here, the first two lines contain mass squared terms of squarks, sleptons, Higgses and sneutrinos. The next three lines include the trilinear scalar couplings. In the last line, M_3 , M_2 , and M_1 represent the Majorana masses corresponding to $SU(3)$, $SU(2)$, and $U(1)$ gauginos $\tilde{\lambda}_3$, $\tilde{\lambda}_2$, and $\tilde{\lambda}_1$, respectively. In addition, the tree-level scalar potential receives the usual D - and F -term contributions [107].

In the $\mu\nu$ SUSY, the quadratic potential includes

$$\begin{aligned} V_{\text{quadratic}} &= \frac{1}{2} S'^T M_S^2 S' + \frac{1}{2} P'^T M_P^2 P' + S'^{-T} M_{S^\pm}^2 S'^+ \\ &\quad + \left(\frac{1}{2} \chi'^{0T} M_n \chi'^0 + \Psi^{-T} M_c \Psi^+ + \text{H.c.} \right) + \dots, \end{aligned} \quad (8)$$

where in the unrotated basis $S'^T = (h_d, h_u, (\tilde{\nu}_i)^R, (\tilde{\nu}_i^c)^R)$, $P'^T = (P_d, P_u, (\tilde{\nu}_i)^S, (\tilde{\nu}_i^c)^S)$, $S'^{\pm T} = (H_d^\pm, H_u^\pm, \tilde{e}_{L_i}^\pm, \tilde{e}_{R_i}^\pm)$, $\Psi^{-T} = (-i\tilde{\lambda}^-, \tilde{H}_d^-, e_{L_i}^-)$, $\Psi^{+T} = (-i\tilde{\lambda}^+, \tilde{H}_u^+, e_{R_i}^+)$ and $\chi'^{0T} = (\tilde{B}^0, \tilde{W}^0, \tilde{H}_d, \tilde{H}_u, \nu_{R_i}, \nu_{L_i})$. The concrete expressions for the independent coefficients of mass matrices M_S^2 , M_P^2 , $M_{S^\pm}^2$, M_n and M_c can be found in Ref. [121]. Using 8×8 unitary matrices R_S , R_P and R_{S^\pm} , the unrotated bases S' , P' and $S^{\pm'}$ can be respectively rotated to the mass eigenvectors S , P and S^\pm :

$$S' = R_S S, \quad P' = R_P P, \quad S^{\pm'} = R_{S^\pm} S^\pm. \quad (9)$$

Through the unitary matrices Z_n , Z_- and Z_+ , neutral and charged fermions can also be rotated to the mass eigenvectors χ^0 and χ , respectively.

3 125 GeV Higgs decay with lepton flavor violation

The corresponding effective amplitude for 125 GeV Higgs decay with lepton flavor violation $h \rightarrow \bar{l}_i l_j$ can be written as

$$\mathcal{M} = \bar{l}_i (F_L^{ij} P_L + F_R^{ij} P_R) l_j, \quad (10)$$

with

$$F_{L,R}^{ij} = F_{L,R}^{(V)ij} + F_{L,R}^{(S)ij}, \quad (11)$$

where $F_{L,R}^{(V)ij}$ denotes the contributions from the vertex diagrams in Fig. 1, and $F_{L,R}^{(S)ij}$ stands for the contributions from the self-energy diagrams in Fig. 2, respectively.

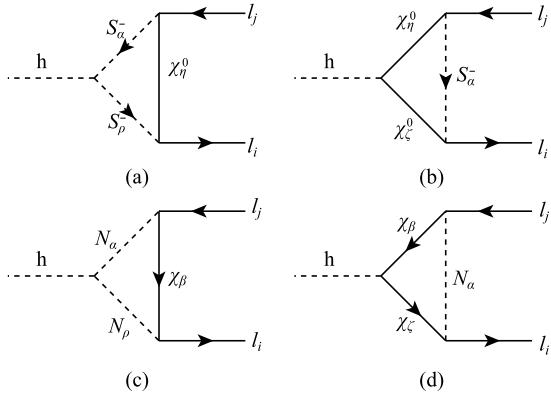


Fig. 1. Vertex diagrams for $h \rightarrow \bar{l}_i l_j$. (a,b) represent the contributions from charged scalar $S_{\alpha,\rho}^-$ and neutral fermion $\chi_{\eta,\zeta}^0$ loops, while (c,d) represent the contributions from neutral scalar $N_{\alpha,\rho}$ ($N=S,P$) and charged fermion $\chi_{\beta,\zeta}$ loops.

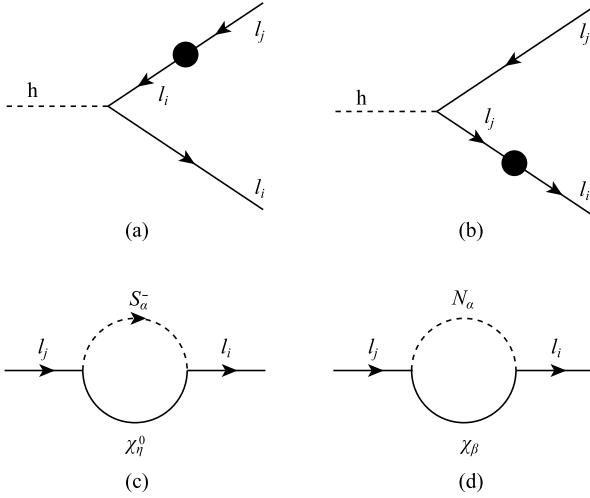


Fig. 2. Self-energy diagrams contributing to $h \rightarrow \bar{l}_i l_j$ in the $\mu\nu$ SSM. The blob in (a,b) indicates the self-energy contributions from (c,d).

The one-loop vertex diagrams for $h \rightarrow \bar{l}_i l_j$ in the $\mu\nu$ SSM are depicted by Fig. 1. Then, we can have

$$F_{L,R}^{(V)ij} = F_{L,R}^{(a)ij} + F_{L,R}^{(b)ij} + F_{L,R}^{(c)ij} + F_{L,R}^{(d)ij}, \quad (12)$$

where $F_{L,R}^{(a,b)ij}$ denotes the contributions from charged scalar $S_{\alpha,\rho}^-$ and neutral fermion $\chi_{\eta,\zeta}^0$ loops, and $F_{L,R}^{(c,d)ij}$ stands for the contributions from the neutral scalar $N_{\alpha,\rho}$ ($N=S,P$) and charged fermion $\chi_{\beta,\zeta}$ loops, respectively. After integrating the heavy freedoms out, we formulate the neutral fermion loop contributions $F_{L,R}^{(a,b)ij}$ as follows:

$$\begin{aligned} F_L^{(a)ij} &= \frac{m_{\chi_\eta^0} C_{1\alpha\rho}^{S^\pm}}{m_W^2} C_L^{S_\rho^- \chi_\eta^0 \bar{l}_i} C_L^{S_\alpha^- * l_j \bar{\chi}_\eta^0} G_1(x_{\chi_\eta^0}, x_{S_\alpha^-}, x_{S_\rho^-}), \\ F_L^{(b)ij} &= \frac{m_{\chi_\zeta^0} m_{\chi_\eta^0}}{m_W^2} C_L^{S_\alpha^- \chi_\zeta^0 \bar{l}_i} C_L^{h \chi_\eta^0 \bar{\chi}_\zeta^0} C_L^{S_\alpha^- * l_j \bar{\chi}_\eta^0} G_1(x_{S_\alpha^-}, x_{\chi_\zeta^0}, x_{\chi_\eta^0}) \\ &\quad + C_L^{S_\alpha^- \chi_\zeta^0 \bar{l}_i} C_R^{h \chi_\eta^0 \bar{\chi}_\zeta^0} C_L^{S_\alpha^- * l_j \bar{\chi}_\eta^0} G_2(x_{S_\alpha^-}, x_{\chi_\zeta^0}, x_{\chi_\eta^0}), \\ F_R^{(a,b)ij} &= F_L^{(a,b)ij} \Big|_{L \leftrightarrow R}. \end{aligned} \quad (13)$$

Here, the concrete expressions for couplings C (and below) can be found in Appendix A and Ref. [123], $x=m^2/m_W^2$, m is the mass for the corresponding particle, and the loop functions G_i are given as

$$\begin{aligned} G_1(x_1, x_2, x_3) &= \frac{1}{16\pi^2} \left[\frac{x_1 \ln x_1}{(x_2 - x_1)(x_1 - x_3)} \right. \\ &\quad \left. + \frac{x_2 \ln x_2}{(x_1 - x_2)(x_2 - x_3)} + \frac{x_3 \ln x_3}{(x_1 - x_3)(x_3 - x_2)} \right], \end{aligned} \quad (14)$$

$$\begin{aligned} G_2(x_1, x_2, x_3) &= \frac{1}{16\pi^2} \left[\frac{x_1^2 \ln x_1}{(x_2 - x_1)(x_1 - x_3)} \right. \\ &\quad \left. + \frac{x_2^2 \ln x_2}{(x_1 - x_2)(x_2 - x_3)} + \frac{x_3^2 \ln x_3}{(x_1 - x_3)(x_3 - x_2)} \right]. \end{aligned} \quad (15)$$

In a similar way, the charged fermion loop contributions $F_{L,R}^{(c,d)ij}$ are

$$\begin{aligned} F_L^{(c)ij} &= \sum_{N=S,P} \frac{m_{\chi_\beta} C_{1\alpha\rho}^N}{m_W^2} C_L^{N_\rho \chi_\beta \bar{l}_i} C_L^{N_\alpha l_j \bar{\chi}_\beta} G_1(x_{\chi_\beta}, x_{N_\alpha}, x_{N_\rho}), \\ F_L^{(d)ij} &= \sum_{N=S,P} \left[C_L^{N_\alpha \chi_\beta \bar{l}_i} C_R^{h \chi_\beta \bar{\chi}_\zeta} C_L^{N_\alpha l_j \bar{\chi}_\beta} G_2(x_{N_\alpha}, x_{\chi_\zeta}, x_{\chi_\beta}) \right. \\ &\quad \left. + \frac{m_{\chi_\zeta} m_{\chi_\beta}}{m_W^2} C_L^{N_\alpha \chi_\zeta \bar{l}_i} C_R^{h \chi_\beta \bar{\chi}_\zeta} C_L^{N_\alpha l_j \bar{\chi}_\beta} G_1(x_{N_\alpha}, x_{\chi_\zeta}, x_{\chi_\beta}) \right], \\ F_R^{(c,d)ij} &= F_L^{(c,d)ij} \Big|_{L \leftrightarrow R}. \end{aligned} \quad (16)$$

In Fig. 2, we show the self-energy diagrams contributing to $h \rightarrow \bar{l}_i l_j$ in the $\mu\nu$ SSM. The contributions from the self-energy diagrams $F_{L,R}^{(S)ij}$ can be given as

$$F_{L,R}^{(S)ij} = F_{L,R}^{(Sa)ij} + F_{L,R}^{(Sb)ij}, \quad (17)$$

with

$$\begin{aligned} F_L^{(Sa)ij} &= \frac{C_L^{\text{hl}_i\bar{l}_i}}{m_{l_j}^2 - m_{l_i}^2} \left\{ m_{l_j}^2 \Sigma_R(m_{l_j}^2) + m_{l_j}^2 \Sigma_{Rs}(m_{l_j}^2) \right. \\ &\quad \left. + m_{l_i} [m_{l_j} \Sigma_L(m_{l_j}^2) + m_{l_j} \Sigma_{Ls}(m_{l_j}^2)] \right\}, \\ F_L^{(Sb)ij} &= \frac{C_L^{\text{hl}_j\bar{l}_j}}{m_{l_i}^2 - m_{l_j}^2} \left\{ m_{l_i}^2 \Sigma_L(m_{l_i}^2) + m_{l_i} m_{l_j} \Sigma_{Rs}(m_{l_i}^2) \right. \\ &\quad \left. + m_{l_j} [m_{l_i} \Sigma_R(m_{l_i}^2) + m_{l_i} \Sigma_{Ls}(m_{l_i}^2)] \right\}, \\ F_R^{(Sa,Sb)ij} &= F_L^{(Sa,Sb)ij} \Big|_{L \leftrightarrow R}. \end{aligned} \quad (18)$$

The Σ of the self-energy diagrams in Fig. 2(c,d) can be obtained below

$$\begin{aligned} \Sigma_L(p^2) &= -\frac{1}{16\pi^2} \left\{ B_1(p^2, m_{\chi_\eta^0}^2, m_{S_\alpha^-}^2) C_L^{S_\alpha^- \chi_\eta^0 \bar{l}_i} C_R^{S_\alpha^- * l_j \bar{\chi}_\eta^0} \right. \\ &\quad + \sum_{N=S,P} B_1(p^2, m_{\chi_\beta^0}^2, m_{N_\alpha}^2) \\ &\quad \cdot C_L^{N_\alpha \chi_\beta \bar{l}_i} C_R^{N_\alpha l_j \bar{\chi}_\beta} \Big\}, \\ m_{l_j} \Sigma_{Ls}(p^2) &= \frac{1}{16\pi^2} \left\{ m_{\chi_\eta^0} B_0(p^2, m_{\chi_\eta^0}^2, m_{S_\alpha^-}^2) \right. \\ &\quad C_L^{S_\alpha^- \chi_\eta^0 \bar{l}_i} C_L^{S_\alpha^- * l_j \bar{\chi}_\eta^0} \\ &\quad + \sum_{N=S,P} m_{\chi_\beta} B_0(p^2, m_{\chi_\beta}^2, m_{N_\alpha}^2) \\ &\quad \left. C_L^{N_\alpha \chi_\beta \bar{l}_i} C_L^{N_\alpha l_j \bar{\chi}_\beta} \right\}, \\ \Sigma_R(p^2) &= \Sigma_L(p^2) \Big|_{L \leftrightarrow R}, \\ m_{l_j} \Sigma_{Rs}(p^2) &= m_{l_j} \Sigma_{Ls}(p^2) \Big|_{L \leftrightarrow R}. \end{aligned} \quad (19)$$

Here, $B_{0,1}(p^2, m_0^2, m_1^2)$ are two-point functions [124–130].

Then, we can obtain the decay width of $h \rightarrow \bar{l}_i l_j$ [9, 14]

$$\Gamma(h \rightarrow \bar{l}_i l_j) \approx \frac{m_h}{16\pi} \left(|F_L^{ij}|^2 + |F_R^{ij}|^2 \right). \quad (20)$$

If interpreted as a signal, the decay width of $h \rightarrow l_i l_j$ is

$$\Gamma(h \rightarrow l_i l_j) = \Gamma(h \rightarrow \bar{l}_i l_j) + \Gamma(h \rightarrow \bar{l}_j l_i), \quad (21)$$

and the branching ratio of $h \rightarrow l_i l_j$ is

$$Br(h \rightarrow l_i l_j) = \Gamma(h \rightarrow l_i l_j) / \Gamma_h, \quad (22)$$

where $\Gamma_h \approx 4.1 \times 10^{-3}$ GeV [131] denotes the total decay width of the 125 GeV Higgs boson.

4 Numerical analysis

In order to obtain transparent numerical results in the $\mu\nu$ SSM, we take the minimal flavor violation (MFV) assumptions for some parameters, which assume

$$\begin{aligned} \kappa_{ijk} &= \kappa \delta_{ij} \delta_{jk}, \quad (A_\kappa \kappa)_{ijk} = A_\kappa \kappa \delta_{ij} \delta_{jk}, \quad \lambda_i = \lambda, \\ (A_\lambda \lambda)_i &= A_\lambda \lambda, \quad Y_{e_{ij}} = Y_{e_i} \delta_{ij}, \quad Y_{\nu_{ij}} = Y_{\nu_i} \delta_{ij}, \end{aligned}$$

$$\begin{aligned} v_{\nu_i^c} &= v_{\nu^c}, \quad (A_\nu Y_\nu)_{ij} = a_{\nu_i} \delta_{ij}, \quad m_{\tilde{\nu}_{ij}^c}^2 = m_{\tilde{\nu}_i^c}^2 \delta_{ij}, \\ m_{\tilde{Q}_{ij}}^2 &= m_{\tilde{Q}_i}^2 \delta_{ij}, \quad m_{\tilde{u}_{ij}^c}^2 = m_{\tilde{u}_i^c}^2 \delta_{ij}, \quad m_{\tilde{d}_{ij}^c}^2 = m_{\tilde{d}_i^c}^2 \delta_{ij}, \end{aligned} \quad (23)$$

where $i, j, k = 1, 2, 3$. $m_{\tilde{\nu}_{ij}^c}^2$ can be constrained by the minimization conditions of the neutral scalar potential seen in Ref. [121]. To agree with experimental observations on quark mixing, one can have

$$\begin{aligned} Y_{u_{ij}} &= Y_{u_i} V_{L_{ij}}^u, \quad (A_u Y_u)_{ij} = A_{u_i} Y_{u_{ij}}, \\ Y_{d_{ij}} &= Y_{d_i} V_{L_{ij}}^d, \quad (A_d Y_d)_{ij} = A_{d_i} Y_{d_{ij}}, \end{aligned} \quad (24)$$

and $V = V_L^u V_L^{d\dagger}$ denotes the CKM matrix.

For the trilinear coupling matrix $(A_e Y_e)$ and soft breaking slepton mass matrices $m_{\tilde{L},\tilde{e}^c}^2$, we will take into account the off-diagonal terms for the matrices, which are named the slepton flavor mixings and are defined by [132–137]

$$m_{\tilde{L}}^2 = \begin{pmatrix} 1 & \delta_{12}^{LL} & \delta_{13}^{LL} \\ \delta_{12}^{LL} & 1 & \delta_{23}^{LL} \\ \delta_{13}^{LL} & \delta_{23}^{LL} & 1 \end{pmatrix} m_L^2, \quad (25)$$

$$m_{\tilde{e}^c}^2 = \begin{pmatrix} 1 & \delta_{12}^{RR} & \delta_{13}^{RR} \\ \delta_{12}^{RR} & 1 & \delta_{23}^{RR} \\ \delta_{13}^{RR} & \delta_{23}^{RR} & 1 \end{pmatrix} m_E^2, \quad (26)$$

$$(A_e Y_e) = \begin{pmatrix} m_{l_1} A_e & \delta_{12}^{LR} m_L m_E & \delta_{13}^{LR} m_L m_E \\ \delta_{12}^{LR} m_L m_E & m_{l_2} A_e & \delta_{23}^{LR} m_L m_E \\ \delta_{13}^{LR} m_L m_E & \delta_{23}^{LR} m_L m_E & m_{l_3} A_e \end{pmatrix} \frac{1}{v_d}. \quad (27)$$

The following numerical results will show that the branching ratio of $h \rightarrow \mu\tau$ depends on the slepton mixing parameters δ_{23}^{XX} ($X = L, R$).

At first, the constraints from some experiments should be considered. Through our previous work [119], we have discussed in detail how the neutrino oscillation data constrain neutrino Yukawa couplings $Y_{\nu_i} \sim \mathcal{O}(10^{-7})$ and left-handed sneutrino VEVs $v_{\nu_i} \sim \mathcal{O}(10^{-4}\text{GeV})$ via the seesaw mechanism. Here, due to the neutrino sector only weakly affecting $h \rightarrow \mu\tau$, we can take no account of the constraints from neutrino experiment data.

The neutral Higgs with mass around 125 GeV reported by ATLAS and CMS contributes a strict constraint on the relevant parameters of the $\mu\nu$ SSM. For a large mass of the pseudoscalar M_A and moderate $\tan\beta$, the SM-like Higgs mass of the $\mu\nu$ SSM is approximately written as [107, 138]

$$m_h^2 \approx q m_Z^2 \cos^2 2\beta + \frac{6\lambda^2 s_w^2 c_w^2}{e^2} m_Z^2 \sin^2 2\beta + \Delta m_h^2. \quad (28)$$

Compared with the MSSM, the $\mu\nu$ SUSY gets an additional term, $\frac{6\lambda^2 s_w^2 c_w^2}{e^2} m_Z^2 \sin^2 2\beta$. Thus, the SM-like Higgs in the $\mu\nu$ SUSY can easily account for the mass around 125 GeV, especially for small $\tan\beta$. Including two-loop leading-log effects, the main radiative corrections Δm_h^2 can be given as [139–141]

$$\begin{aligned}\Delta m_h^2 &= \frac{3m_t^4}{4\pi^2 v^2} \left[\left(t + \frac{1}{2} \tilde{X}_t \right) \right. \\ &\quad \left. + \frac{1}{16\pi^2} \left(\frac{3m_t^2}{2v^2} - 32\pi\alpha_3 \right) (t^2 + \tilde{X}_t t) \right], \\ t &= \log \frac{M_S^2}{m_t^2}, \quad \tilde{X}_t = \frac{2\tilde{A}_t^2}{M_S^2} \left(1 - \frac{\tilde{A}_t^2}{12M_S^2} \right),\end{aligned}\quad (29)$$

where $v = 174$ GeV, α_3 is the strong coupling constant, $M_S = \sqrt{m_{\tilde{t}_1} m_{\tilde{t}_2}}$ with $m_{\tilde{t}_{1,2}}$ denoting the stop masses, $\tilde{A}_t = A_t - \mu \cot\beta$ with $A_t = A_{u_3}$ being the trilinear Higgs-stop coupling and $\mu = 3\lambda v_{\nu^c}$ denoting the Higgsino mass parameter.

We also impose a constraint on the SUSY contribution to the muon magnetic dipole moment a_μ in the $\mu\nu$ SUSY, which is given in Appendix for convenience. The difference between experiment and the SM prediction on a_μ is [142–144]

$$\Delta a_\mu = a_\mu^{\text{exp}} - a_\mu^{\text{SM}} = (24.8 \pm 7.9) \times 10^{-10}, \quad (30)$$

with all errors combining in quadrature. Therefore, the SUSY contribution to a_μ in the $\mu\nu$ SUSY should be constrained as $1.1 \times 10^{-10} \leq \Delta a_\mu \leq 48.5 \times 10^{-10}$, where a 3σ experimental error is considered.

Through analysis of the parameter space of the $\mu\nu$ SUSY in Ref. [107], we can take reasonable parameter values to be $\lambda = 0.1$, $\kappa = 0.4$, $A_\lambda = 500$ GeV, $A_\kappa = -300$ GeV and $A_e = 1$ TeV for simplicity. For the gauginos' Majorana masses, we will choose the approximate GUT relation $M_1 = \frac{\alpha_1^2}{\alpha_2^2} M_2 \approx 0.5M_2$ and

$M_3 = \frac{\alpha_3^2}{\alpha_2^2} M_2 \approx 2.7M_2$. The gluino mass, $m_{\tilde{g}} \approx M_3$, is greater than about 1.2 TeV from the ATLAS and CMS experimental data [145–148]. For simplicity, we could adopt $m_{\tilde{Q}_3} = m_{\tilde{u}_3^c} = m_{\tilde{d}_3^c} = 1.5$ TeV. As key parameters, A_t and $\tan\beta \equiv v_u/v_d$ affect the SM-like Higgs mass. Here, we keep the SM-like Higgs mass $m_h = 125$ GeV as input, and then the value of parameter A_t can be given automatically in the numerical calculation. Then, the free parameters that affect our next analysis are $\tan\beta$, $\mu \equiv 3\lambda v_{\nu^c}$, M_2 , m_L , m_E and slepton mixing parameters δ_{23}^{XX} ($X = L, R$).

It is well known that the lepton flavour violating processes are flavor dependent. The LFV rates for $\mu - \tau$ transitions depend on the slepton mixing parameters

δ_{23}^{XX} ($X = L, R$), which can be confirmed by Fig. 3. The slepton mixing parameters δ_{12}^{XX} and δ_{13}^{XX} ($X = L, R$) hardly affect the LFV rates for $\mu - \tau$ transitions, which play a leading role in the LFV rates for $e - \mu$ and $e - \tau$ transitions. So, we take $\delta_{12}^{XX} = 0$ and $\delta_{13}^{XX} = 0$ ($X = L, R$) here. To produce Fig. 3, we scan the parameter space shown in Table 1. Here the steps are large, because the running of the program is not very fast. However the scanned parameter space is broad enough to contain the possibility of more.

In the scan, we keep the chargino masses $m_{\chi_\beta} > 200$ GeV ($\beta = 1, 2$), the neutral fermion masses $m_{\chi_\eta^0} > 200$ GeV ($\eta = 1, \dots, 7$), and the scalar masses $m_{S_\alpha, P_\alpha, S_\alpha^\pm} > 500$ GeV ($\eta = 2, \dots, 8$), to avoid the range ruled out by the experiments [142]. The results are also constrained by the muon anomalous magnetic dipole moment $1.1 \times 10^{-10} \leq \Delta a_\mu \leq 48.5 \times 10^{-10}$, where a 3σ experimental error is considered. In Ref. [123], we have investigated the signals of the Higgs boson decay channels $h \rightarrow \gamma\gamma$, $h \rightarrow VV^*$ ($V = Z, W$), and $h \rightarrow f\bar{f}$ ($f = b, \tau$) in the $\mu\nu$ SUSY. When the lightest stop mass $m_{\tilde{t}_1} \gtrsim 700$ GeV and the lightest stau mass $m_{\tilde{\tau}_1} \gtrsim 300$ GeV, the signal strengths of these Higgs boson decay channels are in agreement with the SM. Therefore, the scanning results in this paper coincide with the experimental data of these Higgs boson decay channels.

Note that, when the calculation program is scanning one of the slepton mixing parameters δ_{23}^{XX} ($X = L, R$), the other two slepton mixing parameters δ_{23}^{XX} ($X = L, R$) are set to zero. So, we can see the contribution of every slepton mixing parameter alone. Then in Fig. 3, we plot $Br(h \rightarrow \mu\tau)$ varying with slepton mixing parameters δ_{23}^{LR} (a), δ_{23}^{LL} (c), and δ_{23}^{RR} (e) respectively, where the dashed line stands for the upper limit on $Br(h \rightarrow \mu\tau)$ at 95% CL shown in Eq. (2). We also plot $Br(\tau \rightarrow \mu\gamma)$ versus slepton mixing parameters δ_{23}^{LR} (b), δ_{23}^{LL} (d), and δ_{23}^{RR} (f) respectively, where the dashed line denotes the present limit of $Br(\tau \rightarrow \mu\gamma)$ [149]:

$$Br(\tau \rightarrow \mu\gamma) < 4.4 \times 10^{-8}. \quad (31)$$

Here, the red triangles are ruled out by the present limit of $Br(\tau \rightarrow \mu\gamma)$, and the black circles are consistent with the present limit of $Br(\tau \rightarrow \mu\gamma)$.

In Fig. 3, when slepton mixing parameters $\delta_{23}^{XX} = 0$ ($X = L, R$), $Br(h \rightarrow \mu\tau)$ can reach $\mathcal{O}(10^{-23})$ and $Br(\tau \rightarrow \mu\gamma)$ can attain $\mathcal{O}(10^{-27})$, because the contributions can come from the mixing of the particles, which can easily be seen in Eq. (8). These results are too small to detect. However, if the nonzero slepton mixing parameters δ_{23}^{XX} ($X = L, R$) are considered, $Br(h \rightarrow \mu\tau)$ and $Br(\tau \rightarrow \mu\gamma)$ grow quickly. With increasing δ_{23}^{XX} ($X = L, R$), $Br(\tau \rightarrow \mu\gamma)$ can easily go beyond the present experimental limit of $Br(\tau \rightarrow \mu\gamma)$, shown in the plot as the red triangles. Although $Br(h \rightarrow \mu\tau)$ cannot reach

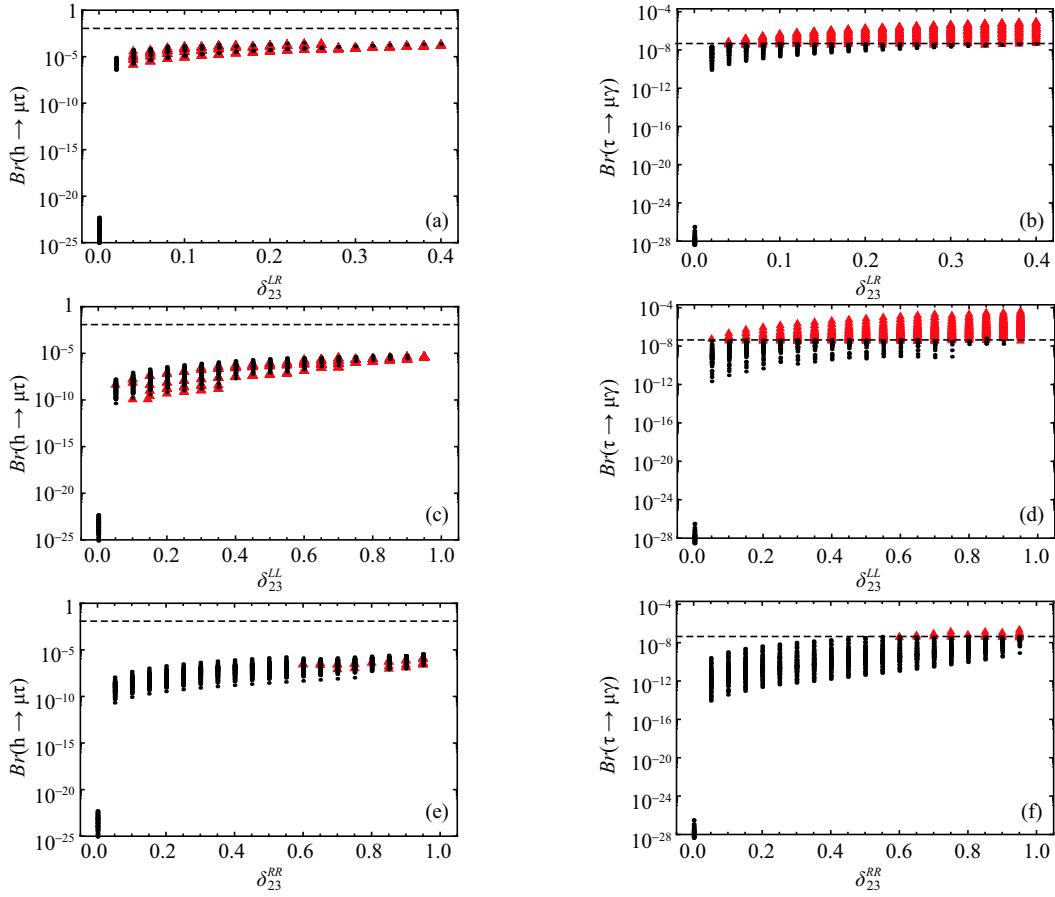


Fig. 3. (color online) $Br(h \rightarrow \mu\tau)$ versus slepton mixing parameters δ_{23}^{LR} (a), δ_{23}^{LL} (c), and δ_{23}^{RR} (e), where the dashed line stands for the upper limit on $Br(h \rightarrow \mu\tau)$ at 95% CL showed in Eq. (2). $Br(\tau \rightarrow \mu\gamma)$ versus slepton mixing parameters δ_{23}^{LR} (b), δ_{23}^{LL} (d), and δ_{23}^{RR} (f), where the dashed line denotes the present limit of $Br(\tau \rightarrow \mu\gamma)$ seen in Eq. (31). Here, the red triangles are ruled out by the present limit of $Br(\tau \rightarrow \mu\gamma)$, and the black circles are consistent with the present limit of $Br(\tau \rightarrow \mu\gamma)$.

the present experimental upper limit of $Br(h \rightarrow \mu\tau)$, $Br(h \rightarrow \mu\tau)$ becomes larger and approaches the present experimental limit with increasing δ_{23}^{XX} ($X = L, R$). Especially in Fig. 3(a), considering nonzero slepton mixing parameters δ_{23}^{LR} , $Br(h \rightarrow \mu\tau)$ can achieve $\mathcal{O}(10^{-4})$, which is below the present experimental limit by just two orders of magnitude. Compared to the MSSM, exotic singlet righthanded neutrino superfields in the $\mu\nu$ SSM induce new sources for lepton-flavor violation, considering that the righthanded neutrino and sneutrinos can mix and couple with the other particles seen in Eq. (8) and Appendix A. In Fig. 3(a,c,e), the red triangles overlap with the black circles, because some parameters strongly affect $Br(\tau \rightarrow \mu\gamma)$ but do not affect $Br(h \rightarrow \mu\tau)$. We will research this further in the following.

To see how other parameters affect the results, we appropriately fix $\delta_{23}^{LR} = 0.02$ and $\delta_{23}^{LL} = \delta_{23}^{RR} = 0.2$. Then, we scan the parameter space shown in Table 2, where $\mu = M_2 = m_L = m_E \equiv M_{\text{SUSY}}$. In the scan-

ning, we also keep the chargino masses $m_{\chi_\beta} > 200$ GeV ($\beta = 1, 2$), the neutral fermion masses $m_{\chi_\eta^0} > 200$ GeV ($\eta = 1, \dots, 7$), and the scalar masses $m_{S_\alpha, P_\alpha, S_\alpha^\pm} > 500$ GeV ($\eta = 2, \dots, 8$), to avoid the range ruled out by the experiments [142]. Then in Fig. 4, we plot $Br(h \rightarrow \mu\tau)$ respectively versus $\tan\beta$ (a) and M_{SUSY} (b), where the dashed line stands for the upper limit on $Br(h \rightarrow \mu\tau)$ at 95% CL shown in Eq. (2). We show $Br(\tau \rightarrow \mu\gamma)$ varying with $\tan\beta$ (c) and M_{SUSY} (d) respectively, where the dashed line denotes the present limit of $Br(\tau \rightarrow \mu\gamma)$ which can be seen in Eq. (31). We also picture the muon anomalous magnetic dipole moment Δa_μ versus $\tan\beta$ (e) and M_{SUSY} (f) respectively, where the gray area denotes the Δa_μ at 3.0σ given in Eq. (30). Here, the red triangles are excluded by the present limit of $Br(\tau \rightarrow \mu\gamma)$, the green squares are eliminated by the Δa_μ at 3.0σ , and the black circles conform to both the present limit of $Br(\tau \rightarrow \mu\gamma)$ and the Δa_μ at 3.0σ .

In Fig. 4(d,f), the numerical results show that $Br(\tau \rightarrow \mu\gamma)$ and the muon anomalous magnetic dipole

moment Δa_μ are decoupling with increasing M_{SUSY} . For large M_{SUSY} , it is hard to give large contribution to Δa_μ . So, the large M_{SUSY} are easily excluded by the Δa_μ at 3.0σ given in Eq. (30), which can be seen in the graph as the green squares. For small M_{SUSY} , there can be a large contribution to $Br(\tau \rightarrow \mu\gamma)$. Therefore, the small M_{SUSY} are easily ruled out by the present experimental limit of $Br(\tau \rightarrow \mu\gamma)$, shown as the red triangles. In Fig. 4(b), $Br(h \rightarrow \mu\tau)$ is non-decoupling with increasing M_{SUSY} , which is in agreement with the research in the MSSM [44, 67]. Due to the introduction of slepton mixing parameters, the non-decoupling behaviour of $Br(h \rightarrow \mu\tau)$ tends to $\mathcal{O}((m_h/M_{\text{SUSY}})^0)$, which is somewhat different from the Appelquist-Carazzone decoupling theorem [150]. (As a side note, in Ref. [151], a non-decoupling behaviour in computation of the Higgs mass showed that it was linked to an ambiguity in the treatment of $\tan\beta$, which is a renormalization scheme dependent parameter.) We can also see that the red triangles overlap with the black circles in Fig. 4(b), because the parameter $\tan\beta$ does not affect $Br(h \rightarrow \mu\tau)$ visibly

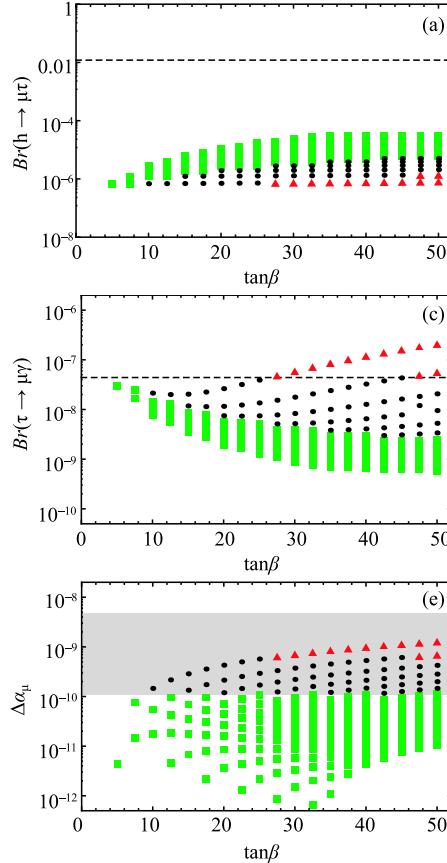


Fig. 4. (color online) $Br(h \rightarrow \mu\tau)$ versus $\tan\beta$ (a) and M_{SUSY} (b), where the dashed line stands for the upper limit on $Br(h \rightarrow \mu\tau)$ at 95% CL shown in Eq. (2). $Br(\tau \rightarrow \mu\gamma)$ versus $\tan\beta$ (c) and M_{SUSY} (d), where the dashed line denotes the present limit of $Br(\tau \rightarrow \mu\gamma)$, which can be seen in Eq. (31). Δa_μ versus $\tan\beta$ (e) and M_{SUSY} (f), where the gray area denotes the Δa_μ at 3.0σ given in Eq. (30). Here, the red triangles are excluded by the present limit of $Br(\tau \rightarrow \mu\gamma)$, the green squares are eliminated by the Δa_μ at 3.0σ , and the black circles simultaneously conform to the present limit of $Br(\tau \rightarrow \mu\gamma)$ and the Δa_μ at 3.0σ .

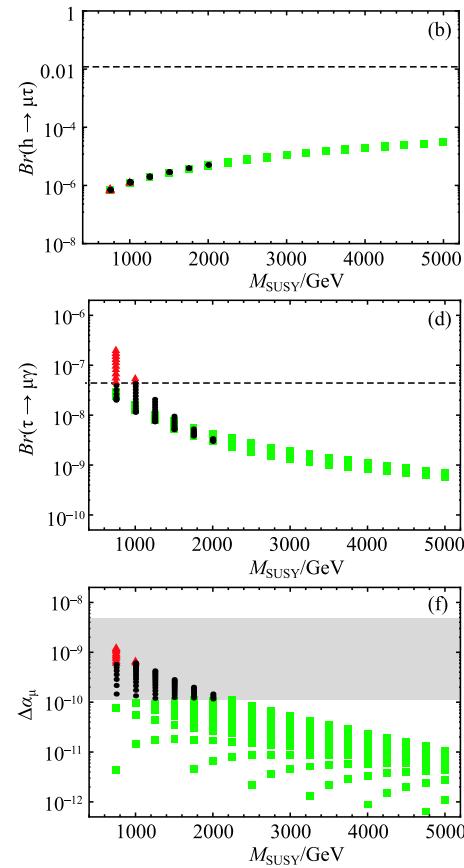
in this parameter space. In Fig. 4(a,c,e), the numerical results show that $Br(h \rightarrow \mu\tau)$, $Br(\tau \rightarrow \mu\gamma)$ and the muon anomalous magnetic dipole moment Δa_μ can have large values when $\tan\beta$ is large.

Table 1. Scanning parameters for Fig. 3.

parameters	min	max	step
$\tan\beta$	5	50	15
$\mu = M_2/\text{GeV}$	500	5000	500
$m_L = m_E/\text{GeV}$	500	5000	500
δ_{23}^{LR}	0	0.4	0.02
δ_{23}^{LL}	0	1.0	0.05
δ_{23}^{RR}	0	1.0	0.05

Table 2. Scanning parameters for Fig. 4, where $\mu = M_2 = m_L = m_E \equiv M_{\text{SUSY}}$.

parameters	min	max	step
$\tan\beta$	5	50	2.5
$M_{\text{SUSY}}/\text{GeV}$	500	5000	250



5 Summary

In this work, we have studied the 125 GeV Higgs decay with lepton flavor violation, $h \rightarrow \mu\tau$, in the framework of the $\mu\nu$ SSM with slepton flavor mixing. The numerical results show that the branching ratio of $h \rightarrow \mu\tau$ depends on the slepton mixing parameters δ_{23}^{XX} ($X = L, R$), because the lepton flavour violating processes are flavor dependent. The branching ratio of $h \rightarrow \mu\tau$ increases with increasing δ_{23}^{XX} ($X = L, R$). Under the experimental constraints of the muon anomalous magnetic

dipole moment, the SM-like Higgs mass around 125 GeV and the present limit of $Br(\tau \rightarrow \mu\gamma)$, the branching ratio of $h \rightarrow \mu\tau$ can reach $\mathcal{O}(10^{-4})$. Compared with the MSSM, exotic singlet righthanded neutrino superfields in the $\mu\nu$ SSM induce new sources for the lepton-flavor violation. Considering that the recent ATLAS and CMS measurements for $h \rightarrow \mu\tau$ do not show a significant deviation from the SM, the experiments still need to make more precise measurements in the future. To detect a Higgs boson lepton flavour violating process is a prospective window to search for new physics.

Appendix A

The couplings

The couplings between CP-even neutral scalars and the other CP-even (or CP-odd) neutral scalars are formulated as

$$\mathcal{L}_{\text{int}} = C_{\alpha\beta\gamma}^S S_\alpha S_\beta S_\gamma + C_{\alpha\beta\gamma}^P S_\alpha P_\beta P_\gamma, \quad (\text{A1})$$

with

$$\begin{aligned} C_{\alpha\beta\gamma}^S &= \frac{-e^2}{4\sqrt{2}s_w^2 c_w^2} \left[v_d R_S^{1\alpha} R_S^{1\beta} R_S^{1\gamma} + v_u R_S^{2\alpha} R_S^{2\beta} R_S^{2\gamma} \right. \\ &\quad \left. + (v_d R_S^{1\alpha} + v_u R_S^{2\alpha}) R_S^{(2+i)\beta} R_S^{(2+i)\gamma} \right] \\ &\quad + \frac{1}{\sqrt{2}} \left[\lambda_i \lambda_i (v_d R_S^{1\alpha} R_S^{2\beta} R_S^{2\gamma} + v_u R_S^{2\alpha} R_S^{1\beta} R_S^{1\gamma}) \right. \\ &\quad \left. - \lambda_i \lambda_j (v_d R_S^{1\alpha} + v_u R_S^{2\alpha}) R_S^{(5+i)\beta} R_S^{(5+j)\gamma} \right] \\ &\quad + \sqrt{2} \kappa_{mij} \kappa_{mkl} v_{\nu_i^c} R_S^{(5+j)\alpha} R_S^{(5+k)\beta} R_S^{(5+l)\gamma} \\ &\quad - \frac{1}{3\sqrt{2}} (A_\kappa \kappa)_{ijk} R_S^{(5+i)\alpha} R_S^{(5+j)\beta} R_S^{(5+k)\gamma} \\ &\quad + \frac{1}{\sqrt{2}} (A_\lambda \lambda)_i R_S^{1\alpha} R_S^{2\beta} R_S^{(5+i)\gamma} \\ &\quad - \frac{1}{\sqrt{2}} \lambda_i \kappa_{ijk} (v_u R_S^{1\alpha} + v_d R_S^{2\alpha}) R_S^{(5+j)\beta} R_S^{(5+k)\gamma}, \end{aligned} \quad (\text{A2})$$

$$\begin{aligned} C_{\alpha\beta\gamma}^P &= \frac{-e^2}{4\sqrt{2}s_w^2 c_w^2} \left[v_d R_S^{1\alpha} R_P^{1\beta} R_P^{1\gamma} \right. \\ &\quad \left. + v_u R_S^{2\alpha} R_P^{2\beta} R_P^{2\gamma} + (v_d R_S^{1\alpha} \right. \\ &\quad \left. + v_u R_S^{2\alpha}) R_P^{(2+i)\beta} R_P^{(2+i)\gamma} \right] \\ &\quad + \frac{1}{\sqrt{2}} \left[\lambda_i \lambda_i (v_d R_S^{1\alpha} R_P^{2\beta} R_P^{2\gamma} + v_u R_S^{2\alpha} R_P^{1\beta} R_P^{1\gamma}) \right. \\ &\quad \left. - \lambda_i \lambda_j (v_d R_S^{1\alpha} + v_u R_S^{2\alpha}) R_P^{(5+i)\beta} R_P^{(5+j)\gamma} \right] \\ &\quad + \sqrt{2} \kappa_{mij} \kappa_{mkl} v_{\nu_i^c} R_S^{(5+l)\alpha} R_P^{(5+j)\beta} R_P^{(5+k)\gamma} \\ &\quad + \frac{1}{\sqrt{2}} (A_\kappa \kappa)_{ijk} R_S^{(5+i)\alpha} R_P^{(5+j)\beta} R_P^{(5+k)\gamma} \\ &\quad - \frac{1}{\sqrt{2}} (A_\lambda \lambda)_i \left[R_S^{1\alpha} R_P^{2\beta} R_P^{(5+i)\gamma} + R_S^{2\alpha} R_P^{1\beta} R_P^{(5+i)\gamma} \right] \end{aligned}$$

$$\begin{aligned} &\quad + R_S^{(5+i)\alpha} R_P^{1\beta} R_P^{2\gamma} \Big] \\ &\quad + \frac{1}{\sqrt{2}} \lambda_i \kappa_{ijk} (v_u R_S^{1\alpha} + v_d R_S^{2\alpha}) R_P^{(5+j)\beta} R_P^{(5+k)\gamma}, \end{aligned} \quad (\text{A3})$$

where the unitary matrices R_S , R_P (and Z_n , Z_- , Z_+ below) can be found in Ref. [121], and the small terms containing $Y_{\nu_i} \sim \mathcal{O}(10^{-7})$ and $v_{\nu_i} \sim \mathcal{O}(10^{-4} \text{ GeV})$ are ignored.

The interaction Lagrangian between CP-even neutral scalars and neutral fermions is formulated as

$$\mathcal{L}_{\text{int}} = S_\alpha \bar{\chi}_\zeta^0 \left(C_L^{S_\alpha \chi_\eta^0 \bar{\chi}_\zeta^0} P_L + C_R^{S_\alpha \chi_\eta^0 \bar{\chi}_\zeta^0} P_R \right) \chi_\eta^0, \quad (\text{A4})$$

where

$$\begin{aligned} C_L^{S_\alpha \chi_\eta^0 \bar{\chi}_\zeta^0} &= \frac{-e}{2s_w c_w} \left(c_w Z_n^{2\eta} - s_w Z_n^{1\eta} \right) \\ &\quad \left(R_S^{1\alpha} Z_n^{3\zeta} - R_S^{2\alpha} Z_n^{4\zeta} + R_S^{(2+i)\alpha} Z_n^{(7+i)\zeta} \right) \\ &\quad - \frac{1}{\sqrt{2}} Y_{\nu_{ij}} \left(R_S^{2\alpha} Z_n^{(7+i)\eta} Z_n^{(4+j)\zeta} \right. \\ &\quad \left. + R_S^{(2+i)\alpha} Z_n^{3\eta} Z_n^{(4+j)\zeta} + R_S^{(5+j)\alpha} Z_n^{3\eta} Z_n^{(7+i)\zeta} \right) \\ &\quad - \frac{1}{\sqrt{2}} \lambda_i \left(R_S^{1\alpha} Z_n^{(4+i)\eta} Z_n^{4\zeta} \right. \\ &\quad \left. + R_S^{2\alpha} Z_n^{(4+i)\eta} Z_n^{3\zeta} + R_S^{(5+i)\alpha} Z_n^{3\eta} Z_n^{4\zeta} \right) \\ &\quad + \frac{1}{\sqrt{2}} \kappa_{ijk} R_S^{(5+i)\alpha} Z_n^{(4+j)\eta} Z_n^{(4+k)\zeta}, \end{aligned} \quad (\text{A5})$$

$$C_R^{S_\alpha \chi_\eta^0 \bar{\chi}_\zeta^0} = \left[C_L^{S_\alpha \chi_\eta^0 \bar{\chi}_\zeta^0} \right]^*, \quad (\text{A6})$$

and

$$P_L = \frac{1}{2}(1 - \gamma^5), \quad P_R = \frac{1}{2}(1 + \gamma^5). \quad (\text{A7})$$

The interaction Lagrangian of neutral scalars and charged fermions can be written as

$$\begin{aligned} \mathcal{L}_{\text{int}} &= S_\alpha \bar{\chi}_\zeta (C_L^{S_\alpha \chi_\beta \bar{\chi}_\zeta} P_L + C_R^{S_\alpha \chi_\beta \bar{\chi}_\zeta} P_R) \chi_\beta \\ &\quad + P_\alpha \bar{\chi}_\zeta (C_L^{P_\alpha \chi_\beta \bar{\chi}_\zeta} P_L + C_R^{P_\alpha \chi_\beta \bar{\chi}_\zeta} P_R) \chi_\beta, \end{aligned} \quad (\text{A8})$$

where the coefficients are

$$\begin{aligned} C_L^{S_\alpha \chi_\beta \bar{\chi}_\zeta} &= \frac{-e}{\sqrt{2}s_W} \left[R_S^{2\alpha} Z_-^{1\beta} Z_+^{2\zeta} + R_S^{1\alpha} Z_-^{2\beta} Z_+^{1\zeta} \right. \\ &\quad \left. + R_S^{(2+i)\alpha} Z_-^{(2+i)\beta} Z_+^{1\zeta} \right] \\ &\quad - \frac{1}{\sqrt{2}} \lambda_i R_S^{(5+i)\alpha} Z_-^{2\beta} Z_+^{2\zeta} \\ &\quad + \frac{1}{\sqrt{2}} Y_{e_{ij}} \left[R_S^{(2+i)\alpha} Z_-^{2\beta} Z_+^{(2+j)\zeta} \right. \\ &\quad \left. - R_S^{1\alpha} Z_-^{(2+i)\beta} Z_+^{(2+j)\zeta} \right] \\ &\quad - \frac{1}{\sqrt{2}} Y_{\nu_{ij}} R_S^{(5+j)\alpha} Z_-^{(2+i)\beta} Z_+^{2\zeta}, \end{aligned} \quad (\text{A9})$$

$$\begin{aligned} C_L^{P_\alpha \chi_\beta \bar{\chi}_\zeta} &= \frac{ie}{\sqrt{2}s_W} \left[R_P^{2\alpha} Z_-^{1\beta} Z_+^{2\zeta} + R_P^{1\alpha} Z_-^{2\beta} Z_+^{1\zeta} \right. \\ &\quad \left. + R_P^{(2+i)\alpha} Z_-^{(2+i)\beta} Z_+^{1\zeta} \right] - \frac{i}{\sqrt{2}} \lambda_i R_P^{(5+i)\alpha} Z_-^{2\beta} Z_+^{2\zeta} \\ &\quad + \frac{i}{\sqrt{2}} Y_{e_{ij}} \left[R_P^{(2+i)\alpha} Z_-^{2\beta} Z_+^{(2+j)\zeta} \right. \\ &\quad \left. - R_P^{1\alpha} Z_-^{(2+i)\beta} Z_+^{(2+j)\zeta} \right] \\ &\quad - \frac{i}{\sqrt{2}} Y_{\nu_{ij}} R_P^{(5+j)\alpha} Z_-^{(2+i)\beta} Z_+^{2\zeta}, \end{aligned} \quad (\text{A10})$$

$$C_R^{S_\alpha \chi_\beta \bar{\chi}_\zeta} = [C_L^{S_\alpha \chi_\beta \bar{\chi}_\zeta}]^*, \quad C_R^{P_\alpha \chi_\beta \bar{\chi}_\zeta} = [C_L^{P_\alpha \chi_\beta \bar{\chi}_\zeta}]^*. \quad (\text{A11})$$

The interaction Lagrangian of charged scalars, charged

Appendix B

Muon MDM in the $\mu\nu$ SUSY

The muon anomalous magnetic dipole moment (MDM) in the $\mu\nu$ SUSY can be given as the effective Lagrangian

$$\mathcal{L}_{\text{MDM}} = \frac{e}{4m_\mu} a_\mu \bar{l}_\mu \sigma^{\alpha\beta} l_\mu F_{\alpha\beta}, \quad (\text{B1})$$

where l_μ denotes the muon which is on-shell, m_μ is the mass of the muon, $\sigma^{\alpha\beta} = \frac{i}{2} [\gamma^\alpha, \gamma^\beta]$, $F_{\alpha\beta}$ represents the electromagnetic field strength and muon MDM $a_\mu = \frac{1}{2}(g-2)_\mu$. Adopting the effective Lagrangian approach, the MDM of the muon can be written by [152–154]

$$a_\mu = 4m_\mu^2 \Re(C_2^R + C_2^{L*} + C_6^R), \quad (\text{B2})$$

where $\Re(\dots)$ denotes the operation to take the real part of the complex number, and $C_{2,6}^{L,R}$ represent the Wilson coefficients of the corresponding effective operators $O_{2,6}^{L,R}$

$$\begin{aligned} O_2^{L,R} &= \frac{eQ_f}{(4\pi)^2} \overline{(iD_\alpha l_\mu)} \gamma^\alpha F \cdot \sigma P_{L,R} l_\mu, \\ O_6^{L,R} &= \frac{eQ_f m_\mu}{(4\pi)^2} \bar{l}_\mu F \cdot \sigma P_{L,R} l_\mu. \end{aligned} \quad (\text{B3})$$

The SUSY corrections of the Wilson coefficients in the

fermions, and neutral fermions can be similarly written by

$$\begin{aligned} \mathcal{L}_{\text{int}} &= S_\alpha^- \bar{\chi}_\beta (C_L^{S_\alpha^- \chi_\eta^0 \bar{\chi}_\beta} P_L + C_R^{S_\alpha^- \chi_\eta^0 \bar{\chi}_\beta} P_R) \chi_\eta^0 \\ &\quad + S_\alpha^{-*} \bar{\chi}_\eta^0 (C_L^{S_\alpha^{-*} \chi_\beta \bar{\chi}_\eta^0} P_L + C_R^{S_\alpha^{-*} \chi_\beta \bar{\chi}_\eta^0} P_R) \chi_\beta, \end{aligned} \quad (\text{A12})$$

where

$$\begin{aligned} C_L^{S_\alpha^- \chi_\eta^0 \bar{\chi}_\beta} &= \frac{-e}{\sqrt{2}s_W c_W} R_{S^\pm}^{2\alpha*} Z_+^{2\beta} \left[c_W Z_n^{2\eta} + s_W Z_n^{1\eta} \right] \\ &\quad - \frac{e}{s_W} R_{S^\pm}^{2\alpha*} Z_+^{1\beta} Z_n^{4\eta} - \frac{\sqrt{2}e}{c_W} R_{S^\pm}^{(5+i)\alpha*} Z_+^{(2+i)\beta} Z_n^{1\eta} \\ &\quad + Y_{\nu_{ij}} R_{S^\pm}^{(2+i)\alpha} Z_+^{2\beta} Z_n^{(4+j)\eta} + Y_{e_{ij}} Z_+^{(2+j)\beta} \\ &\quad \cdot \left[R_{S^\pm}^{1\alpha} Z_n^{(7+i)\eta} - R_{S^\pm}^{(2+i)\alpha} Z_n^{3\eta} \right] \\ &\quad - \lambda_i R_{S^\pm}^{1\alpha} Z_+^{2\beta} Z_n^{(4+i)\eta}, \end{aligned} \quad (\text{A13})$$

$$\begin{aligned} C_L^{S_\alpha^{-*} \chi_\beta \bar{\chi}_\eta^0} &= \frac{e}{\sqrt{2}s_W c_W} \left[R_{S^\pm}^{1\alpha*} Z_-^{2\beta} + R_{S^\pm}^{(2+i)\alpha*} Z_-^{(2+i)\beta} \right] \\ &\quad \cdot \left[c_W Z_n^{2\eta} + s_W Z_n^{1\eta} \right] \\ &\quad - \frac{e}{s_W} Z_-^{1\beta} \left[R_{S^\pm}^{1\alpha*} Z_n^{3\eta} + R_{S^\pm}^{(2+i)\alpha*} Z_n^{(7+i)\eta} \right] \\ &\quad + Y_{\nu_{ij}} R_{S^\pm}^{2\alpha} Z_-^{(2+i)\beta} Z_n^{(4+j)\eta} + Y_{e_{ij}} R_{S^\pm}^{(5+j)\alpha} \\ &\quad \cdot \left[Z_-^{2\beta} Z_n^{(7+i)\eta} - Z_-^{(2+i)\beta} Z_n^{3\eta} \right] \\ &\quad - \lambda_i R_{S^\pm}^{2\alpha} Z_-^{2\beta} Z_n^{(4+i)\eta}, \end{aligned} \quad (\text{A14})$$

$$C_R^{S_\alpha^- \chi_\eta^0 \bar{\chi}_\beta} = [C_L^{S_\alpha^- \chi_\beta \bar{\chi}_\eta^0}]^*, \quad C_R^{S_\alpha^{-*} \chi_\beta \bar{\chi}_\eta^0} = [C_L^{S_\alpha^{-*} \chi_\eta^0 \bar{\chi}_\beta}]^*. \quad (\text{A15})$$

$\mu\nu$ SUSY can be

$$C_{2,6}^{L,R} = C_{2,6}^{L,R(n)} + C_{2,6}^{L,R(c)}. \quad (\text{B4})$$

The effective coefficients $C_{2,6}^{L,R(n)}$ denote the contributions from the neutralinos χ_η^0 and the charged scalars S_α^- loops

$$\begin{aligned} C_2^{R(n)} &= \frac{1}{m_W^2} C_L^{S_\alpha^- \chi_\eta^0 \bar{l}_\mu} C_R^{S_\alpha^{-*} l_\mu \bar{\chi}_\eta^0} \left[-I_3(x_{\chi_\eta^0}, x_{S_\alpha^-}) \right. \\ &\quad \left. + I_4(x_{\chi_\eta^0}, x_{S_\alpha^-}) \right], \\ C_6^{R(n)} &= \frac{m_{\chi_\eta^0}}{m_W^2 m_\mu} C_R^{S_\alpha^- \chi_\eta^0 \bar{l}_\mu} C_R^{S_\alpha^{-*} l_\mu \bar{\chi}_\eta^0} \left[-2I_1(x_{\chi_\eta^0}, x_{S_\alpha^-}) \right. \\ &\quad \left. + 2I_3(x_{\chi_\eta^0}, x_{S_\alpha^-}) \right], \\ C_{2,6}^{L(n)} &= C_{2,6}^{R(n)} |_{L \leftrightarrow R}. \end{aligned} \quad (\text{B5})$$

Similarly, the contributions $C_{2,6}^{L,R(c)}$ coming from the charginos χ_β and the neutral scalars N_α ($N = S, P$) loops are

$$C_2^{R(c)} = \sum_{N=S,P} \frac{1}{m_W^2} C_R^{N_\alpha \chi_\beta \bar{l}_\mu} C_L^{N_\alpha l_\mu \bar{\chi}_\beta} \left[-I_1(x_{\chi_\beta}, x_{N_\alpha}) \right]$$

$$\begin{aligned}
& +2I_3(x_{\chi_\beta}, x_{N_\alpha}) - I_4(x_{\chi_\beta}, x_{N_\alpha}) \Big], \\
C_6^{R(c)} &= \sum_{N=S,P} \frac{m_{\chi_\beta}}{m_W^2 m_\mu} C_R^{N_\alpha \chi_\beta \bar{l}_\mu} C_R^{N_\alpha l_\mu \bar{\chi}_\beta} \Big[2I_1(x_{\chi_\beta}, x_{N_\alpha}) \\
& - 2I_2(x_{\chi_\beta}, x_{N_\alpha}) - 2I_3(x_{\chi_\beta}, x_{N_\alpha}) \Big], \\
C_{2,6}^{L(c)} &= C_{2,6}^{R(c)} |_{L \leftrightarrow R}. \tag{B6}
\end{aligned}$$

Here, the loop functions $I_i(x_1, x_2)$ are given as

$$I_1(x_1, x_2) = \frac{1}{16\pi^2} \left[\frac{1 + \ln x_2}{x_1 - x_2} - \frac{x_1 \ln x_1 - x_2 \ln x_2}{(x_1 - x_2)^2} \right], \tag{B7}$$

$$\begin{aligned}
I_2(x_1, x_2) &= \frac{1}{16\pi^2} \left[-\frac{1 + \ln x_1}{x_1 - x_2} + \frac{x_1 \ln x_1 - x_2 \ln x_2}{(x_1 - x_2)^2} \right], \\
I_3(x_1, x_2) &= \frac{1}{32\pi^2} \left[\frac{3 + 2 \ln x_2}{x_1 - x_2} + \frac{2x_2 + 4x_2 \ln x_2}{(x_1 - x_2)^2} \right. \\
&\quad \left. - \frac{2x_1^2 \ln x_1}{(x_1 - x_2)^3} + \frac{2x_2^2 \ln x_2}{(x_1 - x_2)^3} \right], \tag{B8}
\end{aligned}$$

$$\begin{aligned}
I_4(x_1, x_2) &= \frac{1}{96\pi^2} \left[\frac{11 + 6 \ln x_2}{x_1 - x_2} + \frac{15x_2 + 18x_2 \ln x_2}{(x_1 - x_2)^2} \right. \\
&\quad \left. + \frac{6x_2^2 + 18x_2^2 \ln x_2}{(x_1 - x_2)^3} - \frac{6x_1^3 \ln x_1 - 6x_2^3 \ln x_2}{(x_1 - x_2)^4} \right]. \tag{B9}
\end{aligned}$$

References

- 1 G. Aad et al. (ATLAS Collaboration), Phys. Lett. B, **716**: 1 (2012), arXiv:1207.7214 [hep-ex]
- 2 S. Chatrchyan et al. (CMS Collaboration), Phys. Lett. B, **716**: 30 (2012), arXiv:1207.7235 [hep-ex]
- 3 G. Aad et al. (ATLAS and CMS Collaborations), Phys. Rev. Lett., **114**: 191803 (2015), arXiv:1503.07589 [hep-ex]
- 4 R. Harnik, J. Kopp and J. Zupan, JHEP, **03**: 026 (2013), arXiv:1209.1397
- 5 V. Khachatryan et al. (CMS Collaboration), Phys. Lett. B, **749**: 337–362 (2015), arXiv:1502.07400 [hep-ex]
- 6 V. Khachatryan et al. (CMS Collaboration), CMS-PAS-HIG-16-005
- 7 G. Aad et al. (ATLAS Collaboration), JHEP, **11**: 211 (2015), arXiv:1508.03372 [hep-ex]
- 8 G. Aad et al. (ATLAS Collaboration), arXiv:1604.07730 [hep-ex]
- 9 J. L. Diaz-Cruz and J. J. Toscano, Phys. Rev. D, **62**: 116005 (2000), arXiv:hep-ph/9910233
- 10 T. Han and D. Marfatia, Phys. Rev. Lett., **86**: 1442 (2001), arXiv: hep-ph/0008141
- 11 J.L. Diaz-Cruz, JHEP, **05**: 036 (2003)
- 12 A. Brignole and A. Rossi, Phys. Lett. B, **566**: 217–225 (2003)
- 13 A. Brignole and A. Rossi, Nucl. Phys. B, **701**: 3–53 (2004)
- 14 E. Arganda, A.M. Curiel, M.J. Herrero and D. Temes, Phys. Rev. D, **71**: 035011 (2005)
- 15 J. K. Parry, Nucl. Phys. B, **760**: 38–63 (2007)
- 16 J. L. Diaz-Cruz, D. K. Ghosh and S. Moretti, Phys. Lett. B, **679**: 376–381 (2009)
- 17 A. Arhrib, Y. Cheng and O. C. W. Kong, Phys. Rev. D, **87**: 015025 (2013), arXiv:1210.8241
- 18 K. Agashe and R. Contino, Phys. Rev. D, **80**: 075016 (2009), arXiv:0906.1542
- 19 A. Azatov, M. Toharia and L. Zhu, Phys. Rev. D, **80**: 035016 (2009), arXiv:0906.1990
- 20 S. Casagrande et al., JHEP, **10**: 094 (2008), arXiv:0807.4937
- 21 G. Perez and L. Randall, JHEP, **01**: 077 (2009), arXiv:0805.4652
- 22 A. J. Buras, B. Duling and S. Gori, JHEP, **09**: 076 (2009), arXiv:0905.2318
- 23 M. Blanke, A. J. Buras, B. Duling et al, JHEP, **03**: 001 (2009), arXiv:0809.1073
- 24 G. F. Giudice and O. Lebedev, Phys. Lett. B, **665**: 79 (2008), arXiv:0804.1753
- 25 J. A. Aguilar-Saavedra, Nucl. Phys. B, **821**: 215 (2009), arXiv:0904.2387
- 26 M. E. Albrecht, M. Blanke, A. J. Buras et al, JHEP, **09**: 064 (2009), arXiv:0903.2415
- 27 H. Ishimori, T. Kobayashi, H. Ohki et al, Prog. Theor. Phys. Suppl., **183**: 1 (2010), arXiv:1003.3552
- 28 A. Goudelis, O. Lebedev and J. H. Park, Phys. Lett. B, **707**: 369 (2012), arXiv:1111.1715
- 29 D. McKeen, M. Pospelov and A. Ritz, Phys. Rev. D, **86**: 113004 (2012), arXiv:1208.4597
- 30 I. de Medeiros Varzielas, O. Fischer and V. Maurer, JHEP, **08**: 080 (2015), arXiv:1504.03955
- 31 A. Pilaftsis, Phys. Lett. B, **285**: 68–74 (1992)
- 32 K. A. Assamagan, A. Deandrea and P.-A. Delsart, Phys. Rev. D, **67**: 035001 (2003), arXiv:hep-ph/0207302
- 33 S. Kanemura, K. Matsuda, T. Ota et al, Phys. Lett. B, **599**: 83–91 (2004), arXiv:hep-ph/0406316
- 34 U. Cotti, M. Pineda and G. Tavares-Velasco, arXiv:hep-ph/0501162
- 35 S. Kanemura, T. Ota and K. Tsumura, Phys. Rev. D, **73**: 016006 (2006), arXiv:hep-ph/0505191
- 36 M. Cannoni and O. Panella, Phys. Rev. D, **79**: 056001 (2009), arXiv:0812.2875
- 37 S. Kanemura and K. Tsumura, Phys. Lett. B, **674**: 295–298 (2009), arXiv:0901.3159
- 38 S.-L. Chen, M. Frigerio and E. Ma, Phys. Lett. B, **612**: 29–35 (2005), arXiv:hep-ph/0412018
- 39 M. Cannoni and O. Panella, Phys. Rev. D, **79**: 056001 (2009), arXiv:0812.2875
- 40 E. Iltan, Mod. Phys. Lett. A, **24**: 1361 (2009), arXiv:0809.3594
- 41 G. Blankenburg, J. Ellis and G. Isidori, Phys. Lett. B, **712**: 386–390 (2012), arXiv:1202.5704
- 42 A. Arhrib, Y. Cheng and O. C. W. Kong, Europhys. Lett., **101**: 31003 (2013), arXiv:1208.4669
- 43 A. Dery, A. Efrati, Y. Hochberg and Y. Nir, JHEP, **05**: 039 (2013), arXiv:1302.3229
- 44 M. Arana-Catania, E. Arganda and M. Herrero, JHEP, **09**: 160 (2013) [Erratum-ibid. 10 (2015) 192], arXiv:1304.3371
- 45 M. Arroyo, J. L. Diaz-Cruz, E. Diaz and J. A. Ordzu-Ducuara, arXiv:1306.2343
- 46 A. Celis, V. Cirigliano and E. Passemard, Phys. Rev. D, **89**: 013008 (2014), arXiv:1309.3564
- 47 A. Falkowski, D. M. Straub and A. Vicente, JHEP, **05**: 092 (2014), arXiv:1312.5329
- 48 A. Dery, A. Efrati, Y. Nir et al, Phys. Rev. D, **90**: 115022 (2014), arXiv:1408.1371
- 49 M. D. Campos, A. E. Carcamo Hernández, H. Päs and E. Schumacher, Phys. Rev. D, **91**: 116011 (2015), arXiv:1408.1652
- 50 D. Aristizabal Sierra and A. Vicente, Phys. Rev. D, **90**: 115004 (2014), arXiv:1409.7690
- 51 J. Heeck, M. Holthausen, W. Rodejohann and Y. Shimizu, Nucl. Phys. B, **896**: 281–310 (2015), arXiv:1412.3671
- 52 A. Crivellin, G. D'Ambrosio and J. Heeck, Phys. Rev. Lett., **114**: 151801 (2015), arXiv:1501.00993
- 53 I. Dorner, S. Fajfer, A. Greljo, et al, JHEP, **06**: 108 (2015), arXiv:1502.07784
- 54 Y. Omura, E. Senaha and K. Tobe, JHEP, **05**: 028 (2015),

- arXiv:1502.07824
- 55 A. Crivellin, G. D'Ambrosio and J. Heeck, Phys. Rev. D, **91**: 075006 (2015), arXiv:1503.03477
- 56 A. Vicente, Adv. High Energy Phys. **2015**: 686572 (2015), arXiv:1503.08622
- 57 F. Bishara, J. Brod, P. Uttayarat and J. Zupan, JHEP, **01**: 010 (2016), arXiv:1504.04022
- 58 X.-G. He, J. Tandean and Y.-J. Zheng, JHEP, **09**: 093 (2015), arXiv:1507.02673
- 59 W. Altmannshofer, S. Gori, A. L. Kagan et al, Phys. Rev. D, **93**: 031301 (2016), arXiv:1507.07927
- 60 K. Cheung, W.-Y. Keung and P.-Y. Tseng, Phys. Rev. D, **93**: 015010 (2016), arXiv:1508.01897
- 61 E. Arganda, M. J. Herrero, X. Marcano and C. Weiland, Phys. Rev. D, **93**: 055010 (2016), arXiv:1508.04623
- 62 F. J. Botella, G. C. Branco, M. Nebot and M. N. Rebelo, Eur. Phys. J. C, **76**: 161 (2016), arXiv:1508.05101
- 63 X. Liu, L. Bian, X.-Q. Li and J. Shu, Nucl. Phys. B, **909**: 507-524 (2016), arXiv:1508.05716
- 64 S. Baek and K. Nishiwaki, Phys. Rev. D, **93**: 015002 (2016), arXiv:1509.07410
- 65 W. Huang and Y.-L. Tang, Phys. Rev. D, **92**: 094015 (2015), arXiv:1509.08599
- 66 S. Baek and Z. Kang, JHEP, **03**: 106 (2016), arXiv:1510.00100
- 67 E. Arganda, M. J. Herrero, R. Morales and A. Szynkman, JHEP, **03**: 055 (2016), arXiv:1510.04685
- 68 D. Aloni, Y. Nir and E. Stamouz, JHEP, **1604**: 162 (2016), arXiv:1511.00979
- 69 P. T. Giang, L.T. Hue, D.T. Huong and H. N. Long, Nucl. Phys. B, **864**: 85-112 (2012), arXiv:1204.2902
- 70 D. T. Binh, L. T. Hue, D. T. Huong and H. N. Long, Eur. Phys. J. C, **74**: 2851 (2014), arXiv:1308.3085
- 71 R. Harnik, J. Kopp and J. Zupan, JHEP, **03**: 026 (2013), arXiv:1209.1397
- 72 S. Davidson and P. Verdier, Phys. Rev. D, **86**: 111701 (2012), arXiv:1211.1248
- 73 E. Arganda, M. J. Herrero, X. Marcano and C. Weiland, Phys. Rev. D, **91**: 015001 (2015), arXiv:1405.4300
- 74 S. Bressler, A. Dery and A. Efrati, Phys. Rev. D, **90**: 015025 (2014), arXiv:1405.4545
- 75 J. Kopp and M. Nardecchia, JHEP, **10**: 156 (2014), arXiv:1406.5303
- 76 M. A. López-Osorio, E. Martínez-Pascual and J. J. Toscano, J. Phys. G: Nucl. Part. Phys., **43**: 025003 (2016), arXiv:1408.3307
- 77 C.-J. Lee and J. Tandean, JHEP, **04**: 174 (2015), arXiv:1410.6803
- 78 L. de Lima, C. S. Machado, R. D. Matheus and L. A. F. do Prado, JHEP, **11**: 074 (2015), arXiv:1501.06923
- 79 I. M. Varzielas and G. Hiller, arXiv:1503.01084
- 80 D. Das and A. Kundu, Phys. Rev. D, **92**: 015009 (2015), arXiv:1504.01125
- 81 B. Bhattacherjee, S. Chakraborty and S. Mukherjee, Mod. Phys. Lett. A, **31**: 1650174 (2016), arXiv:1505.02688
- 82 Y. Mao and S. Zhu, Phys. Rev. D, **93**: 035014 (2016), arXiv:1505.07668
- 83 R. Benbrik, C.-H. Chen and T. Nomura, Phys. Rev. D, **93**: 095004 (2016), arXiv:1511.08544
- 84 Y. Omura, E. Senaha and K. Tobe, Phys. Rev. D, **94**: 055019 (2016), arXiv:1511.08880
- 85 M. Sher and K. Thrasher, Phys. Rev. D, **93**: 055021 (2016), arXiv:1601.03973
- 86 M. Buschmann, J. Kopp, J. Liu and X.-P. Wang, JHEP, **1606**: 149 (2016), arXiv:1601.02616
- 87 Y. Farzan and I. M. Shoemaker, JHEP, **07**: 033 (2016), arXiv:1512.09147
- 88 N. Bizot, S. Davidson, M. Frigerio, and J.-L. Kneur, JHEP, **03**: 073 (2016), arXiv:1512.08508
- 89 C.-F. Chang, C.-H. V. Chang, C. S. Nugroho, and T.-C. Yuan, Nucl. Phys. B, **910**: 293-308 (2016), arXiv:1602.00680
- 90 C.-H. Chen and T. Nomura, Eur. Phys. J. C, **76**: 353 (2016), arXiv:1602.07519
- 91 C. Alvarado, R.M. Capdevilla, A. Delgado and A. Martin, arXiv:1602.08506
- 92 A. Hayreter, X.-G. He and G. Valencia, Phys. Lett. B, **760**: 175-177 (2016), arXiv:1603.06326
- 93 K. Huitu, V. Keus, N. Koivunen and O. Lebedev, JHEP, **1605**: 026 (2016), arXiv:1603.06614
- 94 T. T. Thuc, L. T. Hue, H. N. Long and T. P. Nguyen, Phys. Rev. D, **93**: 115026 (2016), arXiv:1604.03285
- 95 S. Baek, T. Nomura and H. Okada, Phys. Lett. B, **759**: 91-98 (2016), arXiv:1604.03738
- 96 J. Herrero-Garcia, N. Rius and A. Santamaria, arXiv:1605.06091
- 97 K. H. Phan, H. T. Hung and L. T. Hue, arXiv:1605.07164
- 98 S. V. Demidov and I.V. Sobolev, JHEP, **1608**: 030 (2016), arXiv:1605.08220
- 99 B. Yang, J. Han and N. Liu, arXiv:1605.09248
- 100 A. Hayreter, X.-G. He and G. Valencia, Phys. Rev. D, **94**: 075002 (2016), arXiv:1606.00951
- 101 L. Wang, S. Yang and X.-F. Han, arXiv:1606.04408
- 102 A. Efrati, J. F. Kamenik and Y. Nir, arXiv:1606.07082
- 103 A. D. Iura, J. Herrero-Garcia and D. Meloni, Nucl. Phys. B, **911**: 388-424 (2016), arXiv:1606.08785
- 104 M. Aoki, S. Kanemura, K. Sakurai and H. Sugiyama, arXiv:1607.08548
- 105 S. Fathy, T. Ibrahim, A. Itani and P. Nath, arXiv:1608.05998
- 106 D. E. López-Fogliani and C. Muñoz, Phys. Rev. Lett., **97**: 041801 (2006), hep-ph/0508297
- 107 N. Escudero, D.E. López-Fogliani, C. Muñoz and R. Ruiz de Austri, JHEP, **12**: 099 (2008) arXiv:0810.1507
- 108 J. Fidalgo, D. E. López-Fogliani, C. Muñoz and R. Ruiz de Austri, JHEP, **10**: 020 (2011), arXiv:1107.4614
- 109 H. P. Nilles, Phys. Rept., **110**: 1 (1984)
- 110 H. E. Haber and G.L. Kane, Phys. Rept., **117**: 75 (1985)
- 111 H. E. Haber, hep-ph/9306207
- 112 S. P. Martin, hep-ph/9709356
- 113 J. Rosiek, Phys. Rev. D, **41**: 3464 (1990), hep-ph/9511250
- 114 J. E. Kim and H. P. Nilles, Phys. Lett. B, **138**: 150 (1984)
- 115 P. Ghosh and S. Roy, JHEP, **04**: 069 (2009), arXiv:0812.0084
- 116 A. Bartl, M. Hirsch, S. Liebler et al. JHEP, **05**: 120 (2009), arXiv:0903.3596
- 117 J. Fidalgo, D. E. López-Fogliani, C. Muñoz and R.R. de Austri, JHEP, **08**: 105 (2009), arXiv:0904.3112
- 118 P. Ghosh, P. Dey, B. Mukhopadhyaya and S. Roy, JHEP, **05**: 087 (2010), arXiv:1002.2705
- 119 H.-B. Zhang, T.-F. Feng, L.-N. Kou and S.-M. Zhao, Int. J. Mod. Phys. A, **28**: 1350117 (2013), arXiv:1307.6284
- 120 H.-B. Zhang, T.-F. Feng, S.-M. Zhao and T.-J. Gao, Nucl. Phys. B, **873**: 300 (2013), arXiv:1304.6248
- 121 H.-B. Zhang, T.-F. Feng, G.-F. Luo et al, JHEP, **07**: 069 (2013) [Erratum-ibid. **10**: 173 (2013)], arXiv:1305.4352
- 122 H.-B. Zhang, T.-F. Feng, S.-M. Zhao and F. Sun, Int. J. Mod. Phys. A, **29**: 1450123 (2014), arXiv:1407.7365
- 123 H.-B. Zhang, T.-F. Feng, F. Sun et al, Phys. Rev. D, **89**: 115007 (2014), arXiv:1307.3607
- 124 G. 't Hooft and M. Veltman, Nucl. Phys. B, **153**: 365 (1979)
- 125 R. Mertig, M. Bohm and A. Denner, Comput. Phys. Commun., **64**: 345 (1991)
- 126 A. Denner, Fortsch. Phys., **41**: 307 (1993)
- 127 A. Denner and S. Dittmaier, Nucl. Phys. B, **658**: 175-202 (2003)
- 128 T. Hahn and M. Perez-Victoria, Comput. Phys. Commun., **118**: 153 (1999)
- 129 T. Hahn, Comput. Phys. Commun., **140**: 418 (2001)
- 130 T. Hahn and C. Schappacher, Comput. Phys. Commun., **143**: 54 (2002)

- 131 S. Heinemeyer et al. (LHC Higgs Cross Section Working Group), CERN-2013-004, arXiv:1307.1347 [hep-ph]
- 132 M. Misiak, S. Pokorski and J. Rosiek, Adv. Ser. Direct. High Energy Phys., **15**: 795 (1998), hep-ph/9703442
- 133 P. Paradisi, JHEP, **10** : 006 (2005), hep-ph/0505046
- 134 J. Girrbach, S. Mertens, U. Nierste and S. Wiesenfeldt, JHEP, **05**: 026 (2010), arXiv:0910.2663
- 135 J. Rosiek, P. H. Chankowski, A. Dedes et al, Comput. Phys. Commun., **181**: 2180 (2010), arXiv:1003.4260
- 136 M. Arana-Catania, S. Heinemeyer and M. J. Herrero, Phys. Rev. D, **88**: 015026 (2013), arXiv:1304.2783
- 137 H.-B. Zhang, T.-F. Feng, Z.-F. Ge and S.-M. Zhao, JHEP, **02**: 012 (2014), arXiv:1401.2704
- 138 H.-B. Zhang, G.-L. Luo, T.-F. Feng et al, Mod. Phys. Lett. A, **29**: 1450196 (2014), arXiv:1409.6837
- 139 M. Carena, J. R. Espinosa, M. Quirós and C. E. M. Wagner, Phys. Lett. B, **355**: 209 (1995)
- 140 M. Carena, M. Quirós and C. E. M. Wagner, Nucl. Phys. B, **461**: 407 (1996)
- 141 M. Carena, S. Gori, N.R. Shah and C. E. M. Wagner, JHEP, **03**: 014 (2012)
- 142 K. A. Olive et al. (Particle Data Group), Chin. Phys. C, **38**: 090001 (2014)
- 143 G. W. Bennett et al. (Muon (g-2) Collaboration), Phys. Rev. D, **73**: 072003 (2006)
- 144 P. J. Mohr, B. N. Taylor and D.B. Newell, Rev. Mod. Phys. **80**: 633 (2008)
- 145 ATLAS Collaboration, Phys. Rev. D, **86**: 092002 (2012)
- 146 ATLAS Collaboration, JHEP, **10**: 130 (2013)
- 147 CMS Collaboration, JHEP, **01**: 077 (2013)
- 148 CMS Collaboration, JHEP, **07**: 122 (2013)
- 149 B. Aubert et al. (BaBar Collaboration), Phys. Rev. Lett., **104**: 021802 (2010), arXiv:0908.2381
- 150 T. Appelquist and J. Carazzone, Phys. Rev. D, **11**: 2856 (1975)
- 151 P. Draper and H. E. Haber, Eur. Phys. J. C, **73**: 2522 (2013)
- 152 T. F. Feng, L. Sun and X. Y. Yang, Nucl. Phys. B, **800**: 221 (2008)
- 153 T. F. Feng, L. Sun and X. Y. Yang, Phys. Rev. D, **77**: 116008 (2008)
- 154 T. F. Feng and X. Y. Yang, Nucl. Phys. B, **814**: 101 (2009)